

The Petz (lite) recovery map for scrambling channel

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Based on : 2310.18491 (hep-th)

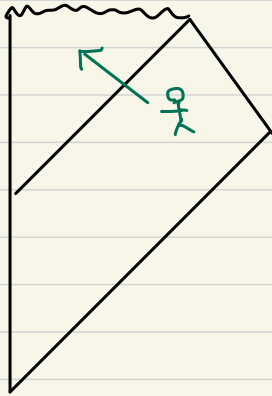
with A. Miyata (KITS), Y. Nakayama (Kyoto)

+ Work in Progress

with R. Myers (Perimeter) and S. Ruan (YITP)

BH Complementarity : Two equivalent observers in QG.

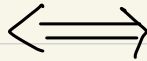
Infalling observer



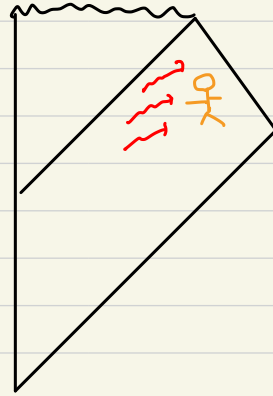
See BH interior

Interior geometry

Equivalent



Distant observer

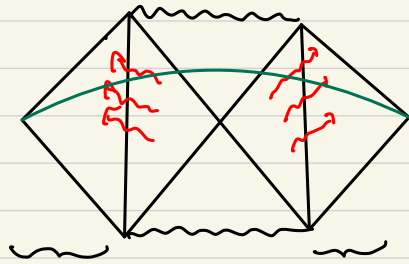


Collecting Hawking quanta

Entanglement

Can't see the BH interior and Hawking radiation by the same observer

Semiclassically they appear to be independent

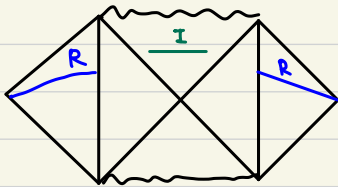


Non gravitating
Heat bath

: AdS BH with non gravitating heat Bath

- In a Cauchy Slice, BH interior looks independent from d.o.f in the 'radiation region'

The island formula



$$S(p_R) = \min_{\text{I}} \text{ext} \left[\frac{A(\partial \text{I})}{4G} + S_{\text{bulk}}^{\text{dFT}}(\text{I} \cup \text{R}) \right]$$

⇒ The island region (in the BH interior) and the radiation region are not independent

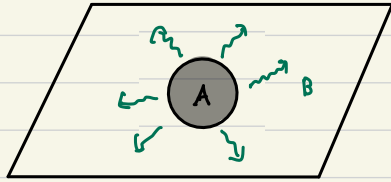
⇒ BH interior region is reconstructable from Hawking radiation

Talk today is about

This reconstruction protocol

- in
- (1) the SYK model
 - (2) a Holographic DCFT
(a doubly holographic setup)

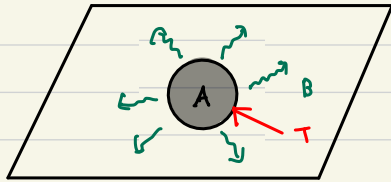
A toy model: the Hayden Preskill setup



- A BH after the Page time (old BH)
(maximally entangled with HR)

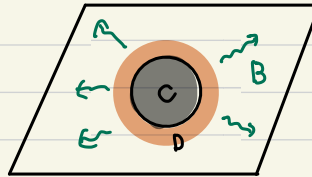
$$|\Psi\rangle = \frac{1}{\sqrt{d_A}} \sum_{i=1}^{d_A} |\psi_i\rangle_A \otimes |i\rangle_B \equiv |\text{EPR}\rangle_{AB}$$

Throwing a diary into the BH



B: early radiation

BH evaporation
⇒



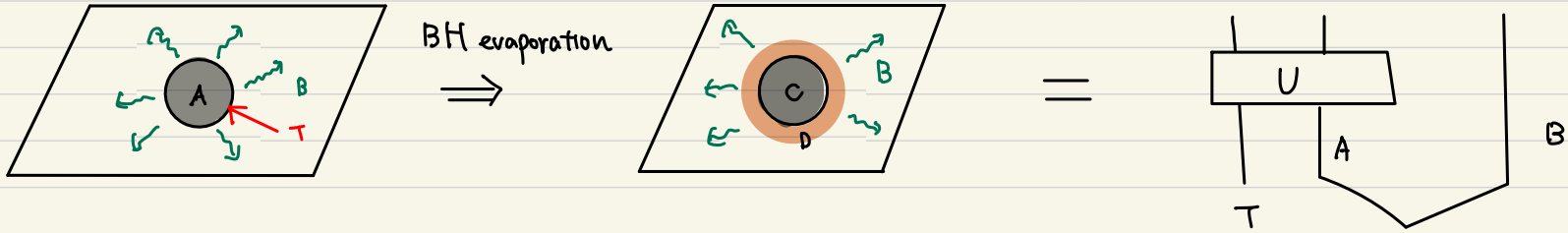
C: remaining BH : H_C

D: late radiation : H_D

$$|\Psi\rangle_{CDB} = |\text{EPR}\rangle_{AB} \otimes |T\rangle_T \Rightarrow I_B \otimes U_{CDE=AT} |\text{EPR}\rangle_{AB} \otimes |T\rangle_T$$

$U_{AT \rightarrow CD}$: a Unitary matrix modeling evaporation dynamics $H_A \otimes H_T \rightarrow H_C \otimes H_D$

Circuit diagram. [Hayden Preskill]



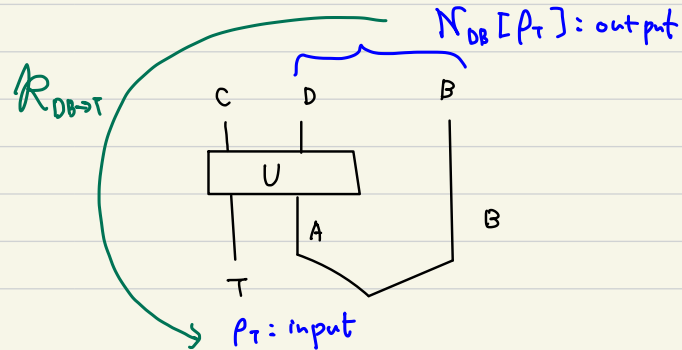
This defines a map $N: T \rightarrow DB$

$$P_T \Rightarrow \text{tr}_C (I_B \otimes U_{AT \rightarrow CD}) (P_T \otimes |EPR\rangle_{AB} \langle EPR|) (U_{AT \rightarrow CD}^\dagger \otimes I_B) \equiv N_{T \rightarrow DB} [P_T]$$

o Reconstructing BH interior from HR

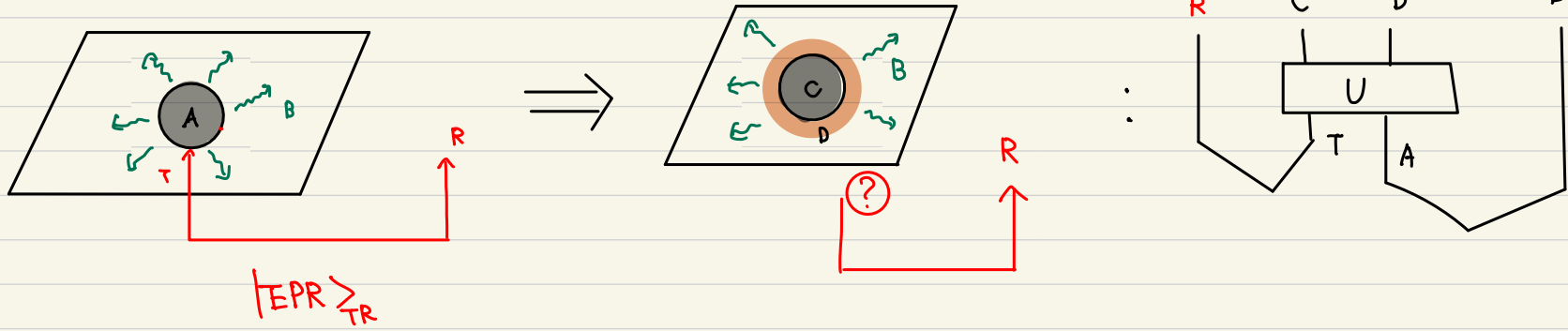
\Leftrightarrow Recovery map: $DB \rightarrow T$

$$R_{DB \rightarrow T} \circ N_{T \rightarrow DB} = 1_T$$



Where is the diary?

- Entangling T with a reference system R



- Information of the diary is in the Hilbert space R is entangled with

Decoupling thm

$$\int dU \|\rho_{Rc} - \rho_R \otimes \rho_c\|_1 \leq \left(\frac{d_T}{d_D}\right)^2$$

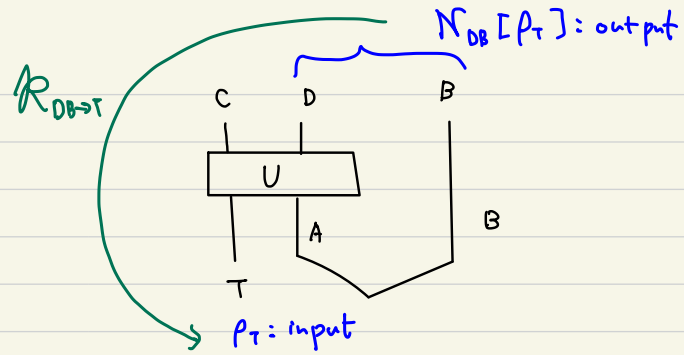
\Rightarrow If $d_T < d_D$ then, $\rho_{Rc} = \rho_R \otimes \rho_c$

$\Rightarrow R$ is mostly entangled with

DB : Hawking radiation !!

Petz recovery map

For a given N which is recoverable,
The recovery map is



$$\mathcal{R}_\sigma[P_{DB}] = \sigma_T^{1/2} N^\dagger \left[N^{-1/2} [P_T] \rho_{DB} N^{-1/2} [P_T] \right] \sigma_T^{1/2}$$

Petz recovery map

$\sigma_T =$ a full rank density matrix on T . $\sigma_T = \frac{\mathbb{1}}{d_T}$

Naturally appear by regarding the setup as a

quantum error correcting code (knill Lafflane)

Petz gets Simplified in a chaotic system

- Scrambling: Information of σ_T spread over entire BD after the scrambling time,

$$N[\sigma_T] \sim \frac{1}{d_0 d_B} I_{DB}$$

$$\mathcal{R}_\sigma[\rho_{DB}] = \sigma_T^{1/2} N^\dagger \left[N^{-1/2}[\sigma_T] \rho_{DB} N^{-1/2}[\sigma_T] \right] \sigma_T^{1/2} \sim N^\dagger[\rho_{DB}]$$

We check this in

1. HCP setup
2. the SYK model

$$S(N^\dagger N[\rho] \parallel \rho) = 0$$

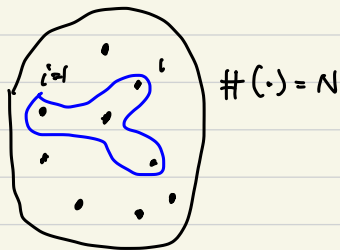
$$N^\dagger N = 1 \text{ on } H_{\text{code}}$$

the Recovery map in the SYK model

The SYK model

0+1 dimensional theory of N Majorana fermions

$$H = (i)^{\frac{q}{2}} \sum_{1 \leq i_1 \leq i_2 \dots \leq i_q \leq N} J_{i_1 \dots i_q} \psi_{i_1} \dots \psi_{i_q}$$

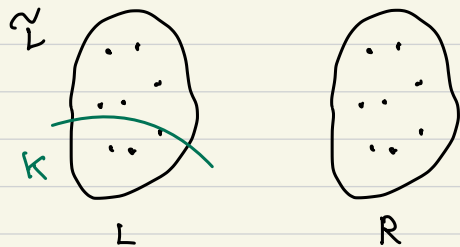


$J_{i_1 \dots i_q}$: Random coupling drawn from Gaussian

$$\langle J_{i_1 \dots i_q}^2 \rangle = \frac{J^2}{(q-1)!} N^{q-2}$$

- Maximally chaotic (OTOCs grow fastest)
- Solvable in the large N limit
- Low energy effective action coincides with the boundary action of JT gravity

Hayden Preskill in the SYK



Two copies of the SYK

L

\Leftrightarrow

Original BH A

R

\Leftrightarrow

Early radiation B

$$|TFD\rangle_{LR} = \frac{1}{\sqrt{2}} \sum e^{i\frac{\beta E}{2}} |E_n\rangle |E_n\rangle \quad \Leftrightarrow$$

$|EPR\rangle_{AB}$

K (ψ_i $i=1 \dots k$) \Leftrightarrow

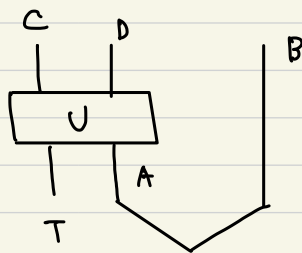
Late radiation

\tilde{L} (ψ_i $i=k, \dots, N$) \Leftrightarrow

Remaining BH

Subspace: $\text{Span} \{ |\beta\rangle, \psi_i \tilde{L} |\beta\rangle \} \quad \Leftrightarrow$

$H_T = \text{Span} \{ |0\rangle, |1\rangle \}$



the HP setup

We expect: $\mathcal{R}_{RK \rightarrow T}^{STK} = a N_{RK \rightarrow T}^{\dagger STK}$

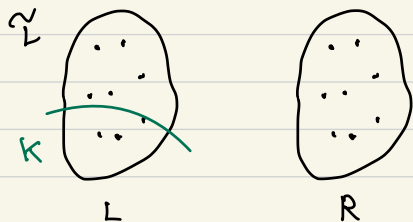
a : constant to ensure $\text{tr}[\mathcal{R}^{STK}] = 1$

check: $\langle T^1 | \mathcal{R} \cdot N [|T^2\rangle \langle T^3|] |T^4\rangle = \langle T^1 | T^2 \rangle \langle T^3 | T^4 \rangle$

We do this by writing the LHS in terms of a correlator and compute it in the large N limit.

$$\langle 1 | N_{KR \rightarrow T}^{\dagger} N_{T \rightarrow KR} [|0\rangle \langle 0|] |1\rangle \propto \langle \text{TFD} | \psi_{i,L}(t-i\epsilon) (\mathbb{I}_L \otimes \rho_{KR}) \psi_{i,L}(t+i\epsilon) | \text{TFD} \rangle$$

Modular flowed correlator



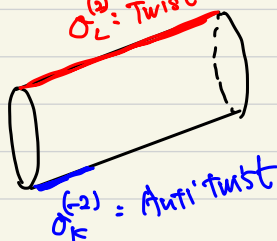
$$\rho_{KR} = \text{tr}_L [| \text{TFD} \rangle \langle \text{TFD} |]$$

The result

$$\langle \text{TFD} | \psi_{iL}(t-is) (I_L \otimes \rho_{KR}) \psi_{iL}(t+is) | \text{TFD} \rangle$$

$$= \frac{G_{2\beta}(2\beta-2s)}{G_{2\beta}(2s)} \left[1 - \frac{G_{2\beta}(\beta)}{\sin\left(\frac{\pi s}{\beta}\right)} \left(\frac{2\beta J}{g^2 C} \right) \frac{k}{N} e^{\frac{\pi}{\beta} t} \right]$$

Similar to OTOC: $\langle \psi(\tau_1) \psi(\tau_2) \psi(\tau_3) \psi(\tau_4) \rangle$ with $\tau_1 > \tau_3 > \tau_2 > \tau_4$

$$\langle \beta | \psi_{iL} \rho_{KR} \psi_{iL} | \beta \rangle =$$


$$\sigma_K^{(-2)} = 1 + \frac{k}{N} \sum_{i=1}^k \psi_{iK} \psi_{iK} + \dots$$

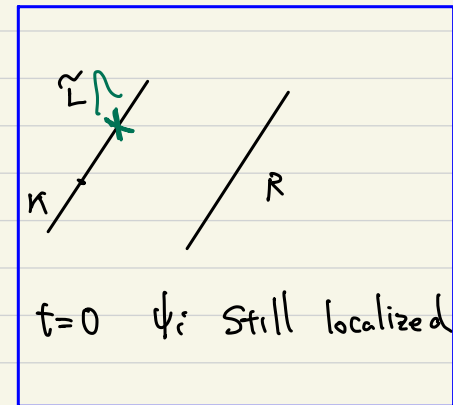
$$\langle \beta^{\otimes 2} | \psi_{iL} \sigma_L^{(2)} \sigma_K^{(-2)} \psi_{iL} | \beta^{\otimes 2} \rangle \Rightarrow \langle \beta^{\otimes 2} | \psi_{iL} \psi_{jK} \psi_{jK} \psi_{iL} | \beta^{\otimes 2} \rangle$$

an Interpretation

$$\langle 1 | N_{KR \rightarrow T}^\dagger N_{T \rightarrow KR} [|0\rangle_T \langle 0|] |1\rangle = \frac{G_{2\beta}(2\beta - 2S)}{G_{2\beta}(2S)} \left[1 - \frac{G_{2\beta}(\beta)}{\sin\left(\frac{\pi S}{\beta}\right)} \left(\frac{2\beta J}{g^2 C} \right) \frac{K}{N} e^{\frac{\pi}{\beta} t} \right]$$

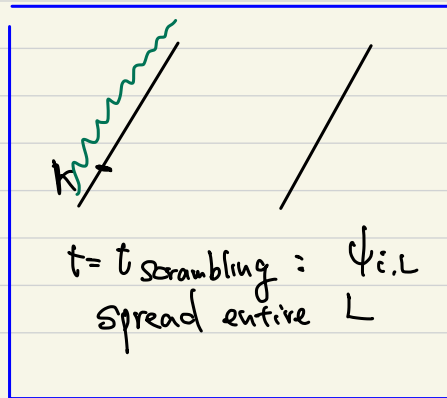
$t \sim 0 \neq 0 \Rightarrow N^\dagger N \neq 1$: recovery is not possible

$t \sim \frac{\beta}{\pi} \log \frac{N}{K} = 2 t_{\text{scrambling}} \Rightarrow N^\dagger N = 1$: recovery is possible



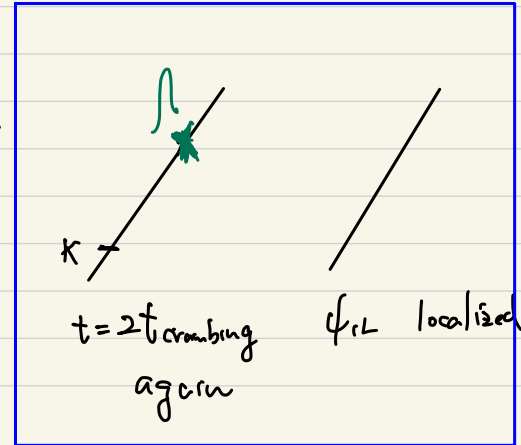
Spreading

\Rightarrow



recovery

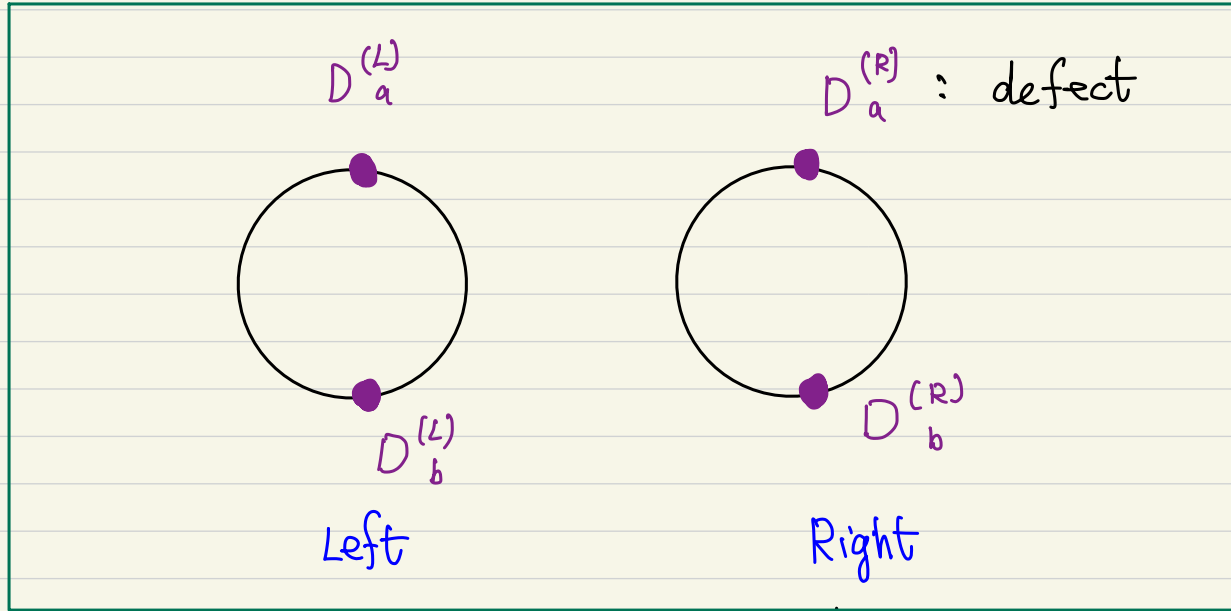
\Rightarrow



HP decoupling in 1+1 D DCFT

1+1 D DCFT setup

Two copies of 2d CFT on $S^1 \times \mathbb{R}$ with two defects

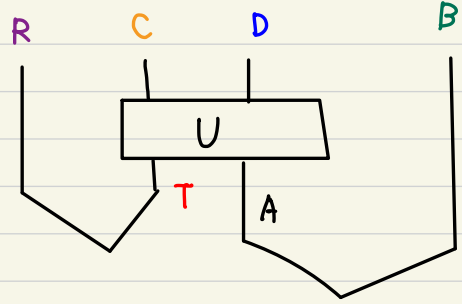
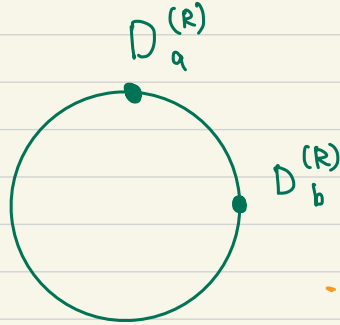
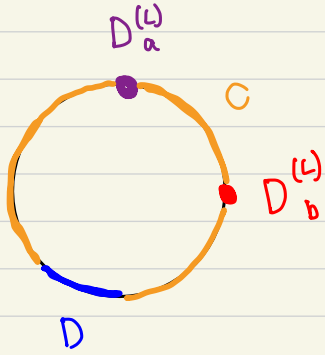


- Each defect is interacting with ambient d.o.f
- L and R form a TFD state

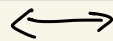
Identification with the HP setup

DCFT

HP

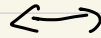


The right system



Early Hawking radiation

$D_a^{(L)}$



The reference system

$D_b^{(L)}$



The diary thrown into the BH

C: the ambient d.o.f

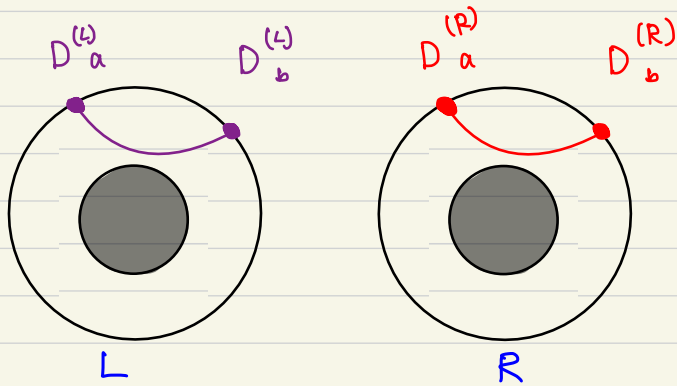


The remaining BH

the Holographic description in AdS_3

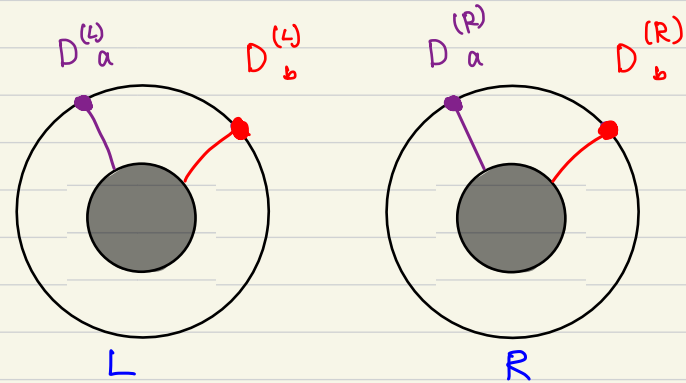
Two defects are connected by a brane in the bulk.

Warm phase



Two defects on the same boundary are connected in the bulk

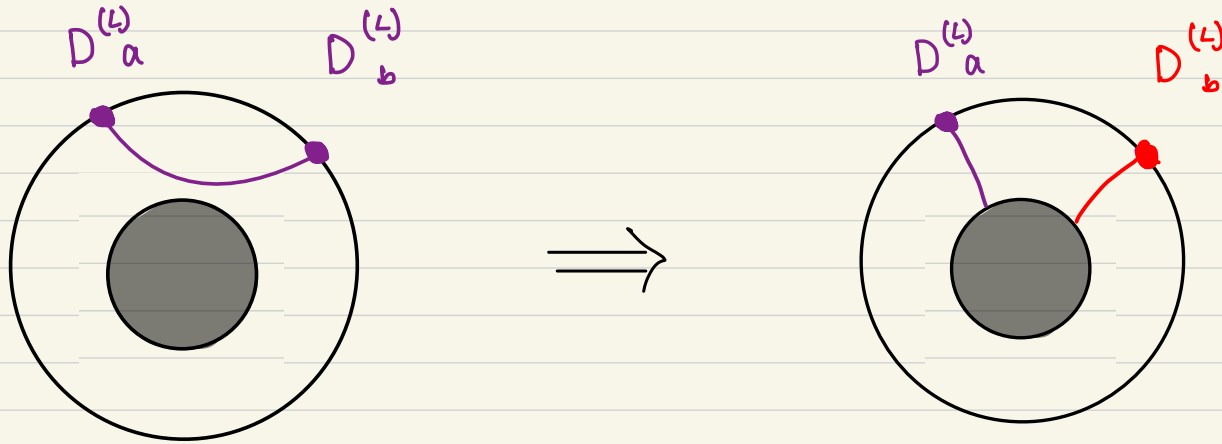
Hot phase



Two defects on the opposite boundary are connected in the bulk

The phase structure

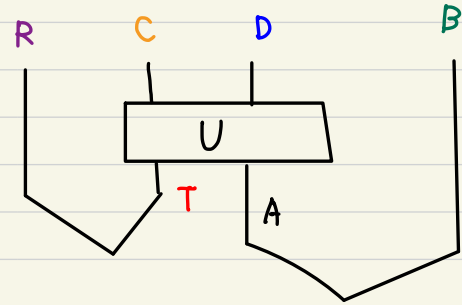
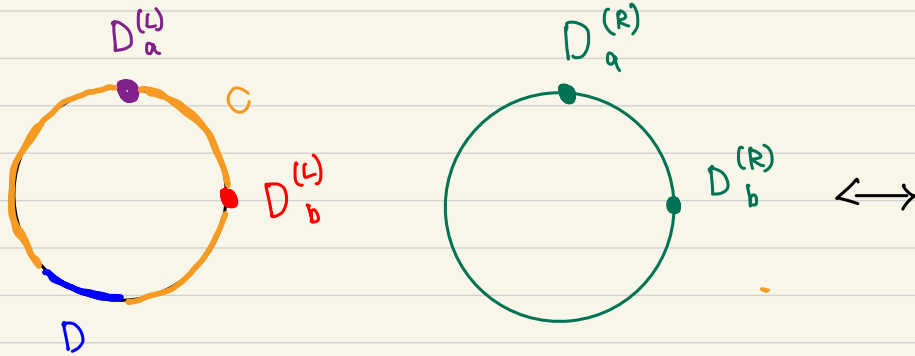
The dimensionless parameter : $\tau = T \times L_2$
temperature The distance between two defects



$\tau \ll 1$: the Warm phase dominates.

$\tau \gg 1$: the Hot phase dominates

Decoupling in the DCFT Setup



$$I_{RC} = S_R + S_C - S_{RC} = 0$$

\Leftrightarrow

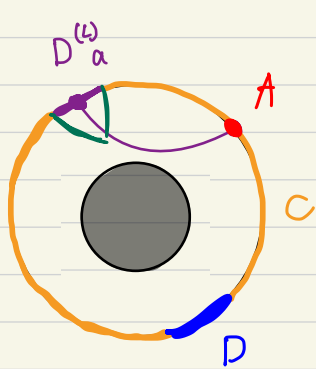
$$\rho_{RC} = \rho_R \otimes \rho_C$$

Vanishing of the mutual information

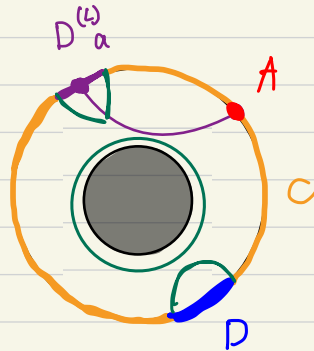
Decoupling

Holographic calculation of I_{RC} : Warm phase

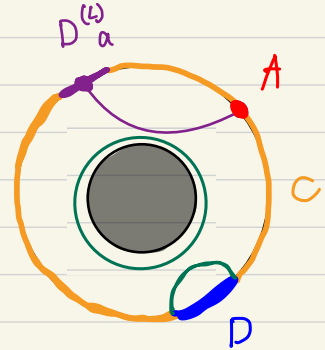
Each entropy is computed by the appropriate geodesic length in the bulk



$$S_R = 2 \log g$$



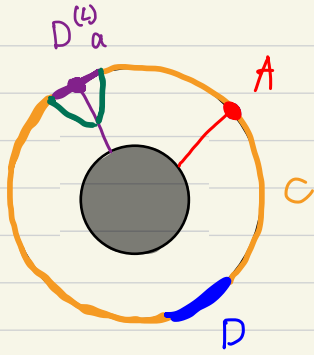
$$S_C = 2 \log g + S_{th} + S_{EE,\beta}(D)$$



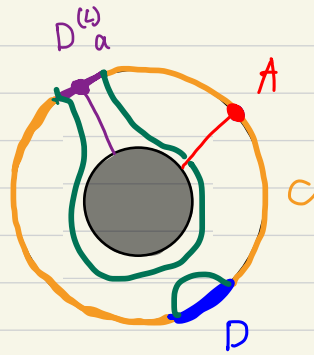
$$S_{RC} = S_{th} + S_{EE,\beta}(D)$$

$\Rightarrow I_{RC} = 4 \log [g]$: Decoupling is **not** achieved in the Warm-phase

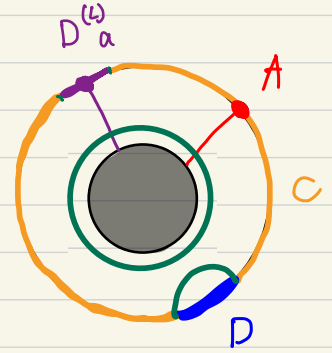
Holographic calculation of I_{RC} : Hot phase



$$S_R = 2 \log g$$



$$S_C = S_{th} + 2 \log g + S_{p,EE}(D)$$



$$S_{RC} = S_{th} + 4 \log g + S_{p,EE}(D)$$

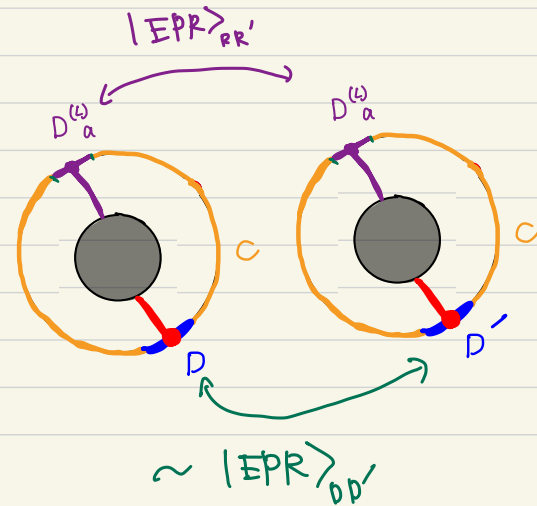
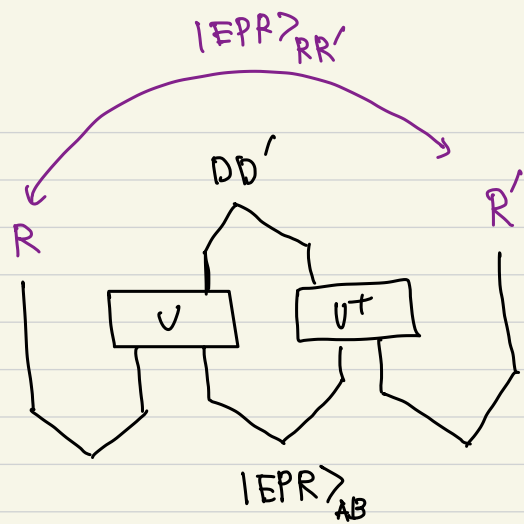
$I_{RC} = 0$: Decoupling is archived in the Hot phase

Recovery map in the DCFT setup

The Petz recovery map is equivalent to the Yoshida-Kitaev protocol

- Prepare a copy of $|EPR\rangle_{RT}$
- Simulate the BH dynamics
- Post select to $|EPR\rangle_{DD'}$

◦ The Yoshida-Kitaev has a natural DCFT realization when $D_b^{(L)} \in D$



Summary

- We studied the Petz recovery map in chaotic system
 - First step to reconstruct BH interior from Hawking radiation
- We found the expression of the recovery map significantly gets simplified in chaotic system
 - { the Hayden Preskill model
 - { the SYK model

Future work

- Generalizations to higher D (Brane world setup)
- de Sitter (Preliminary result: 2308.09748)

with Balasubramanian, Nomura