# Asymmetric Mediator in Scotogenic Model



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# 1. Introduction -- Motivation



No, There is a relation !!!

Asymmetric Dark Matter (ADM)

Asymmetries are generated simultaneously

Same origin for DM and Baryon

# 1. Introduction – goal for the work

- Dark Matter
- Neutrino Oscillation
- Baryon Asymmetry
- $\Omega_{\rm DM} / \Omega_B \simeq 5$

Scotogenic Model E. Ma , Phys. Rev. D 73 (2006) 077301

> Leptogenesis M. Fukugita and T. Yanagida, Phys. Lett. B174 (1986)45-47

**Asymmetric Dark Mater** 

David E. Kaplan, Markus A. Luty, and Kathryn M. Zurek, Phys. Rev. D.79. (2009) 115016



T. Hugle, M. Platscher, and K. Schmitz,

Phys. Rev. D 98 (2018) 023020

# 2. Our Model

Scotogenic Model E. Ma , Phys. Rev. D 73 (2006) 077301

Standard Model +  $N_i$  (i = 1, 2, 3) (RH neutrinos) +  $\eta$  (SU(2)<sub>L</sub> doublet scalar)

• Symmetry

 $\mathrm{SU}(3)_{\mathcal{C}} \times \mathrm{SU}(2)_L \times \mathrm{U}(1)_Y \times \mathbb{Z}_2$ 

 $\mathbb{Z}_2$ : Stability for DM  $\sigma$ Radiative v mass

	フェルミオン場			スカラー場		
場	L	$e_R$	N	H	$\eta$	$\sigma$
$\overline{\mathrm{SU}(2)_L}$	2	1	1	2	2	1
$Z_2$	+	+	_	+	_	_

#### **Our Model**



 $\sigma$  is assumed to be a few GeV

• Lagrangian
$$\mathcal{L} \supset -h_{lpha i} ar{L}_{lpha} ilde{\eta} N_i + rac{1}{2} M_i$$

Yukawa

Majorana mass

$$V(H,\eta,\sigma) \supset \frac{\lambda_8}{2} \left[ (H^{\dagger}\eta)^2 + h.c. \right] + \frac{\mu}{\sqrt{2}} \left[ \sigma(H^{\dagger}\eta) + h.c. \right]$$
  
Radiative seesaw DM

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# 2. Our Model

Z<sub>2</sub> forbid Dirac mass instead coupling with  $\eta$  $X : h_{\alpha i} \overline{L_{\alpha}} \widetilde{H} N_i \longrightarrow O : h_{\alpha i} \overline{L_{\alpha}} \widetilde{\eta} N_i$ Mass generation at1-loop

Radiative generation



Neutrino mass matrix

$$(\mathcal{M}_{\nu})_{\alpha\beta} \simeq \frac{\lambda_8 v^2}{32\pi^2} \sum_i \frac{h_{\alpha i}^* h_{\beta i}^*}{M_i} \begin{bmatrix} \ln \frac{M_i^2}{m_\eta^2} - 1 \end{bmatrix} \qquad \begin{array}{c} M_i \\ m_\eta \\ v \colon \end{array}$$

 $M_i$  : RH neutrino mass  $m_\eta$  : SU(2)<sub>L</sub> doublet mass v : Higgs VEV

#### $\lambda_8$ is key to generate neutrino mass

## 2. Our Model

#### **Leptogenesis**: Lepton Number creation with RH v Decay





## 3. How to "create" Dark Matter



#### Connecting Baryon # and DM # via $\Delta\eta$



Mass of  $\sigma \sim$  GeV

## 3. How to "create" Dark Matter



### What are important in our scenario



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## 4. Condition for Baryon Number

Approcimation for B #

 $\eta_B \simeq -0.01 \epsilon_1 \kappa_1$ 

 $\epsilon_1$  : asymmetry parameter  $\kappa_1$  : efficiency factor

 $\epsilon_1$  and  $\kappa_1$  are calculated by Yukawa

Casas-Ibarra parametrization

$$h_{\alpha i} = \left( U D_{\nu}^{\frac{1}{2}} R^{\dagger} D_{\Lambda}^{\frac{1}{2}} \right)_{\alpha i} \quad (\mathcal{D}_{\Lambda})_{ii} = \frac{2\pi^2}{\lambda_8} \xi_i \frac{2M_i}{v^2}$$

$$\epsilon_1$$
 is given by  $(r_{ij} = M_j/M_i, \eta_i = m_\eta/M_i)$   
 $\epsilon_i = \frac{1}{8\pi} \frac{1}{(h^{\dagger}h)_{ii}} \sum_{j \neq i} \operatorname{Im} \left[ \left\{ (h^{\dagger}h)_{ij} \right\}^2 \right] \frac{1}{\sqrt{r_{ji}}} F(r_{ji}, \eta_i)$ 

 $\kappa_1$  is given by

$$\kappa_1(K_1) \simeq \frac{1}{1.2K_1 \left[\ln K_1\right]^{0.8}}$$

$$K_1 = \frac{1}{8\pi} \sqrt{\frac{90}{8\pi^3 g_*}} \frac{M_{\rm Pl}}{M_1} \left(h^{\dagger} h\right)_{11} \left(1 - \eta_1\right)^2$$

if  $K_1 > 1$ , this is well satisfied

#### **R** complex orthogonal matrix

# 4. Condition for Baryon Number

 $(1) No \eta \eta \to HH$ 

$$<->\Gamma_{\eta\eta\to HH} < H(T=m_{\eta})$$

<->  $\lambda_8 < 3.9 imes 10^{-8} \sqrt{m_\eta/\text{GeV}}$ 

 $\lambda_8$  contributes neutrino masses

$$(\mathcal{M}_{\nu})_{\alpha\beta} \simeq \frac{\lambda_8 v^2}{32\pi^2} \sum_i \frac{h_{\alpha i}^* h_{\beta i}^*}{M_i} \left[ \ln \frac{M_i^2}{m_\eta^2} - 1 \right]$$

Select parameters so that neutrino oscillation can be explained

Blue: 
$$\eta_B$$
  
Black:  $\eta_B^{obs} = 6.1 \times 10^{-10}$ 



## 4. Conditions for pair annihilation of $\eta$

 $2\eta$  pair-annihilates sufficiently

 $Y^{\mathrm{f}}_{\eta} \ll Y_{\Delta\eta}$ 

Note  $n_\eta < n_{\overline{\eta}}$ 

**Estimate with following approximation** 

$$\boldsymbol{Y_{\eta}^{f}} \equiv \frac{\boldsymbol{n_{\eta}^{f}}}{\boldsymbol{s}} = 2 \times \frac{3.80 \, x_{f}}{\left(g_{*s}/g_{*}^{1/2}\right) M_{\mathrm{Pl}} m_{\eta} \left\langle \sigma_{\mathrm{g}} v_{\mathrm{rel}} \right\rangle}$$

$$x_{\rm f} \equiv \frac{m_{\eta}}{T_{\rm f}} = \ln \left[ 0.038 \left( g/g_*^{1/2} \right) M_{\rm Pl} m_{\eta} \left\langle \sigma_{\rm g} v_{\rm rel} \right\rangle \right] - \frac{1}{2} \ln \left\{ \ln \left[ 0.038 \left( g/g_*^{1/2} \right) M_{\rm Pl} m_{\eta} \left\langle \sigma_{\rm g} v_{\rm rel} \right\rangle \right] \right\}$$



## 4 . Condition for pair annihilation of $\eta$



#### 4. Condition for asymmetric Dark Matter



(3) 
$$\sigma$$
 is produced finally by decay of  $\eta$   

$$m_{f} > T > T_{BBN}$$

$$T_{f} > T > T_{BBN}$$

$$T_{f} : \text{freez-out of pair annih.}$$

$$T_{BBN} : \text{Big-bang Nuclepsynthesis}$$
As a whole  

$$m_{\eta} \sim 10^{4} \text{GeV}, \quad \lambda_{8} < 10^{-8} \left(\frac{m_{\eta}}{\text{GeV}}\right)^{\frac{1}{2}}, \quad 10^{-11} \left(\frac{m_{\eta}}{\text{GeV}}\right)^{\frac{1}{2}} < \frac{\mu}{\text{Gev}} < 10^{-10} \left(\frac{m_{\eta}}{\text{GeV}}\right)^{\frac{3}{2}}$$

# 5. Summary



We study a Scotogenic Model with a real scalar

**1** Decay of RH  $N_1$  : Lepton # generation

(2) The mediator  $\eta$  pair-annihilate with keeping asymmetry  $n_{\Delta\eta}$ 

**(3)** DM  $\sigma$  is produced by  $\eta$  decay

#### Summay

Dark Matter Neutrino Oscillation

Baryon Asymmetry Asymmetric DM condition  $\Omega_{\rm DM}$  /  $\Omega_B \simeq 5$ 

can be simultaneously explained in one framework !