## The emergence of $(3+1)$－dimensional space－time in the type IIB matrix model

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## 1. Introduction

To obtain the 4D space-time from superstring theory, which is a promising candidate of the unified theory including quantum gravity, non-perturbative effects of superstring theory are considered to be important.
-The type IIB matrix model [Ishibashi-Kawai-Kitazawa-Tsuchiya ('96)]
A non-perturbative formulation of superstring theory.
In this model, space-time does not exist a priori but emerges
dynamically from the degrees of freedom of matrices.

$$
\begin{aligned}
S & =-N \operatorname{Tr}\left(\frac{1}{4}\left[A^{\mu}, A^{\nu}\right]\left[A_{\mu}, A_{\nu}\right]+\frac{1}{2} \bar{\Psi} \Gamma^{\mu}\left[A_{\mu}, \Psi\right]\right) \\
& =S_{\mathrm{b}}+S_{\mathrm{f}}, S_{\mathrm{b}}=-\frac{N}{4} \operatorname{Tr}\left(\left[A^{\mu}, A^{\nu}\right]\left[A_{\mu}, A_{\nu}\right]\right), S_{\mathrm{f}}=-\frac{N}{2} \operatorname{Tr}\left(\bar{\Psi} \Gamma^{\mu}\left[A_{\mu}, \Psi\right]\right)
\end{aligned}
$$

Related talks by
Yamamori and Piensuk (Wed.) J. Nishimura (Thu.)

$$
\Gamma^{\mu}: \text { 10D Gamma matrices }
$$

$$
(\mu=0, \ldots, 9)
$$

$N \times N$ Hermitian matrices | $\boldsymbol{A}_{\boldsymbol{\mu}}:$ 10D Lorentz vector |
| :--- |
| $\mathbf{\Psi}:$ 10D Majorana-Weyl spinor |

under $\operatorname{SO}(9,1)$ transformation
space-time is extracted from the eigenvalue distribution of $\boldsymbol{A}_{\boldsymbol{\mu}}$.

## Does real space-time emerge?

The type IIB matrix model [ishibashi-Kawai-Kitazawa-Tsuchiya ('96)]
Partition function: $Z=\int d A d \Psi e^{i\left(S_{\mathrm{b}}+S_{\mathrm{f}}\right)}=\int \underset{\text { complex } \rightarrow \text { sign problem! }}{d A e^{i S_{\mathrm{b}}} \operatorname{Pf} \mathcal{M}(A)}$
Definition of expectation values: $\langle\mathcal{O}\rangle=\frac{1}{Z} \int d A \mathcal{O} e^{i S_{\mathrm{b}}} \operatorname{Pf} \mathcal{M}(A)$
Even if $A_{\mu}$ are Hermitian, the expectation values of eigenvalues are complex in general. In the original model, complex phase of $\left\langle A_{0}\right\rangle \sim e^{-3 \pi i / 8}$ and that of $\left\langle A_{i}\right\rangle \sim e^{\pi i / 8}$.
$\rightarrow$ Real space-time cannot be realized!

## How does real space-time emerge in this model?

## Classical solutions

Classical EOM for the type IIB matrix model: $\frac{\delta S}{\delta A_{\mu}}=\left[A^{\nu},\left[A_{\nu}, A_{\mu}\right]\right]=0$ All simultaneously diagonalizable $\boldsymbol{A}_{\mu}$ are classical solutions, so there is no expansion in general for such solutions.
In the previous work [KH-Matsumoto-Nishimura-Tsuchiya-Yosprakob ('19)], to implement the effects of the IR regularization required for the Lorentzian model, we added the quadratic term of $\boldsymbol{A}_{\mu}$ (mass term) $-\frac{1}{2} N \gamma\left[\operatorname{Tr}\left(A_{0}\right)^{2}-\operatorname{Tr}\left(A_{i}\right)^{2}\right]$. Classical EOM with mass term: $\frac{\boldsymbol{\delta} \boldsymbol{S}}{\boldsymbol{\delta} \boldsymbol{A}_{\mu}}=\left[\boldsymbol{A}^{\nu},\left[\boldsymbol{A}_{\nu}, \boldsymbol{A}_{\mu}\right]\right]-\gamma \boldsymbol{A}_{\mu}=\mathbf{0}$

- $\gamma>0$ : classical solutions with an expanding behavior

$\leftarrow$ Classical solution for only 3d space expands. ※Dimensionality of expanding space is arbitrary. But not for $\gamma<0$ !

We introduce the same mass term when we perform simulations of this model.

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## 2. Complex Langevin method (CLM)

We perform numerical simulations based on the complex Langevin method (CLM) to overcome the sign problem.

$$
\boldsymbol{Z}=\int \boldsymbol{d} \boldsymbol{x} \boldsymbol{w}(\boldsymbol{x}), \quad \boldsymbol{x} \in \mathbb{R} \longrightarrow \quad \underset{\text { complex-valued function }}{ } \quad \text { [Parisi ('83), Klauder ('84)] }
$$

Complex Langevin equation ( $t_{\mathrm{L}}$ : Langevin time)

$$
\frac{d z_{k}}{d t_{\mathrm{L}}}=\frac{1}{\underset{w}{w} \frac{\partial w}{\partial z_{k}}+\underset{\text { Gaussian noise, real }}{\eta_{k}\left(t_{\mathrm{L}}\right)} P\left(\eta_{k}\left(t_{\mathrm{L}}\right)\right) \propto \exp \left(-\frac{1}{4} \int d t_{\mathrm{L}} \sum_{k}\left[\eta_{k}\left(t_{\mathrm{L}}\right)\right]^{2}\right) .}
$$

- Necessary and sufficient condition to justify the CLM [Nagata-Nishimura-Shimasaki ('16)]
The probability distribution of the drift term should be exponentially suppressed.



## Application of CLM to the type IIB matrix model

- Order of eigenvalues of $\boldsymbol{A}_{\mathbf{0}}\left(\boldsymbol{\alpha}_{\boldsymbol{i}}\right)$ [Nishimura-Tsuchiya ('19)]

$$
A_{0}=\operatorname{diag}\left(\alpha_{1}, \ldots, \alpha_{N}\right)
$$

$$
\alpha_{1}<\alpha_{2}<\cdots<\alpha_{N} \longmapsto \frac{\alpha_{1}=0, \alpha_{2}=e^{\tau_{1}}, \alpha_{3}=e^{\tau_{1}}+e^{\tau_{2}}, \ldots, \alpha_{N}=\sum_{k=1}^{N-1} e^{\tau_{k}},{ }^{\text {Chift symmetry }}}{}
$$

Complexify variables
$\tau_{k} \in \mathbb{R}$

Hermitian matrices: $\boldsymbol{A}_{\boldsymbol{i}} \in \mathrm{SU}(\boldsymbol{N})$$\Rightarrow$| $\tau_{k} \in \mathbb{C}$ |
| :---: |
| General complex matrices: $\boldsymbol{A}_{\boldsymbol{i}} \in \mathrm{GL}(\boldsymbol{N}, \mathbb{C})$ |

- Complex Langevin equations $\boldsymbol{t}_{\mathrm{L}}$ : Langevin time

$$
\begin{gathered}
\frac{d \tau_{k}}{d t_{\mathrm{L}}}=-\frac{\partial S_{\mathrm{eff}}}{\partial \tau_{k}}+\eta_{k}\left(t_{\mathrm{L}}\right) \\
\left.\begin{array}{ll}
\left.\frac{d\left(A_{i}\right)_{k l}}{d t_{\mathrm{L}}}=-\frac{\partial S_{\mathrm{eff}}}{\partial\left(\boldsymbol{A}_{i}\right)_{l k}}+\left(\eta_{\mathrm{L}}\right)\right) \propto \exp \left(-\frac{1}{4} \int d t_{\mathrm{L}} \sum_{k l}\left[\eta_{k}\left(t_{\mathrm{L}}\right)\right]^{2}\right) & \mathrm{P}\left(\eta_{i}\left(t_{\mathrm{L}}\right)\right) \propto \exp \left(-\frac{1}{4} \int d t_{\mathrm{L}} \sum_{i} \operatorname{Tr}\left[\eta_{i}\left(t_{\mathrm{L}}\right)\right]^{2}\right)
\end{array}\right\} \text { Gaussian noise } \\
S_{\mathrm{eff}}=-\frac{i}{4} N\left(2 \operatorname{Tr}\left[A_{0}, A_{i}\right]^{2}-\operatorname{Tr}\left[A_{i}, A_{j}\right]^{2}\right)-\frac{i}{2} N \gamma\left(\operatorname{Tr}\left(A_{0}\right)^{2}-\operatorname{Tr}\left(A_{i}\right)^{2}\right)-2 \log \Delta(\alpha)-\sum_{k=1}^{N-1} \tau_{k}-\log \operatorname{Pf} \mathcal{M}(A), \Delta(\alpha)=\prod_{k<l}^{N}\left(\alpha_{k}-\alpha_{l}\right) \\
\hline
\end{gathered}
$$

## Some tricks to make the CLM work

Singular drift problem

$$
Z=\int d A e^{i S_{\mathrm{b}}} \operatorname{Pf} \mathcal{M}(A)
$$

$\mathcal{M}$ that appears in $\operatorname{Pf} \mathcal{M}$ has eigenvalues accumulating near zero, which causes the singular drift problem. Due to this problem, the CLM fails. We introduce a deformation term in fermionic action to avoid the problem.
$S_{\mathrm{f}}=-\frac{N}{2} \operatorname{Tr}\left(\bar{\Psi}_{\alpha}\left(\Gamma^{\mu}\right)_{\alpha \beta}\left[A_{\mu}, \Psi_{\beta}\right]+i m_{\mathrm{f}} \bar{\Psi}_{\alpha}\left(\Gamma_{7} \Gamma_{8}^{\dagger} \Gamma_{9}\right)_{\alpha \beta} \Psi_{\beta}\right)$
$\boldsymbol{m}_{\mathrm{f}}$ : deformation parameter $\left\{\begin{array}{l}\boldsymbol{m}_{\mathrm{f}} \rightarrow \infty \\ m_{\mathrm{f}} \rightarrow 0\end{array} \quad \begin{array}{l}\text { : } \begin{array}{l}\text { bosonic } \\ \text { (Due to decoupling of fermionic degrees of freedom) }\end{array} \\ \text { SUSY }\end{array}\right.$
Stabilization
Cf.) [Attanasio-Jäger ('18)] : dynamical stabilization in CL simulation of QCD
To stabilize the CLM, $\boldsymbol{A}_{\boldsymbol{i}} \rightarrow \frac{1}{1+\eta}\left(\boldsymbol{A}_{\boldsymbol{i}}+\eta \boldsymbol{A}_{i}^{\dagger}\right)$ after each Langevin step
$\begin{cases}\boldsymbol{\eta}=0: \text { do nothing } & \text { Here, } \boldsymbol{\eta}=\mathbf{0 . 0 1} \\ \boldsymbol{\eta}=1: \text { Hermitianize } & \text { ※Justifiable when dominant configurations are close to Hermitian. }\end{cases}$

## How to determine whether time and space are real

- Distribution of $\boldsymbol{\alpha}_{i}$ (eigenvalues of $\boldsymbol{A}_{0}$ ) $\boldsymbol{\Delta}\left\langle\alpha_{i}\right\rangle \equiv\left\langle\alpha_{i+1}\right\rangle-\left\langle\alpha_{i}\right\rangle \propto e^{i \theta_{t}}$

Definition of time: $\bar{\alpha}_{k}=\frac{1}{n} \sum_{i=1}^{n} \alpha_{k+i} \in \mathbb{C}, t_{\rho}=\sum_{k=1}^{\rho}\left|\bar{\alpha}_{k+1}-\bar{\alpha}_{k}\right|$
$\boldsymbol{\theta}_{\mathbf{t}}=0$ : real time
$A_{\mu} \rightarrow U A_{\mu} U^{\dagger}, U$ : unitary matrix diagonalizing $A_{0}$


- $\theta_{\mathrm{S}}(t) \quad \theta_{\mathrm{S}}(t)=0$ : real space $\quad R^{2}(t)=\left\langle\frac{1}{n} \operatorname{tr}\left(\bar{A}_{i}(t)\right)^{2}\right\rangle=e^{2 i \theta_{\mathrm{s}}(t)}\left|R^{2}(t)\right|$
- $\boldsymbol{T}_{i j}(t)$ : order parameter of the SSB ( $\left|\boldsymbol{R}^{2}(\boldsymbol{t})\right|$ : extent of space)


## Complex phase of time



Band-diagonal structure

$$
\mathcal{A}_{p q}=\frac{1}{9} \sum_{i=1}^{9}\left|\left(A_{i}\right)_{p q}\right|^{2} \quad N=48, n=16, \gamma=4, m_{\mathrm{f}}=3.5, \eta=0.01
$$



※This structure appears when the space expands.
Locality of time is ensured by the band-diagonal structure. $\rightarrow$ One can read off the time evolution.

Complex phase of space
$N=48, n=16, \gamma=4, m_{\mathrm{f}}=3.5, \eta=0.01 \quad R^{2}(t)=e^{2 i \theta_{\mathrm{s}}(t)}\left|R^{2}(t)\right|$


Almost real space appears.
$\theta_{\mathrm{s}}(t) \sim 0$
For $\gamma=\mathbf{0}$,
$\theta_{\mathrm{s}}(t)=\pi / 8$, which corresponds to complex phase of $\left\langle A_{i}\right\rangle \sim e^{\pi i / 8}$.


## SSB of SO(9) rotational symmetry



## 4. Conclusion

$\checkmark$ We perform numerical simulations of the type IIB matrix model, which is a candidate for the non-perturbative formulation of superstring theory by using the complex Langevin method (CLM) to overcome the sign problem. We add the deformation term (coefficient: $\boldsymbol{m}_{\mathrm{f}}$ ) in the fermionic action to avoid the failure of the CLM.
$\bullet$ In the original model, $\left\langle\boldsymbol{A}_{\mu}\right\rangle$ are complex. $\rightarrow$ Space-time cannot be real!

- We introduce the mass term (coefficient: $\gamma$ ) to realize the real space-time.

By simulating at $N=48$, for $\gamma=4, m_{\mathrm{f}}=3.5$ we find that

- Time: approaching real compared with the $\gamma=0$ case
- Space: real, 3d expanding space, band-diagonal structure

Results shown in this talk are natural from the viewpoint of $Z=\int d A e^{i S_{\mathrm{E}}(A)} \operatorname{Pf} \mathcal{M}_{\mathrm{E}}(\boldsymbol{A})$, where $\boldsymbol{A}_{0}=-i \boldsymbol{A}_{\mathbf{1 0}}$. However, the distribution of $\left\langle\boldsymbol{\alpha}_{i}\right\rangle$ shows some departure from being vertical unlike what is suggested from $\boldsymbol{A}_{\mathbf{0}}=-\boldsymbol{i} \boldsymbol{A}_{\mathbf{1 0}}$.

## 4. Outlook

After taking the $N \rightarrow \infty$ limit, we should take the $\gamma \rightarrow 0$ limit.


Now, simulations with larger $N$, smaller $\gamma$ and $m_{\mathrm{f}}$ are running.

## Lorentz-invariant mass term

$$
\begin{aligned}
& \begin{array}{ll}
S_{\mathrm{b}}=-\frac{1}{4} N \boldsymbol{N}\left[2 \operatorname{Tr}\left(F_{0 i}\right)^{2}-\operatorname{Tr}\left(\boldsymbol{F}_{i j}\right)^{2}\right]-\frac{1}{2} N \gamma\left[\operatorname{Tr}\left(A_{0}\right)^{2}-\operatorname{Tr}\left(A_{i}\right)^{2}\right] & \beta=1 /\left(g^{2} N\right), F_{\mu \nu}=i\left[\boldsymbol{A}_{\mu}, \boldsymbol{A}_{\nu}\right] \\
\beta>0, \gamma>0
\end{array} \\
& A_{0}=e^{-i \frac{3 \pi}{8} u} \tilde{A}_{0}, A_{i}=e^{i \frac{\pi}{8} u} \tilde{A}_{i}\left(\tilde{A}_{\mu}: \text { Hermitian }\right) \\
& \int d A e^{i S_{\mathrm{b}}}=\int d \tilde{A} e^{-\tilde{S}_{\mathrm{b}}} \\
& \text { Real part is positive. } \\
& \tilde{S}_{\mathrm{b}}(\tilde{A})=\frac{1}{4} \boldsymbol{N} \boldsymbol{\beta}\left[2 e^{i \frac{\pi}{2}(1-u)} \operatorname{Tr}\left(\tilde{F}_{\mathbf{0} i}\right)^{2}+e^{-i \frac{\pi}{2}(1-u)} \operatorname{Tr}\left(\tilde{F}_{i j}\right)^{2}\right] \quad \text { Real part is negative. } \\
& \text { For } \mathbf{0}<\boldsymbol{u} \leq \mathbf{1} \\
& +\frac{1}{2} N \gamma\left[e^{\left.-i \frac{\pi}{2}\left(1+\frac{3 u}{2}\right) \operatorname{Tr}\left(\tilde{A}_{0}\right)^{2}+e^{i \frac{\pi}{2}\left(1+\frac{u}{2}\right)} \operatorname{Tr}\left(\tilde{A}_{i}\right)^{2}\right]}\right. \\
& \text { the integration co }
\end{aligned}
$$

## In this case, the equivalence to the Euclidean model is lost!

Cf) Classical solutions in the Lorentzian type IIB matrix model [KH-Matsumoto-Nishimura-Tsuchiya-Yosprakob ('19)]
The same mass term was used to obtain classical solutions of the Lorentzian type IIB matrix model with expanding behavior:

$$
\left[A^{\nu},\left[A_{\nu}, A_{\mu}\right]\right]-\gamma A_{\mu}=0
$$

If we switch the sign here, we do not obtain such classical solutions with the nice behavior.


## Complex phase of time


$N=22, n=4, m_{\mathrm{f}}=1.5, \eta=0.01$
$\bar{\alpha}_{k}=\frac{1}{n} \sum_{i=1}^{n} \alpha_{k+i} \in \mathbb{C}, t_{\rho}=\sum_{k=1}^{\rho}\left|\bar{\alpha}_{k+1}-\bar{\alpha}_{k}\right|$
$\Delta\left\langle\alpha_{i}\right\rangle \equiv\left\langle\alpha_{i+1}\right\rangle-\left\langle\alpha_{i}\right\rangle \propto e^{i \theta_{t}}$

- $\gamma=0$

Complex phase of $\left\langle A_{0}\right\rangle \sim e^{-3 \pi i / 8}$ Corresponds to the dashed line $1^{\text {st }}$ order phase transition (?)

- $\gamma=6$

Real time emerges at late times. (Real-time phase)

## Band-diagonal structure


※This structure appears when the space expands.
Locality of time is ensured by the band-diagonal structure. $\rightarrow$ One can read off the time evolution.

## Complex phase of space and extent of space



At late times, expanding real space emerges.

## SSB of SO(9) rotational symmetry



