



VARIANCE REDUCTION FOR THE LATTICE HVP

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TABLE OF CONTENTS

1. Low-mode averaging (LMA) and its variants
2. Multigrid / Deflation
3. Multigrid low-mode averaging (MG LMA)
4. Where is the variance?
5. Cost
6. Conclusion



LOW-MODE AVERAGING (LMA) AND ITS VARIANTS

MOTIVATION - LOW-MODE AVERAGING (LMA)

- **Idea** [Neff et al. hep-lat/0106016, DeGrand and Schaefer hep-lat/0401011, Giusti et al. hep-lat/0402002]: Decompose the quark propagator into two pieces
 - ▶ One piece: should contain **most of the variance**
 - ▶ Other piece: **negligible variance**
- Determine N_c lowest modes of $D, Q = \gamma^5 D$, eo-preconditioned D, Q
- Write $S = D^{-1} = \text{truncated spectral/singular sum} + \text{remainder}$

$$Q^{-1} = \underbrace{\sum_{i=1}^{N_c} \frac{1}{\lambda_i} \xi_i \xi_i^\dagger}_{Q_{LMA}^{-1}} + \underbrace{P Q^{-1} P^\dagger}_{Q_{rest}^{-1} = Q^{-1} - Q_{LMA}^{-1}}, \quad (1)$$

with

$$Q\xi_i = \lambda_i \xi_i, \quad |\lambda_i| = \text{small}, \quad P = 1 - \sum_{i=1}^{N_c} \xi_i \xi_i^\dagger.$$

THE TWO-POINT CORRELATOR

- Two-point connected light-quark vector correlator
- In the time-momentum representation [[Bernecker and Meyer 1107.4388](#)] (local-local), $S = D^{-1}$

$$G(t) = \frac{1}{|\Omega_0|} \sum_{y \in \Omega_0} \sum_{\vec{x} \in \Sigma_0} C(y_0 + t, \vec{x}|y), \quad (2)$$

$$C(x|y) = \text{tr} [\Gamma_1 S(x|y) \Gamma_2 S(x|y)^\dagger], \quad (3)$$

- Stochastic sources: introduce extra noise
- Point sources: costs L^3
- Ideally, but unrealistic: full lattice volume average

THE TWO-POINT CORRELATOR WITH LMA

- Plug in decomposition of propagator

$$G(t) = G_{ee}(t) + \underbrace{G_{re}(t) + G_{er}(t)}_{G_x(t)} + G_{rr}(t) \quad (4)$$

- Get 3-4 terms: eigen-eigen, cross (rest-eigen + eigen-rest), rest-rest

$G_{ee}(t)$: exact, volume-averaged, at its gauge noise

$G_{rr}(t)$: little variance contribution \rightarrow few sources

$G_x(t)$: 10-30% contribution to total noise \gg gauge noise

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THE VARIOUS PROBLEMS OF LMA

1. **V^2 -problem:** number of required low modes scales $O(V)$ with the volume, on state-of-the-art lattices at the physical point
 - ▶ **1000-6000 eigenmodes** [Kuberski 2312.13753, Blum et al. 1801.07224, Borsanyi et al. 1711.04980, Blum et al. 1512.09054]
 - ▶ Memory requirements
 - ▶ Storage and I/O requirements (people don't store them anymore!)



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Note

Number of eigenmodes are limited by memory / resources.



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2. **Cross-term-problem:** Cross term has lots of noise → expensive!
 - ▶ **Method 1:** all-mode averaging, AMA, [Blum et al. 1208.4349, Shintani et al. 1402.0244, Blum et al. 1801.07224, Blum et al. 1512.09054]
 - ▶ **Method 2:** truncated solver method (TSM) + bias correction [Kuberski 2312.13753, Borsanyi et al. 1711.04980]
 - ▶ **Method 3:** stochastically evaluate the rest-eigen piece
 - ▶ ...

MULTIGRID / DEFLATION

LOCAL COHERENCE / WEAK APPROX. PROPERTY

- Low modes of Dirac operator are locally coherent [[Luscher 0706.2298](#)]

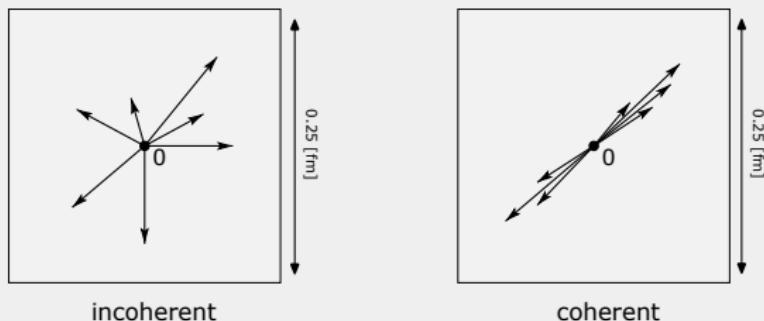


Figure: (Local) coherence of low modes (taken from Ref. [[Luscher 1002.4232](#)]).

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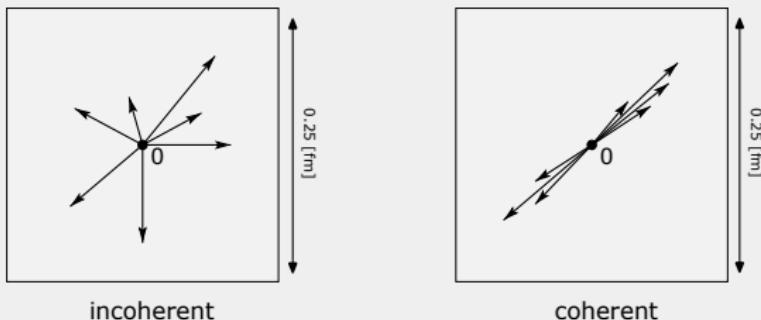


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Conclusion

Using domain decomposition / coarsening on 10-100 low modes is enough to span the $O(V)$ low-mode space!

MULTIGRID / DEFLATION

- Setup subspace(s) as in the previous slide (domain-decomposed low modes)
- Define restrictors R and prolongators T from/to these subspaces

$$R: \psi \mapsto \theta, \quad \theta(i) = \langle \phi_i | \psi \rangle, \quad (5)$$

$$T: \theta \mapsto \psi = \sum_i \theta(i) \phi_i, \quad (6)$$

- Define the **coarse-grid Dirac operator(s)** as $D_c = RDT$

$$\boxed{D_c} = \boxed{R} \cdot \boxed{D} \cdot \boxed{T}$$

- Connection to solver: sloppy D_c^{-1} as preconditioner for the Dirac equation

$$LD\psi = L\eta \quad \text{with} \quad L = TD_c^{-1}R \quad (\text{left preconditioning})$$

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Main message

- De

Coarse-grid operator has smaller dimension, smaller condition number and is thus **cheaper to invert!**

- Connection to solver: sloppy D_c^{-1} as preconditioner for the Dirac equation

$$LD\psi = L\eta \quad \text{with} \quad L = TD_c^{-1}R \quad (\text{left preconditioning})$$



MULTIGRID LOW-MODE AVERAGING (MG LMA)

PROPAGATOR

- Decompose the quark propagator $S = D^{-1}$ using the coarsenings

$$S = \sum_{i=0}^{N-1} S_i = \underbrace{S - K_1}_{=S_0} + \underbrace{K_1 - K_2}_{=S_1} + \underbrace{K_2 - K_3}_{=S_2} + \cdots + \underbrace{K_{N-1}}_{=S_{N-1}}, \quad (7)$$

$K_i = T_i(D_{c,i})^{-1}R_i$, S_i = deflated propagator on level i.

- Each level is defined by a different domain decomp./coarse grid



TWO-POINT CORRELATOR

- Plug into the correlator
- For the correlator we find a **matrix of correlators**:

$$C_{ij}(x,y) = \text{tr} [\Gamma_1 S_i(x|y) \Gamma_2 S_j(y|x)], \quad C = \sum_{i,j} C_{ij}. \quad (8)$$

- $i, j = 0, \dots, N-1$ correspond to **MG-level** (with 0 the fine grid)
- Grouping the N^2 correlators into levels (see figure on next slide) gives us

$$G(t) = \sum_{k=0}^{N-1} G_{Lk}(t). \quad (9)$$

GROUPING OF CORRELATORS

C_{00}	C_{10}	C_{20}	C_{30}
C_{01}	C_{11}	C_{21}	C_{31}
C_{02}	C_{12}	C_{22}	C_{32}
C_{03}	C_{13}	C_{23}	C_{33}

C_{rr}	C_{re}
C_{er}	C_{ee}

$$G = G_{L0} + G_{L1} + G_{L2} + G_{L3}$$

$$G = \underbrace{G_{rr} + G_x}_{G_{L0}} + \underbrace{G_{ee}}_{G_{L1}}$$

- Each level-contribution can be evaluated with a different strategy, i.e. number and type of sources!

Main message

Evaluating G_{Lk} requires inversions of the Dirac operator $D_{c,k}$ on level k and coarser, but not finer levels!



WHERE IS THE VARIANCE?

ENSEMBLES

Name	Size [$T \times L^3$]	L [fm]	$m_\pi L$
E7 ¹	64×32^3	2.1 fm	3.2
F7 ²	96×48^3	3.2 fm	4.8
G7 ¹	128×64^3	4.2 fm	6.4
H7 ¹	192×96^3	6.3 fm	9.6

Table: All ensembles have a pion mass $m_\pi = 270$ MeV and a lattice spacing of $a = 0.0658$ fm with $N_f = 2$ $O(a)$ -improved Wilson fermions.

¹Generated by Tim Harris using openQCD 2.4.2 [Lüscher et al. (2012-2023)]

²CLS lattice from Ref. [CLS (2012-2023)]

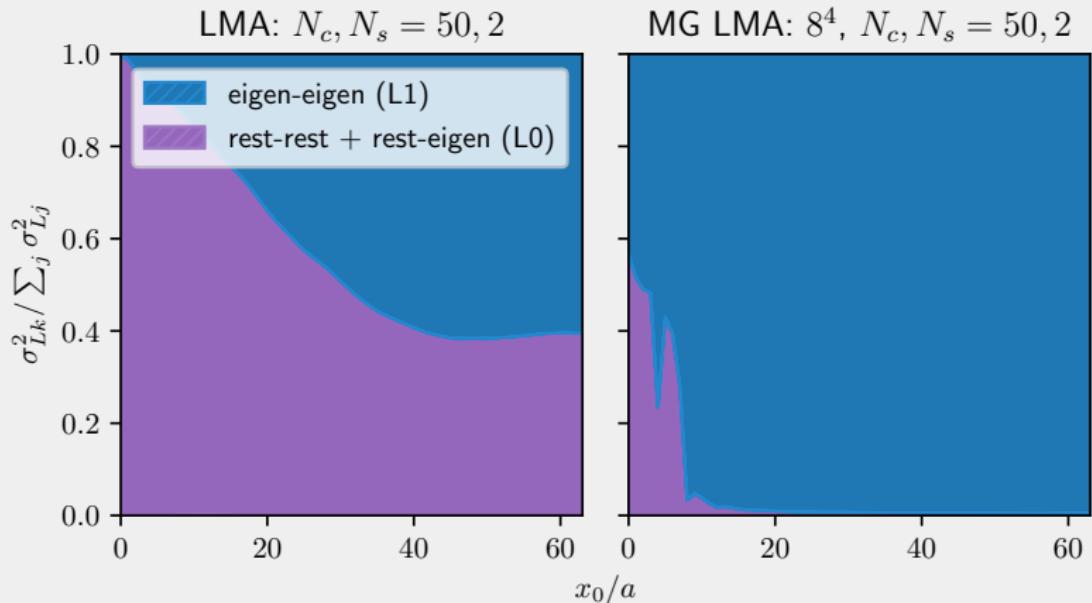
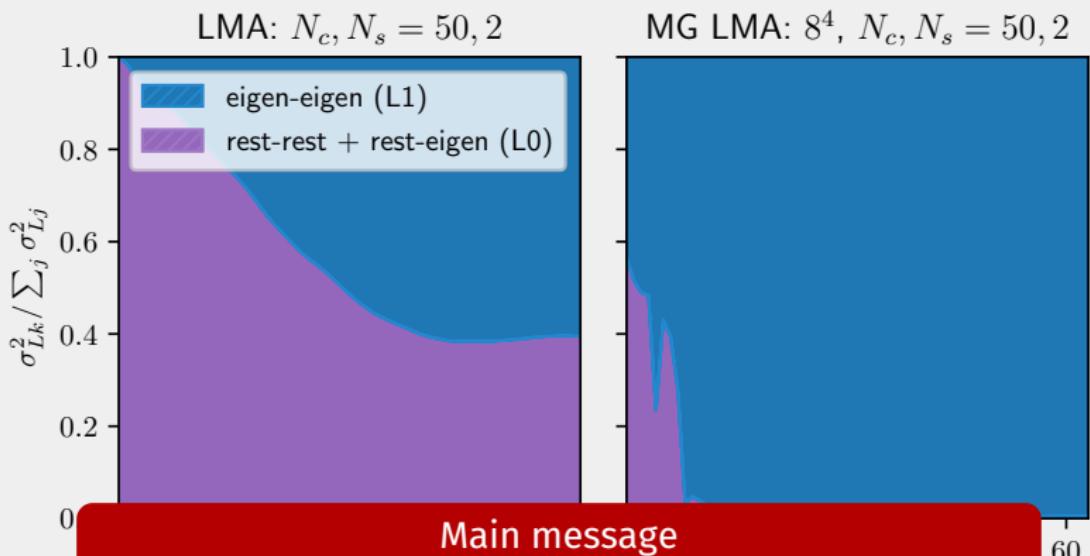


Figure: Relative variance for LMA (left) and MG LMA (right) to the vector correlator with **one stochastic source** for each term.



Main message

Figure: Relative variances of the LMA and MG LMA for the G7 correlation function.

We observe a significant variance contribution from the cheap-to-evaluate L1-term w.r.t LMA.

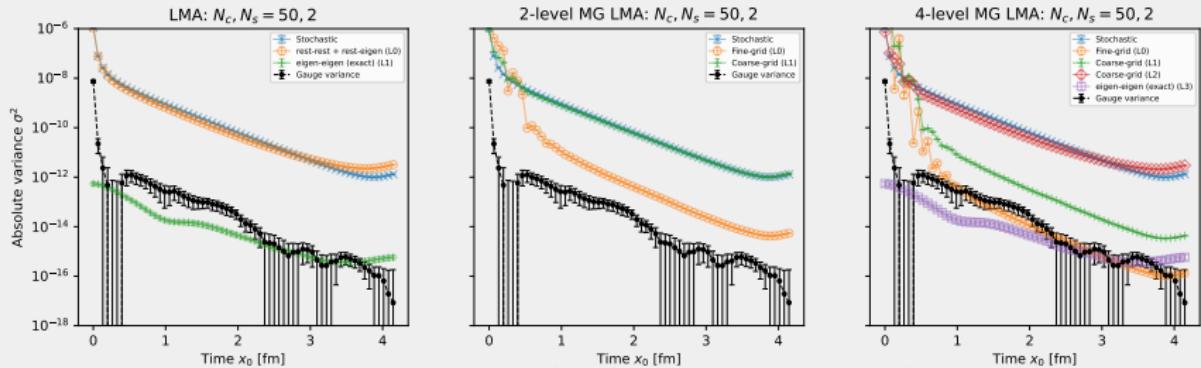


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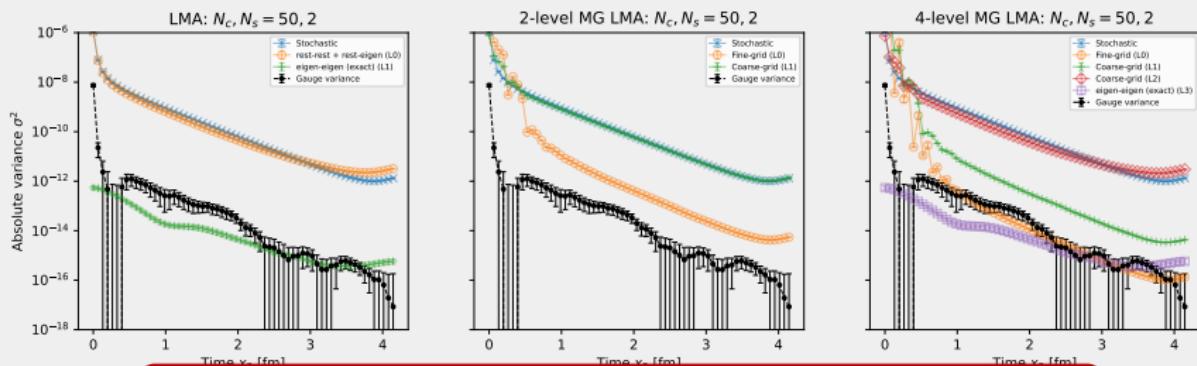


Figure: A
correlat
variance

Main message

We are able to push the remaining Lo noise down to the gauge noise using only a few stochastic sources.

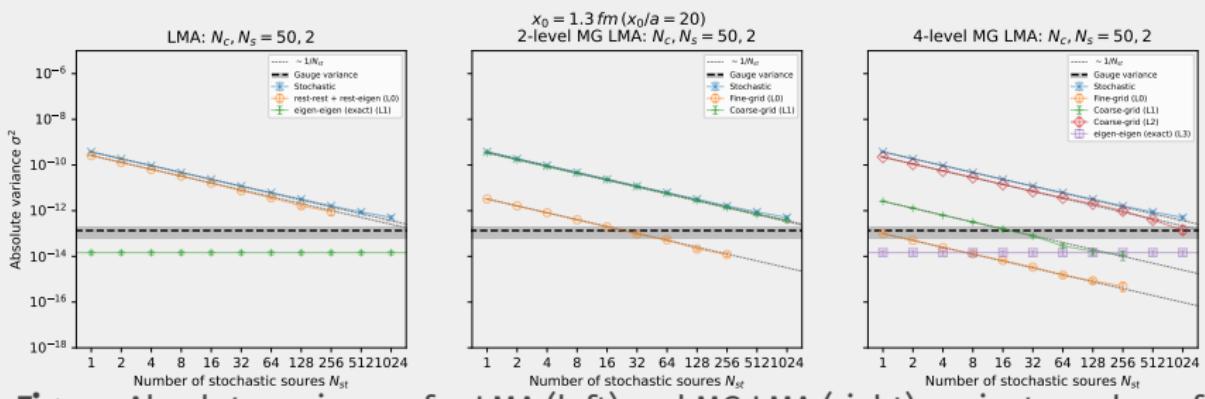


Figure: Absolute variances for LMA (left) and MG LMA (right) against number of stochastic sources N_{st} . The black line is the gauge variance.

VARIANCE VS. VOLUME

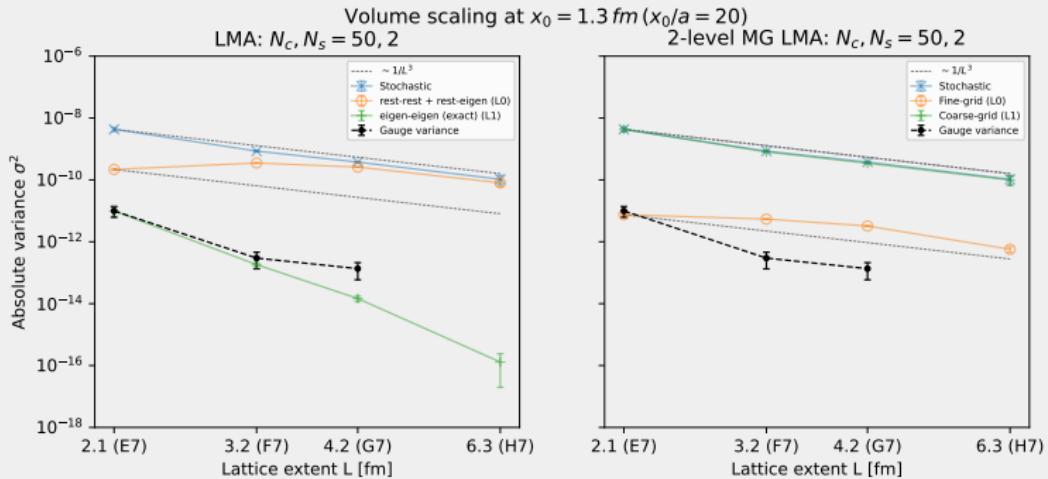


Figure: Absolute variances for LMA (left) and MG LMA (right) against the lattice extent L . The black line is the gauge variance.

VARIANCE VS. VOLUME

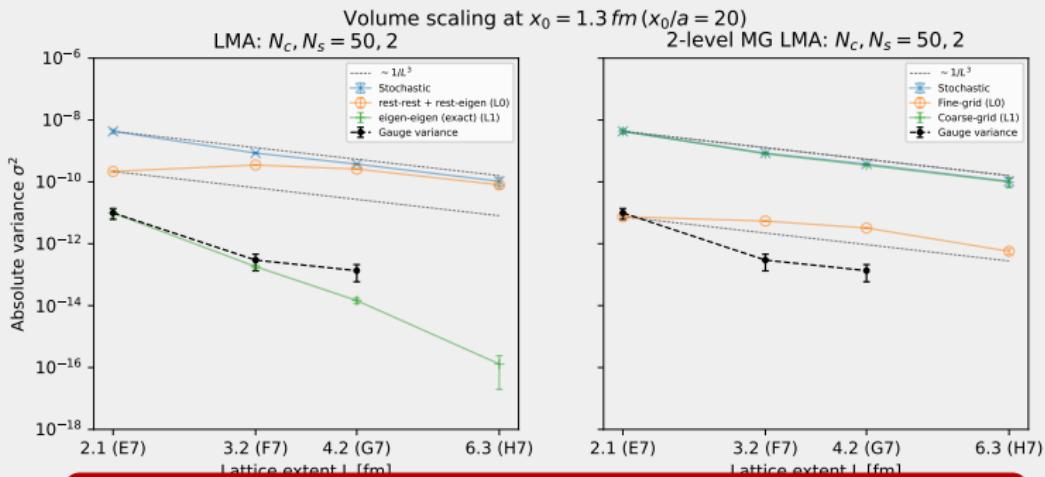


Figure: A
extents L

Main message

MG LMA with a constant number of low modes scales well with the volume.

lattice

COST

Table: Cost breakdown to reach the gauge variance for G7 (4.2 fm).

Estimator	# modes	# sources	meas. cost ¹	model cost ¹
Stochastic	0	Lo: 4096	16384	16384
LMA ²	50	Lo: 2048	8192	8192
2-lvl MG LMA ²	50	Lo: 16 [*] L1: 2048***	557.8	80.7
4-lvl MG LMA ²	50	Lo: 1 [*] L1: 16 ^{**} L2: 1024***	466.7	14.4

My 🐍 implementation:

* fine-grid	128×64^3	inv: 11.1 ± 0.4 sec	(iter: 46.53 ± 0.23)
** coarse-grid	32×16^3	inv: 37.3 ± 2.4 sec	(iter: 1417 ± 22)
*** coarse-grid	16×8^3	inv: 0.667 ± 0.041 sec	(iter: 502.1 ± 5.8)

¹Unit = fine-grid inversions.²Cost of determination of low modes not included (or add 100 - 200 to the cost).

CONCLUSION



CONCLUDING SUMMARY

- Subspaces based on **domain-decomposed / coarsened** low modes
- **Correlator decomposition** into-MG levels
- Method can be defined recursively
- **Every level-contribution** → separate statistics
- 50 low modes capture all the variance (**independent of the lattice volume!**)
- Fewer low modes & more variance contribution than LMA

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Key idea

Hierarchical evaluation: noisy part is cheaper to evaluate!



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BACKUP SLIDES: TABLE OF CONTENTS

- 7 Detailed setups
- 8 Variance contribution - All ensembles
- 9 Absolute variance - All ensembles
- 10 Variance vs. sources - All ensembles
- 11 Cost - All ensembles



DETAILED SETUPS



BACKUP SLIDE: DETAILED SETUPS

Estimator	# modes	Sources	Levels
Stochastic	N/A	semwall	Lo: only fine-grid
LMA	50	semwall exact	Lo: (rest-rest + rest-eigen) L1: (eigen-eigen)
2-level MG LMA	50	semwall	Lo: fine-grid L1: block size 8^4
3-level MG LMA	50	semwall exact	Lo: fine-grid L1: block size 8^4 L2: (eigen-eigen)
4-level MG LMA	50	semwall exact	Lo: fine-grid L1: block size 4^4 L2: block size 8^4 L3: (eigen-eigen)

VARIANCE CONTRIBUTION - ALL ENSEMBLES

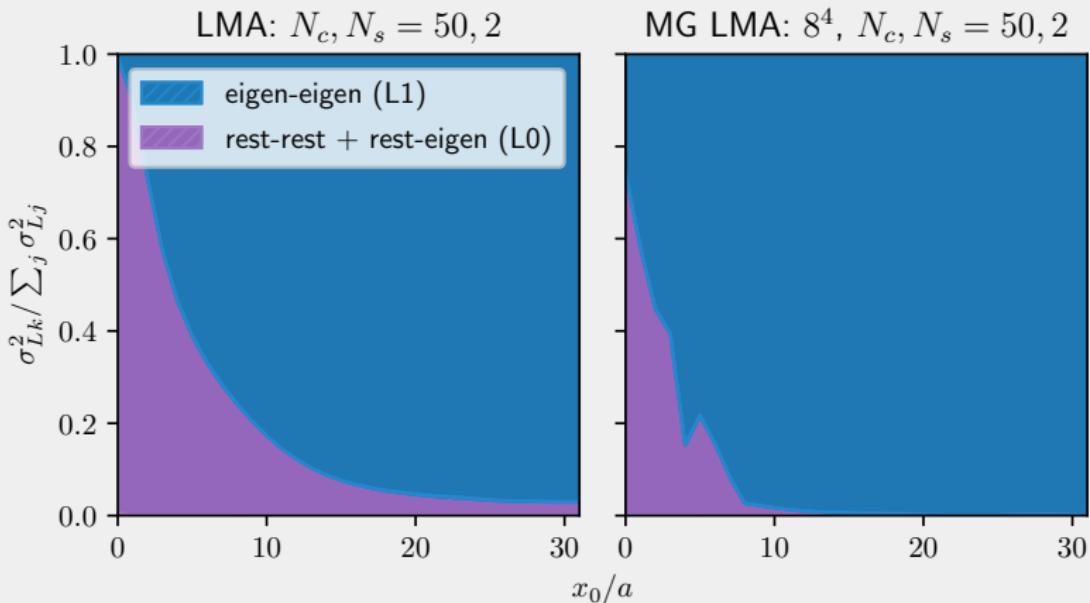
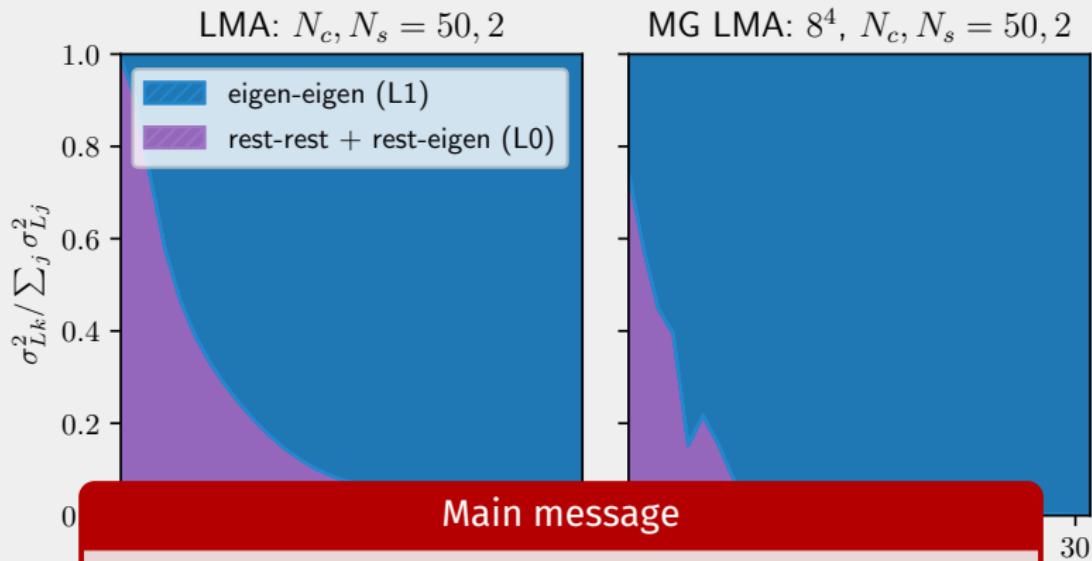


Figure: Relative variance for LMA (left) and MG LMA (right) to the vector correlator with **one stochastic source** for each term.



Main message

We observe a significant variance contribution from the cheap-to-evaluate L1-term w.r.t LMA.

Figure: P

correlate with one stochastic source for each term

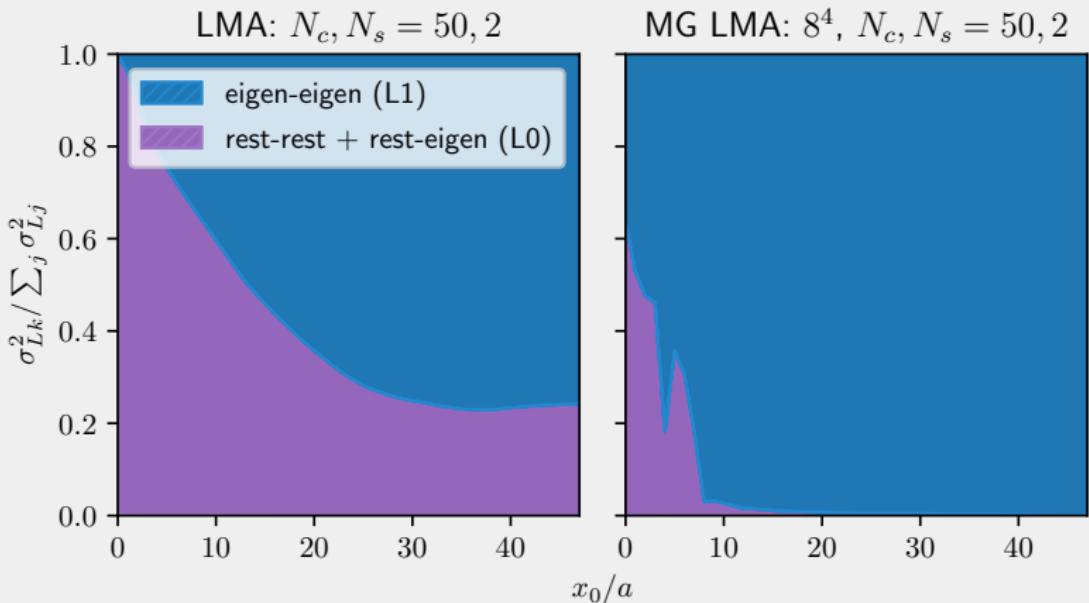
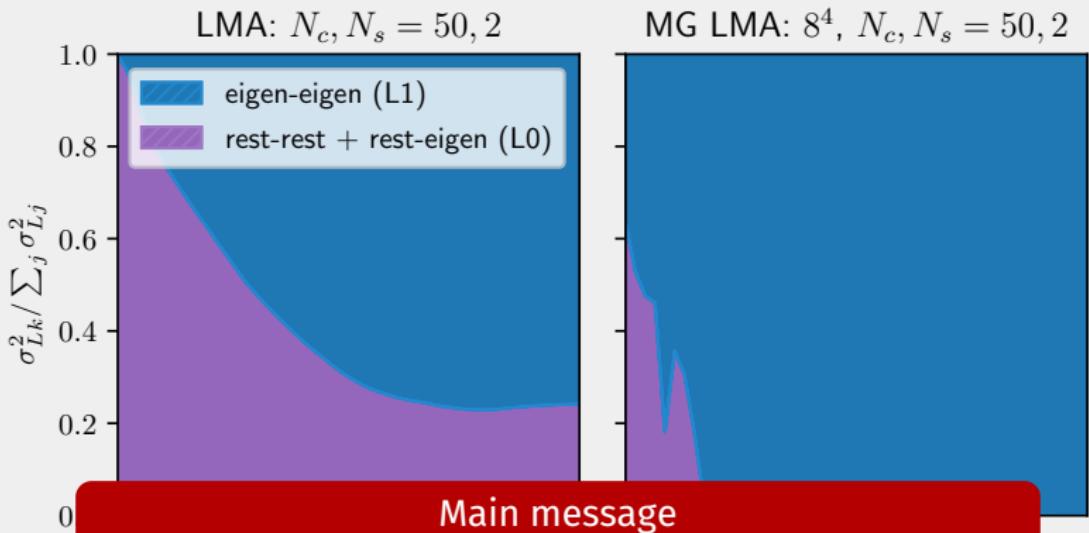


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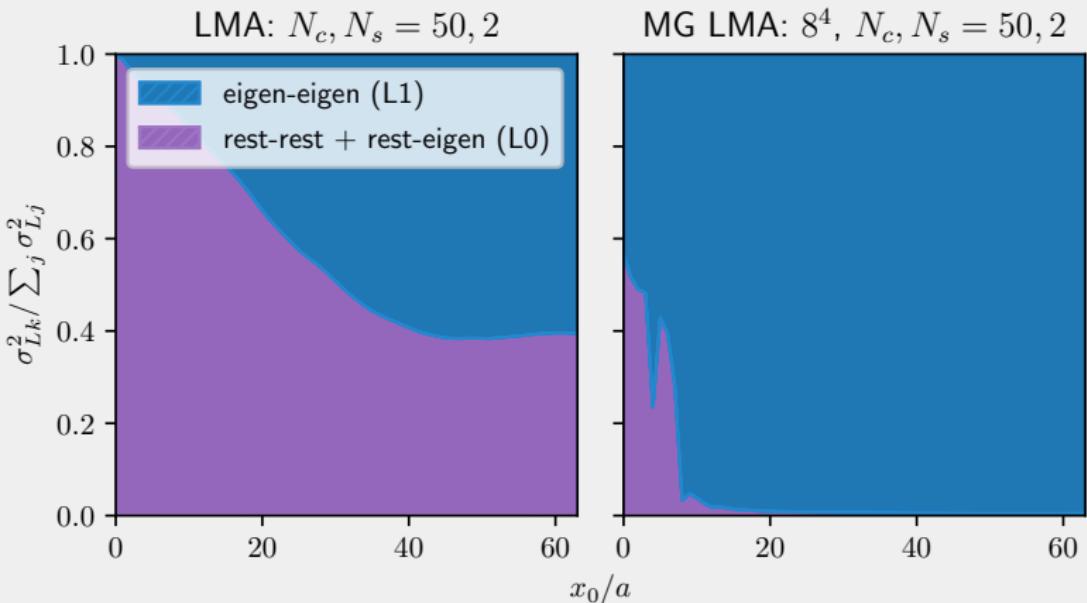
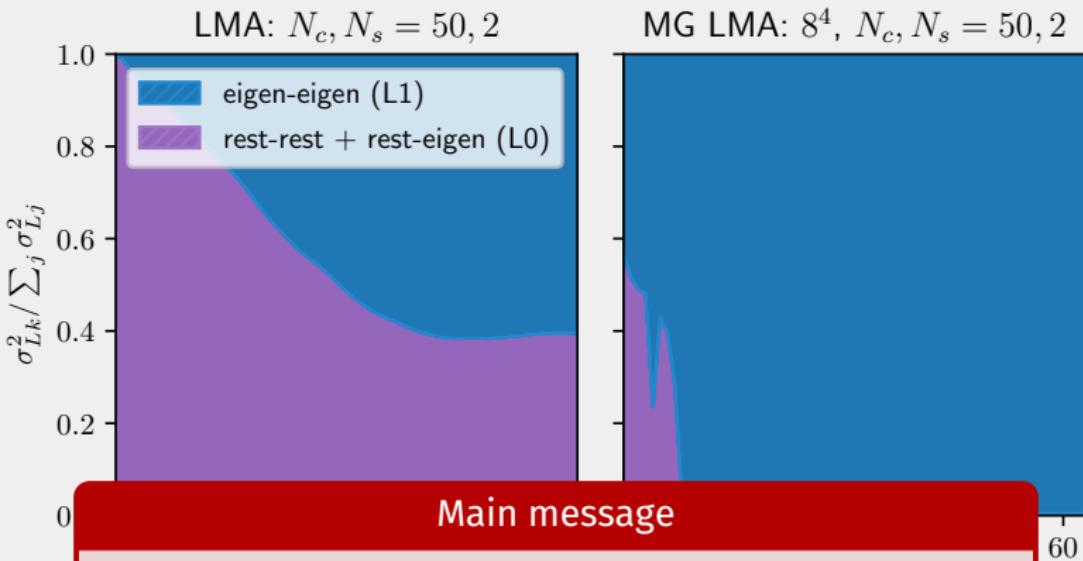


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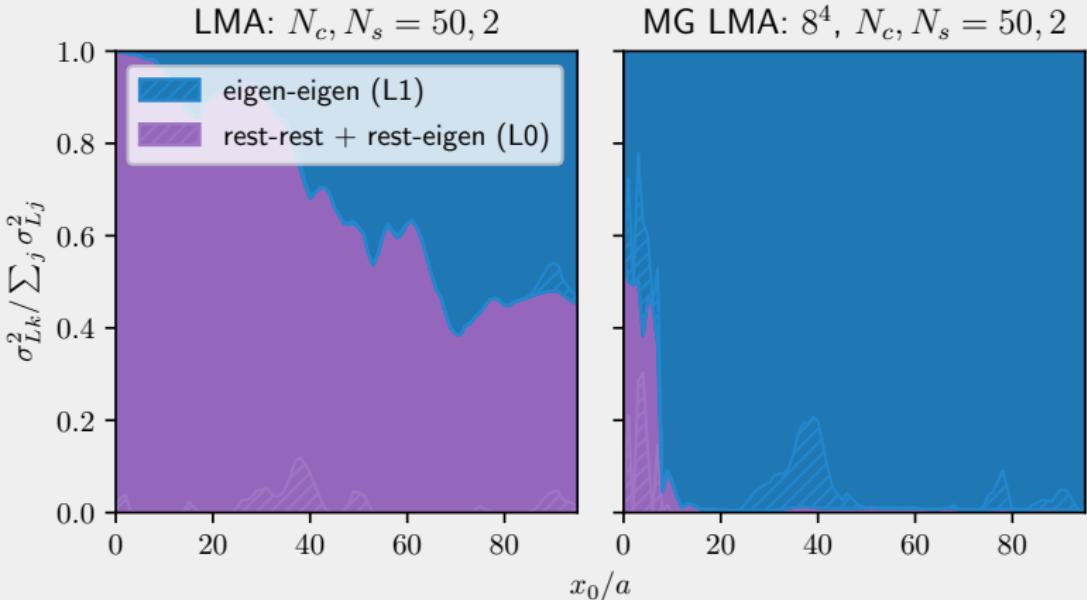
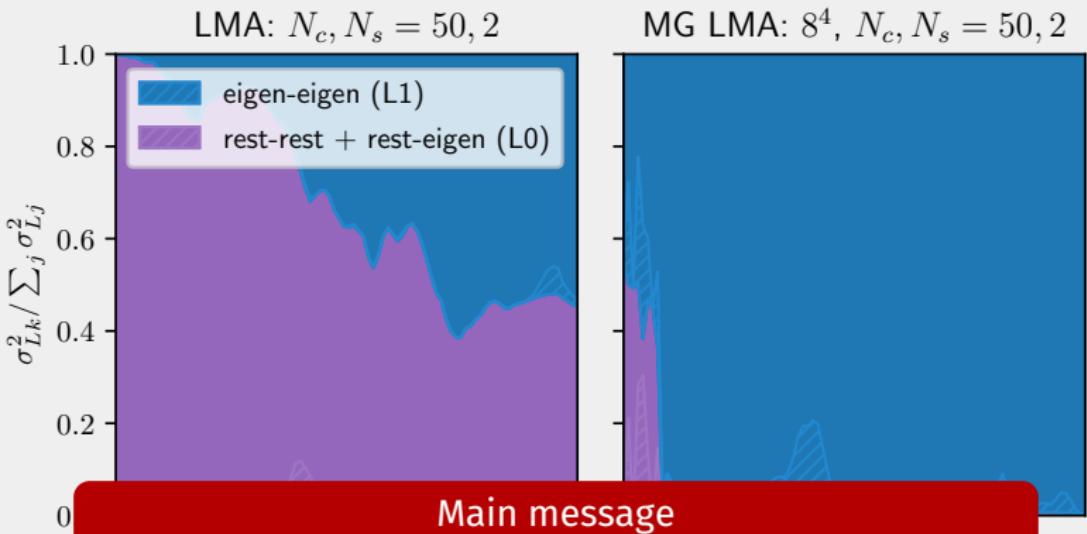


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ABSOLUTE VARIANCE - ALLENSEMBLES

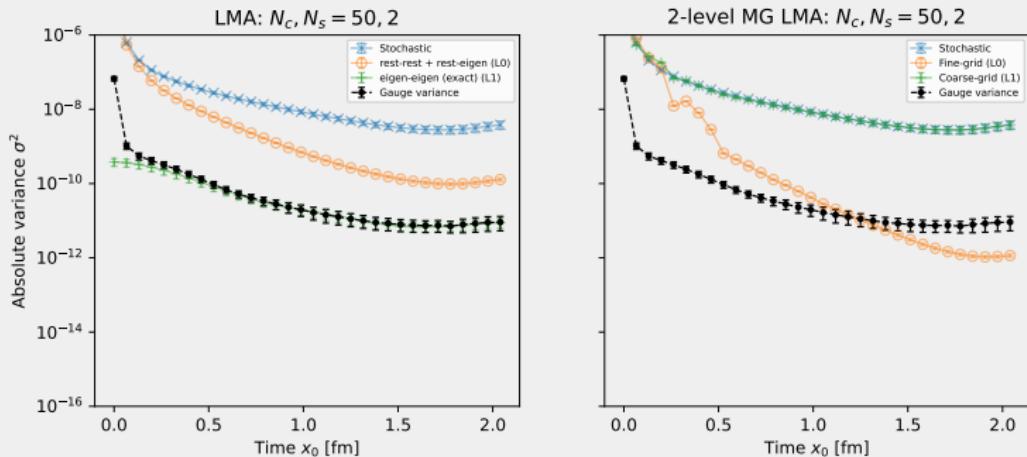
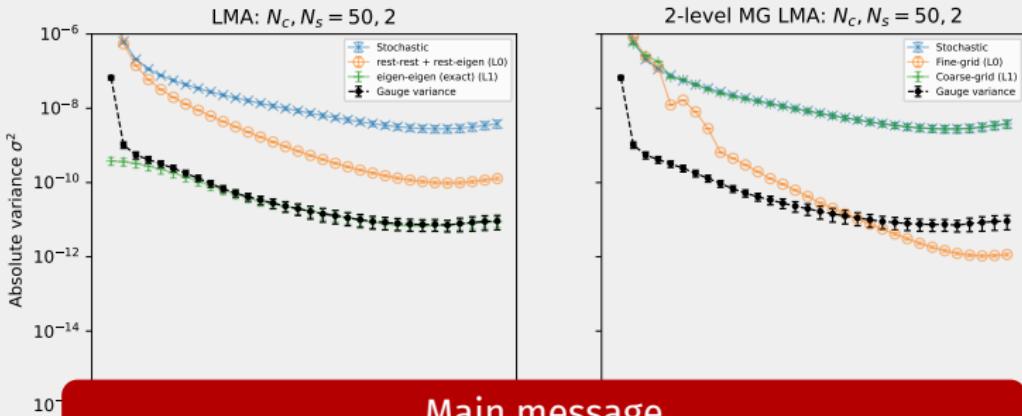


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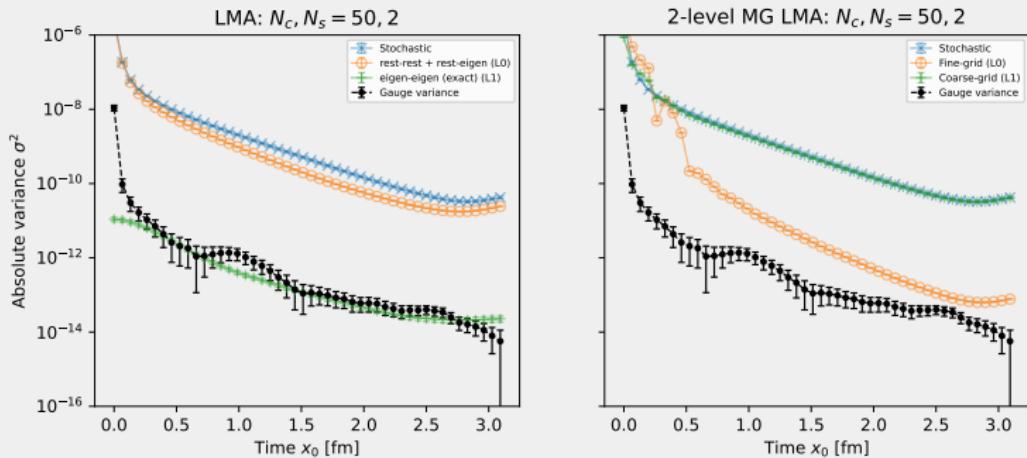
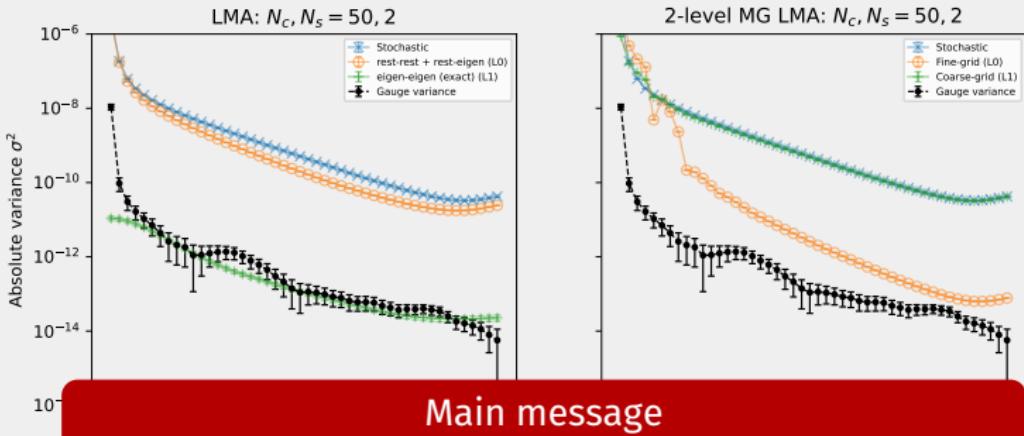


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We are able to push the remaining Lo noise down to the gauge noise using only a few stochastic sources.

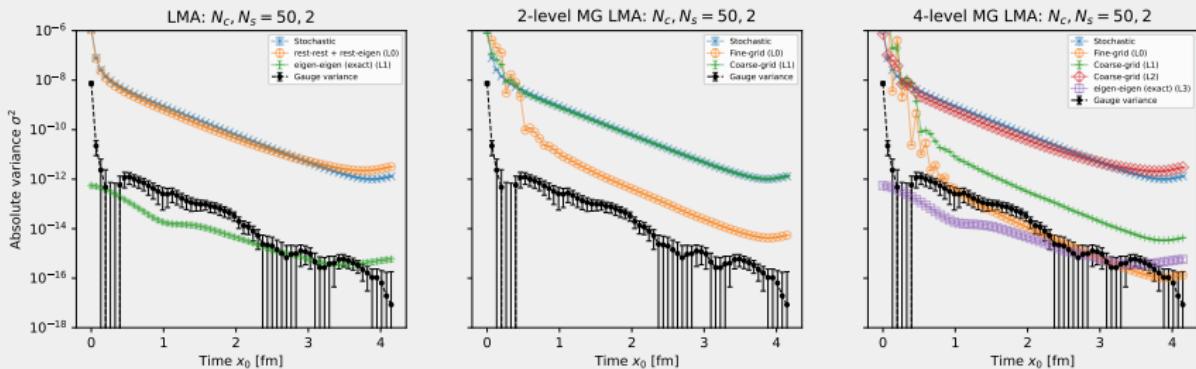


Figure: Absolute variances for LMA (left) and MG LMA (right) to the vector correlator with **one stochastic source** for each term. The black line is the gauge variance.

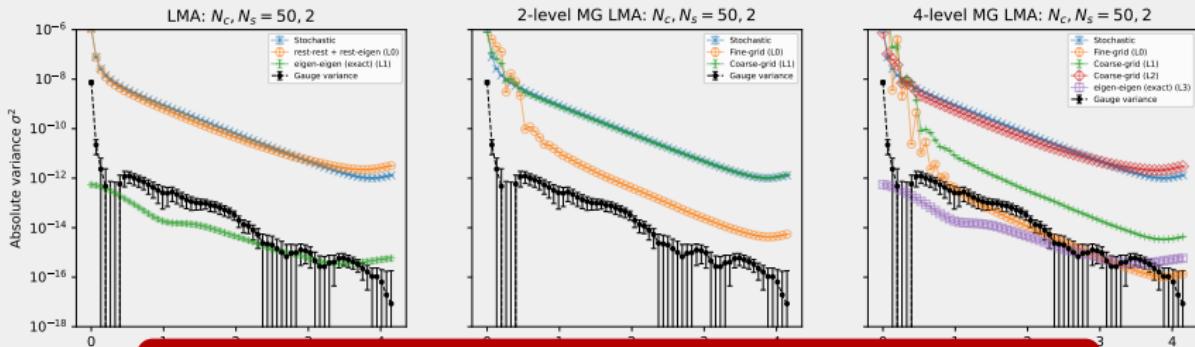


Figure: A comparison of the absolute variance of the correlation function for the gauge variance and the gauge

Main message

We are able to push the remaining Lo noise down to the gauge noise using only a few stochastic sources.

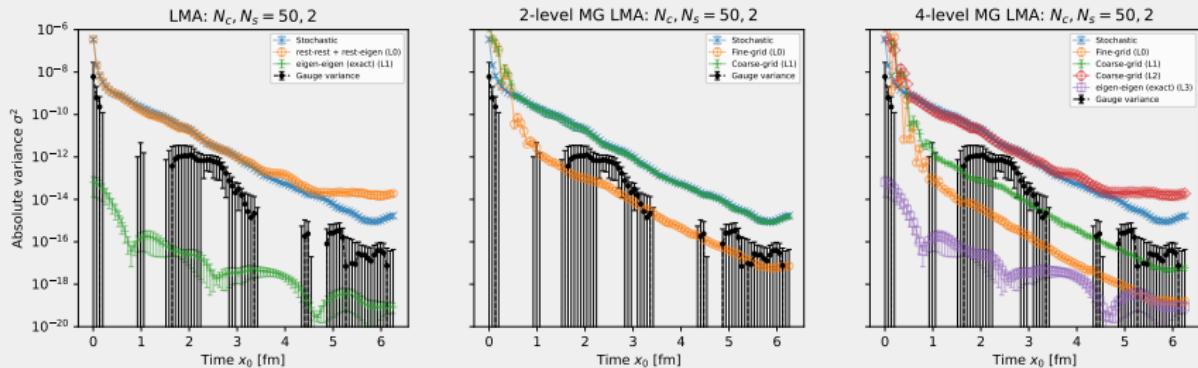


Figure: Absolute variances for LMA (left) and MG LMA (right) to the vector correlator with **one stochastic source** for each term. The black line is the gauge variance.

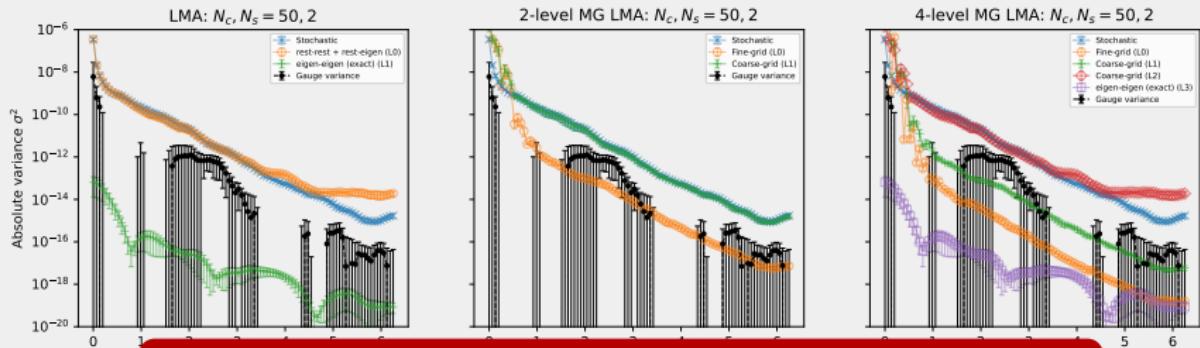


Figure: A
correlat
variance

Main message

We are able to push the remaining Lo noise down to the gauge noise using only a few stochastic sources.



VARIANCE VS. SOURCES - ALL ENSEMBLES

VARIANCE VS. SOURCES: E7

2.1

3.2

4.2

6.2

fm

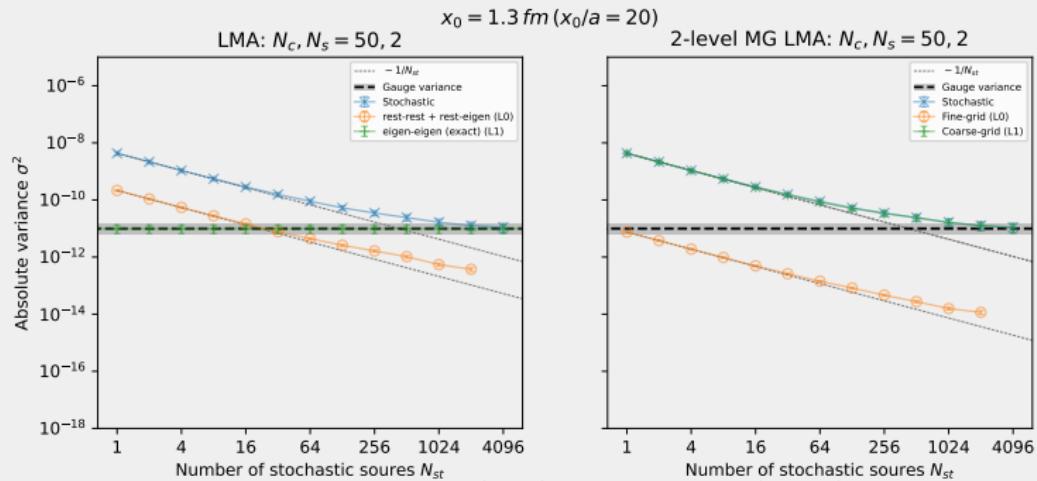


Figure: Absolute variances for LMA (left) and MG LMA (right) against number of stochastic sources N_{st} . The black line is the gauge variance.

VARIANCE VS. SOURCES: F7

2.1

3.2

4.2

6.2

fm

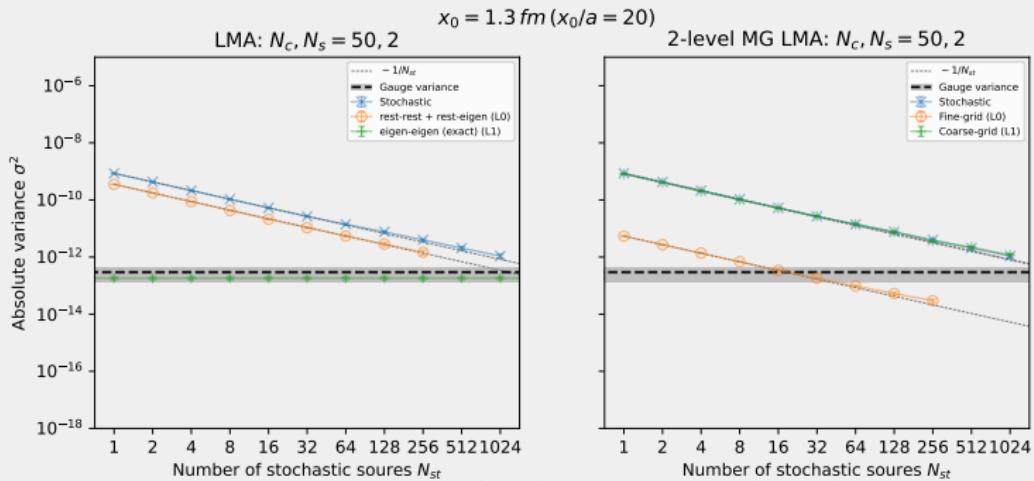


Figure: Absolute variances for LMA (left) and MG LMA (right) against number of stochastic sources N_{st} . The black line is the gauge variance.

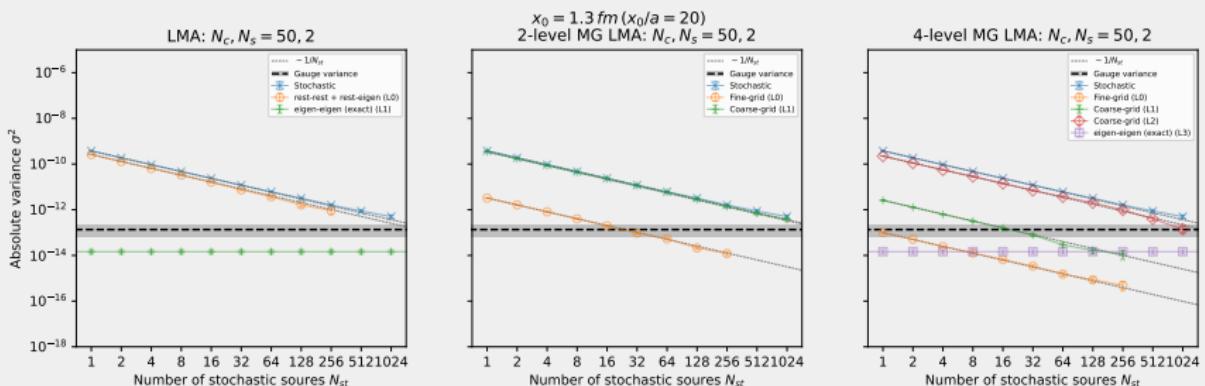


Figure: Absolute variances for LMA (left) and MG LMA (right) against number of stochastic sources N_{st} . The black line is the gauge variance.

VARIANCE VS. SOURCES: H7

2.1

3.2

4.2

6.2

fm

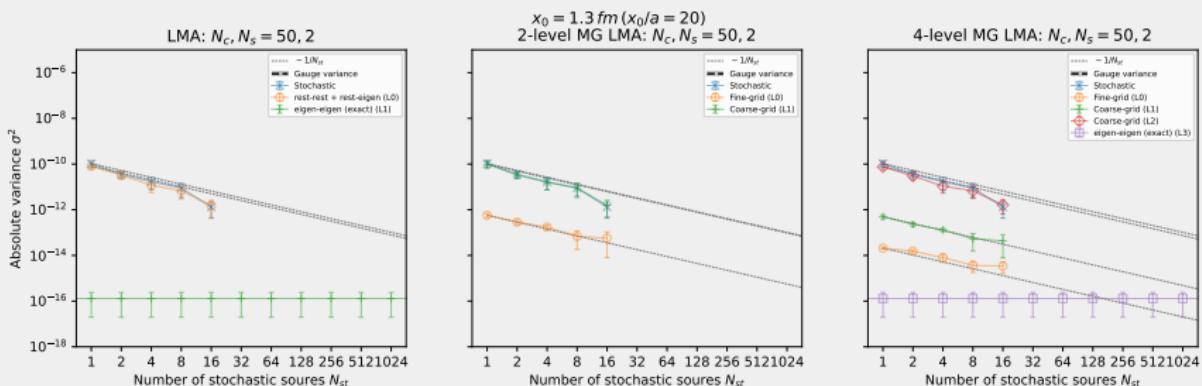


Figure: Absolute variances for LMA (left) and MG LMA (right) against number of stochastic sources N_{st} . The black line is the gauge variance.



COST - ALL ENSEMBLES

Table: Cost breakdown to reach the gauge variance for **E7 (2.1 fm)**.

Estimator	# modes	# sources	meas. cost ¹	model cost ¹
Stochastic	0	Lo: 1024	4096	4096
LMA ²	50	Lo: 16	64	64
2-lvl MG LMA ²	50	Lo: 1* L1: 1024**	100.4	12.3
3-lvl MG LMA ²	50	Lo: 1* L1: 16**	5.5	4.1

My 🐍 implementation:

- * fine-grid 64×32^3 inv: 5.32 ± 0.03 sec (iter: 35.65 ± 0.15)
- ** coarse-grid 8×4^3 inv: 0.125 ± 0.000 sec (iter: 140.5 ± 0.3)

¹Unit = fine-grid inversions.

²Cost of determination of low modes not included (or add 100 - 200 to the cost).

Table: Cost breakdown to reach the gauge variance for F_7 (3.2 fm).

Estimator	# modes	# sources	meas. cost ¹	model cost ¹
Stochastic	0	Lo: 2048	8192	8192
LMA ²	50	Lo: 1024	4096	4096
2-lvl MG LMA ²	50	Lo: 16* L1: 2048**	462.3	80.7
3-lvl MG LMA ²	50	Lo: 16* L1: 1024**	263.2	72.3

My 🐍 implementation:

- * **fine-grid** 96×48^3 inv: 8.42 ± 0.04 sec (iter: 43.77 ± 0.15)
- ** **coarse-grid** 12×6^3 inv: 0.409 ± 0.002 sec (iter: 337.6 ± 1.3)

¹Unit = fine-grid inversions.²Cost of determination of low modes not included (or add 100 - 200 to the cost).

Table: Cost breakdown to reach the gauge variance for **G7 (4.2 fm)**.

Estimator	# modes	# sources	meas. cost ¹	model cost ¹
Stochastic	0	Lo: 4096	16384	16384
LMA ²	50	Lo: 2048	8192	8192
2-lvl MG LMA ²	50	Lo: 16 [*] L1: 2048***	557.8	80.7
4-lvl MG LMA ²	50	Lo: 1 [*] L1: 16 ^{**} L2: 1024***	466.7	14.4

My 🐍 implementation:

* fine-grid	128×64^3	inv: 11.1 ± 0.4 sec	(iter: 46.53 ± 0.23)
** coarse-grid	32×16^3	inv: 37.3 ± 2.4 sec	(iter: 1417 ± 22)
*** coarse-grid	16×8^3	inv: 0.667 ± 0.041 sec	(iter: 502.1 ± 5.8)

¹Unit = fine-grid inversions.

2Cost of determination of low modes not included (or add 100 - 200 to the cost).