

# The emergence of (3+1)-dimensional space-time in the type IIB matrix model

Kohta Hatakeyama (KEK)

In collaboration with Konstantinos N. Anagnostopoulos (Natl. Tech. Univ. of Athens), Takehiro Azuma (Setsunan Univ.), Mitsuaki Hirasawa (INFN), Jun Nishimura (KEK, SOKENDAI), Stratos Kovalkov Papadoudis (Natl. Tech. Univ. of Athens), and Asato Tsuchiya (Shizuoka Univ.)

Anagnostopoulos-Azuma-KH-Hirasawa-Nishimura-Papadoudis-Tsuchiya, work in progress

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# 1. Introduction

To obtain the 4D space-time from superstring theory, which is a promising candidate of the unified theory including quantum gravity, non-perturbative effects of superstring theory are considered to be important.

The type IIB matrix model [Ishibashi-Kawai-Kitazawa-Tsuchiya ('96)]

A non-perturbative formulation of superstring theory.

In this model, space-time does not exist a priori but emerges

dynamically from the degrees of freedom of matrices.

$$S = -N \operatorname{Tr} \left( \frac{1}{4} [A^{\mu}, A^{\nu}] [A_{\mu}, A_{\nu}] + \frac{1}{2} \overline{\Psi} \Gamma^{\mu} [A_{\mu}, \Psi] \right)$$

$$= S_{\mathrm{b}} + S_{\mathrm{f}} , S_{\mathrm{b}} = -\frac{N}{4} \operatorname{Tr} \left( [A^{\mu}, A^{\nu}] [A_{\mu}, A_{\nu}] \right), S_{\mathrm{f}} = -\frac{N}{2} \operatorname{Tr} \left( \overline{\Psi} \Gamma^{\mu} [A_{\mu}, \Psi] \right)$$

$$\Gamma^{\mu} : 10D \text{ Gamma matrices}$$

$$(\mu = 0, \dots, 9)$$

$$N \times N \text{ Hermitian matrices}$$

$$\frac{A_{\mu} : 10D \text{ Lorentz vector}}{\Psi : 10D \text{ Majorana-Weyl spinor}} \text{ under SO(9,1) transformation}$$

space-time is extracted from the eigenvalue distribution of  $A_{\mu}$ .

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#### Does real space-time emerge?

◆ The type IIB matrix model [Ishibashi-Kawai-Kitazawa-Tsuchiya ('96)] Partition function:  $Z = \int dAd\Psi e^{i(S_{\rm b}+S_{\rm f})} = \int dAe^{iS_{\rm b}} \operatorname{Pf}\mathcal{M}(A)$ complex → sign problem! Definition of expectation values:  $\langle \mathcal{O} \rangle = \frac{1}{Z} \int dA \mathcal{O} e^{iS_{\rm b}} \operatorname{Pf}\mathcal{M}(A)$ 

Even if  $A_{\mu}$  are Hermitian, the expectation values of eigenvalues are complex in general.

In the original model, complex phase of  $\langle A_0 \rangle \sim e^{-3\pi i/8}$  and that of  $\langle A_i \rangle \sim e^{\pi i/8}$ .

 $\rightarrow$  Real space-time cannot be realized!

How does real space-time emerge in this model?

#### **Classical solutions**

Classical EOM for the type IIB matrix model:  $\frac{\delta S}{\delta A_{\mu}} = [A^{\nu}, [A_{\nu}, A_{\mu}]] = 0$ All simultaneously diagonalizable  $A_{\mu}$  are classical solutions, so there is no expansion in general for such solutions. In the previous work [KH-Matsumoto-Nishimura-Tsuchiya-Yosprakob ('19)], to implement the effects of the IR regularization required for the Lorentzian model, we added the quadratic term of  $A_{\mu}$  (mass term)  $-\frac{1}{2}N\gamma \left[ \text{Tr}(A_0)^2 - \text{Tr}(A_i)^2 \right]$ .

Classical EOM with mass term:  $\frac{\delta S}{\delta A_{\mu}} = [A^{\nu}, [A_{\nu}, A_{\mu}]] - \gamma A_{\mu} = 0$ 

•  $\gamma > 0$ : classical solutions with an expanding behavior



← Classical solution for only 3d space expands. ※Dimensionality of expanding space is arbitrary. But not for  $\gamma < 0$ !

We introduce the same mass term when we perform simulations of this model.

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## 2. Complex Langevin method (CLM)

We perform numerical simulations

based on the complex Langevin method (CLM) to overcome the sign problem.

 $Z = \int dx w(x), \quad x \in \mathbb{R} \longrightarrow z \in \mathbb{C}$ complexify variable [Parisi ('83), Klauder ('84)]  $\bullet$  Complex Langevin equation ( $t_{\rm L}$ : Langevin time)  $\frac{dz_{k}}{dt_{L}} = \frac{1}{w} \frac{\partial w}{\partial z_{k}} + \frac{\eta_{k}(t_{L})}{Gaussian \text{ noise, real}} P(\eta_{k}(t_{L})) \propto \exp\left(-\frac{1}{4} \int dt_{L} \sum_{k} [\eta_{k}(t_{L})]^{2}\right)$ Gaussian noise, real log (probability) Necessary and sufficient condition to justify the CLM

[Nagata-Nishimura-Shimasaki ('16)]

The probability distribution of the drift term should be exponentially suppressed.



## Application of CLM to the type IIB matrix model

igstarrow Order of eigenvalues of  $A_0$   $(lpha_i)$  [Nishimura-Tsuchiya ('19)]

 $A_0 = \operatorname{diag}(\alpha_1, \ldots, \alpha_N)$ Change of variables N-1 $\alpha_1 < \alpha_2 < \cdots < \alpha_N \Longrightarrow \alpha_1 = 0, \ \alpha_2 = e^{\tau_1}, \ \alpha_3 = e^{\tau_1} + e^{\tau_2}, \ \ldots, \ \alpha_N = \sum e^{\tau_k}$ shift symmetry Complexify variables  $au_k \in \mathbb{R}$  $au_k \in \mathbb{C}$ Hermitian matrices:  $A_i \in SU(N)$  General complex matrices:  $A_i \in GL(N,\mathbb{C})$ • Complex Langevin equations  $t_{\rm L}$ : Langevin time  $\frac{d\tau_{k}}{dt_{\rm L}} = -\frac{\partial S_{\rm eff}}{\partial \tau_{k}} + \eta_{k}(t_{\rm L}) \qquad P(\eta_{k}(t_{\rm L})) \propto \exp\left(-\frac{1}{4}\int dt_{\rm L}\sum_{k}[\eta_{k}(t_{\rm L})]^{2}\right) \\ \frac{d(A_{i})_{kl}}{dt_{\rm L}} = -\frac{\partial S_{\rm eff}}{\partial (A_{i})_{lk}} + (\eta_{i})_{kl}(t_{\rm L}) \qquad P(\eta_{i}(t_{\rm L})) \propto \exp\left(-\frac{1}{4}\int dt_{\rm L}\sum_{i}{\rm Tr}[\eta_{i}(t_{\rm L})]^{2}\right) \right) \qquad \text{Gaussian noise}$  $S_{\text{eff}} = -\frac{i}{4}N\left(2\operatorname{Tr}\left[A_{0},A_{i}\right]^{2} - \operatorname{Tr}\left[A_{i},A_{j}\right]^{2}\right) - \frac{i}{2}N\gamma\left(\operatorname{Tr}(A_{0})^{2} - \operatorname{Tr}(A_{i})^{2}\right) - 2\log\Delta(\alpha) - \sum_{k=1}^{N-1}\tau_{k} - \log\operatorname{Pf}\mathcal{M}(A) , \ \Delta(\alpha) = \prod_{k=1}^{N}(\alpha_{k} - \alpha_{l})$ 

## Some tricks to make the CLM work

#### Singular drift problem

$$Z=\int dA e^{iS_{
m b}} {
m Pf} {\cal M}(A)$$

 $\mathcal{M}$  that appears in  $\mathbf{Pf}\mathcal{M}$  has eigenvalues accumulating near zero, which causes the singular drift problem. Due to this problem, the CLM fails. We introduce a deformation term in fermionic action to avoid the problem.

$$S_{
m f} = -rac{N}{2} \operatorname{Tr} \left( ar{\Psi}_{lpha} (\Gamma^{\mu})_{lphaeta} [A_{\mu}, \Psi_{eta}] + im_{
m f} ar{\Psi}_{lpha} (\Gamma_{7} \Gamma_{8}^{\dagger} \Gamma_{9})_{lphaeta} \Psi_{eta} 
ight)$$

 $m_{
m f}$ : deformation parameter  $\left\{egin{array}{c} m_{
m f} o \infty : {
m bosonic} \ ({
m Due \ to \ decoupling \ of \ fermionic \ degrees \ of \ freedom)} \ m_{
m f} o 0 \ : {
m SUSY} \end{array}
ight.$ 

#### Stabilization

Cf.) [Attanasio-Jäger ('18)] : dynamical stabilization in CL simulation of QCD

To stabilize the CLM,  $A_i 
ightarrow rac{1}{1+\eta} (A_i + \eta A_i^\dagger)$  after each Langevin step

 $\int_{\eta=1}^{\eta=0: \text{ do nothing}} Here, \eta = 0.01.$   $\eta = 1: \text{ Hermitianize} \text{ %Justifiable when dominant configurations are close to Hermitian.}$ 

## 3. Results

#### How to determine whether time and space are real

• Distribution of  $lpha_i$  (eigenvalues of  $A_0$ )  $\Delta \langle lpha_i \rangle \equiv \langle lpha_{i+1} 
angle - \langle lpha_i 
angle \propto e^{i heta_{ ext{t}}}$ 

Definition of time: 
$$\bar{\alpha}_k = \frac{1}{n} \sum_{i=1}^n \alpha_{k+i} \in \mathbb{C}, \ t_\rho = \sum_{k=1}^\rho |\bar{\alpha}_{k+1} - \bar{\alpha}_k|$$

 $heta_{
m t}=0$ : real time

 $A_{\mu} 
ightarrow U A_{\mu} U^{\dagger}, \ U$ : unitary matrix diagonalizing  $A_{0}$ 



•  $T_{ij}(t)$  : order parameter of the SSB

 $(|R^2(t)|$ : extent of space)

#### Complex phase of time



#### Band-diagonal structure



※This structure appears when the space expands.

Locality of time is ensured by the band-diagonal structure.  $\rightarrow$ One can read off the time evolution.



#### SSB of SO(9) rotational symmetry



# 4. Conclusion

- We perform numerical simulations of the type IIB matrix model, which is a candidate for the non-perturbative formulation of superstring theory by using the complex Langevin method (CLM) to overcome the sign problem.
- We add the deformation term (coefficient:  $m_{f}$ ) in the fermionic action to avoid the failure of the CLM.
- In the original model,  $\langle A_{\mu} \rangle$  are complex.  $\rightarrow$  Space-time cannot be real!
- We introduce the mass term (coefficient:  $\gamma$ ) to realize the real space-time.
- igstarrow By simulating at N=48, for  $\gamma=4, m_{
  m f}=3.5\,$  we find that
  - Time: approaching real compared with the  $\gamma = 0$  case
  - Space: real, 3d expanding space, band-diagonal structure

Results shown in this talk are natural from the viewpoint of  $Z = \int dA e^{iS_{\rm E}(A)} {
m Pf} {\cal M}_{\rm E}(A)$ , where  $A_0 = -iA_{10}$ .

However, the distribution of  $\langle \alpha_i \rangle$  shows some departure from being vertical unlike what is suggested from  $A_0 = -iA_{10}$ .

# 4. Outlook

#### After taking the $N ightarrow \infty$ limit, we should take the $\gamma ightarrow 0$ limit.



 $\begin{array}{l} -\langle \alpha_i \rangle \propto e^{-i\frac{3\pi}{8}} \blacklozenge & \text{We expect that real time emerges at late times} \\ (\gamma = 0) & \text{by taking the } N \rightarrow \infty \text{ and } \gamma \rightarrow 0 \text{ limits.} \\ (\text{Flat region of the distribution of } \langle \alpha_i \rangle \text{ appears} \\ & \text{with smaller values of } \gamma.) \end{array}$ 

The expansion of 3d space will be pronounced for smaller values of  $m_{
m f}$ .

(The attractive force among eigenvalues of  $A_{\mu}$ is suppressed by cancellation of the contribution among bosons and fermions. [Aoki-Iso-Kawai-Kitazawa-Tada ('99)])

#### Now, simulations with larger N, smaller $\gamma$ and $m_{\rm f}$ are running.

#### Lorentz-invariant mass term

$$S_{\rm b} = -\frac{1}{4}N\beta \Big[ 2\operatorname{Tr}(F_{0i})^{2} - \operatorname{Tr}(F_{ij})^{2} \Big] - \frac{1}{2}N\gamma \Big[ \operatorname{Tr}(A_{0})^{2} - \operatorname{Tr}(A_{i})^{2} \Big] \qquad \beta = 1/(g^{2}N), \ F_{\mu\nu} = i[A_{\mu}, A_{\nu}]$$

$$\beta > 0, \ \gamma > 0$$

In this case, the equivalence to the Euclidean model is lost!

Cf) Classical solutions in the Lorentzian type IIB matrix model

[KH-Matsumoto-Nishimura-Tsuchiya-Yosprakob ('19)]

The same mass term was used to obtain classical solutions of the Lorentzian type IIB matrix model with expanding behavior:  $\begin{bmatrix} A \nu & A \end{bmatrix}$ 

$$[A^{
u}, [A_{
u}, A_{\mu}]] - {m \gamma} {m A}_{\mu} = 0 \; .$$

If we switch the sign here, we do not obtain such classical solutions with the nice behavior.



#### Complex phase of time



 $N=22,\;n=4,\;m_{
m f}=1.5,\;\eta=0.01$  $ar{lpha}_k = rac{1}{n}\sum_{i=1}^n lpha_{k+i} \in \mathbb{C}, \; t_
ho = \sum_{k=1}^
ho |ar{lpha}_{k+1} - ar{lpha}_k|$  $\Delta \left\langle lpha_i 
ight
angle \equiv \left\langle lpha_{i+1} 
ight
angle - \left\langle lpha_i 
ight
angle \propto e^{i heta_{ ext{t}}}$ •  $\gamma = 0$ Complex phase of  $\langle A_0 
angle \sim e^{-3\pi i/8}$ Corresponds to the dashed line 1<sup>st</sup> order phase transition (?) •  $\gamma = 6$ Real time emerges at late times. (Real-time phase)

#### Band-diagonal structure



%This structure appears when the space expands.

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