Connecting theories of flavour to Kaon physics

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Outline:

- 1. General Introduction to the Flavour Problem
- 2. Effective Field Theories for BSM
- 3. Explicit BSM examples

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Hierarchy problem dark matter/dark energy flavour hierarchies neutrino masses gravity

The (two) flavour problems

- 1. **The SM flavour problem**: The measured Yukawa pattern doesn't seem accidental
	- \Rightarrow Is there any deeper reason for that?

- 2. **The NP flavour problem**: If we regard the SM as an EFT valid below a certain energy cutoff Λ , why don't we see any deviations in flavour changing processes?
	- \Rightarrow Which is the flavour structure of BSM physics?

$$
\mathcal{L}_{\rm Yukawa} \supset Y_u^{ij} \bar{Q}_L^i Hu_R^j
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Exact $U(2)^n$ limit

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An approximate $U(2)^n$ is acting on the light families!

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- **What happens when we switch on NP?**

no breaking of the $U(2)^n$ flavour symmetry at low energies

What's the problem for BSM?

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How to satisfy all the constraints at the same time?

Effective Field Theories for BSM

The EFT approach

- Since we haven't observed any clear sign of NP yet at low energies, we can work in an EFT context
	- \Rightarrow Agnostic of the nature of new physics, describe more than one UV model with the same operators
	- \Rightarrow Try to derive model-independent bounds
- We use the SMFFT
	- \Rightarrow Build all possible operators with SM fields and respecting SM symmetries
- The remnant of high-energy new physics is contained in the Wilson **Coefficients**
	- \Rightarrow With flavour, we have a lot of free degrees of freedom
	- \Rightarrow We need a criterium to infer their magnitude

The $U(2)^n$ symmetry for BSM

$$
q_{3L} \sim (\mathbf{1}, \mathbf{1}) \qquad \qquad \ell_{3L} \sim (\mathbf{1}, \mathbf{1})
$$

$$
Q_L = (Q_L^1, Q_L^2) \sim (\mathbf{\bar{2}}, \mathbf{1}) \qquad \qquad L_L = (\ell_L^1, \ell_L^2) \sim (\mathbf{1}, \mathbf{\bar{2}})
$$

Unbroken $U(2)^5$

The $U(2)^n$ symmetry for BSM

$$
q_{3L} \sim (1, 1) \qquad \qquad \ell_{3L} \sim (1, 1) Q_L = (Q_L^1, Q_L^2) \sim (\bar{2}, 1) \qquad \qquad L_L = (\ell_L^1, \ell_L^2) \sim (1, \bar{2}) V_q \sim (2, 1) \qquad \qquad V_\ell \sim (1, 2)
$$

Unbroken
$$
U(2)^5
$$

\n
$$
Y_u = y_t \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix}
$$
\n
$$
\overline{q}_{3L} \Gamma q_{3L} \checkmark
$$
\n
$$
\overline{q}_{3L} \Gamma Q \checkmark
$$

Soft symmetry breaking

$$
Y_u = y_t \begin{pmatrix} \Delta & V_q \\ 0 & 1 \end{pmatrix}
$$

 $\bar{q}_{3L}\Gamma q_{3L}$ \checkmark

$$
\bar{q}_{3L}\Gamma(V_qQ)\checkmark
$$

Which operators?

$$
\mathcal{L}_{\text{eff}}=-\frac{1}{\Lambda^2}(\bar{q}_{3L}\gamma_\mu\sigma^a q_{3L})(\bar{\ell}_{3L}\gamma^\mu\sigma^a\ell_{3L})-\frac{c_{13}}{\Lambda^2}(\bar{q}_{3L}\gamma_\mu q_{3L})(\bar{\ell}_{3L}\gamma^\mu\ell_{3L})
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• Left-handed fields only

 \Rightarrow Only ones contributing to di-neutrino modes without considering right-handed ν

•
$$
q_{3L} \equiv q_L^b + \theta_q e^{i\phi_q} \hat{V}_q^{\dagger} \cdot Q_L
$$

\n $q_L^b = \begin{pmatrix} V_{j3}^* u_L^j \\ b_L \end{pmatrix}$ $Q_L^i = \begin{pmatrix} V_{ji}^* u_L^j \\ d_L^i \end{pmatrix}$ $\hat{V}_q \equiv (V_{td}^*, V_{ts}^*)$

 \Rightarrow θ_q and ϕ_q are small mixing angles

Correlations among all these modes is essential to prove NP scenarios

Observables: R_{D(*)}

$$
R_{D^{(*)}} = \frac{\mathcal{B}(B \to D^{(*)}\tau\bar{\nu})}{\mathcal{B}(B \to D^{(*)}\ell\bar{\nu})}
$$

- Test of Lepton Flavour Universality between the 3rd and light lepton families
- Ratios allow cancelling hadronic uncertainties and experimental uncertainties

$$
\Delta R_{D^{(*)}} = \left[\frac{R_{D^{(*)}}}{R_{D^{(*)}}^{\rm SM}} - 1 \right] \approx 2 R_0 (1 - \theta_q \cos \phi_q) \qquad R_0 = \frac{1}{\Lambda^2} \frac{1}{\sqrt{2} G_F}
$$

Observables: $B^+ \to K^+ \nu \bar{\nu}$ 2311.14647]

[2311.14647]

Hadronic Tagging

$$
\mathcal{B}(B^+ \to K^+ \nu \bar{\nu}) = (1.1^{+0.9+0.8}_{-0.8-0.5}) \times 10^{-5}
$$

Inclusive Tagging

 $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu}) = (2.7 \pm 0.5 \pm 0.5) \times 10^{-5}$

Combined

$$
\mathcal{B}(B^+ \to K^+ \nu \bar{\nu}) = (2.7 \pm 0.5^{+0.5}_{-0.4}) \times 10^{-5}
$$

 2.7σ tension wrt SM $3.5\,\sigma$ evidence wrt background only

Observables: $B^+ \rightarrow K^+ \nu \bar{\nu}$

- From the theory point of view, the $B^+ \to K^+ \nu \bar \nu$ mode is very clean
	- No charm pollution
	- Pollution from $B \to \tau(\to K\nu)\bar{\nu}$, but removable experimentally
	- Very precise, already hitting the wall with precision

$$
\frac{\mathcal{B}(B \to K^{(*)} \nu \bar{\nu})}{\mathcal{B}(B \to K^{(*)} \nu \bar{\nu})_{\rm SM}} = \left[\frac{2}{3} + \frac{1}{3}\left|1 - \frac{\pi s_w^2 \Delta R_{D^{(*)}}}{2 \alpha X_t (1 - \theta_q \cos \phi_q)} \theta_q e^{i \phi_q} (1 - c_{13})\right|^2\right]
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distinguishes among UV models

Kaon decays

$$
\frac{\mathcal{B}(K^+ \to \pi^+ \nu \bar{\nu})}{\mathcal{B}(K^+ \to \pi^+ \nu \bar{\nu})_{\text{SM}}} = \frac{2}{3} + \left| \frac{1}{3} + \frac{\Delta R_{D^{(*)}} \theta_q^2 (1 - c_{13})}{2(\alpha/\pi) C_{sd,\tau}^{\text{SM,eff}} (1 - \theta_q \cos \phi_q)} \right|^2
$$

$$
\frac{\mathcal{B}(K_L \to \pi^0 \nu \bar{\nu})}{\mathcal{B}(K_L \to \pi^0 \nu \bar{\nu})_{\text{SM}}} = \frac{2}{3} + \left| \frac{1}{3} - \frac{\Delta R_{D^{(*)}} \theta_q^2 (1 - c_{13})}{2(\alpha/\pi)(X_t/s_w^2)(1 - \theta_q \cos \phi_q)} \right|^2
$$

- Same proportionality to $\Delta R_{D(*)}$ as in $B \to K \nu \bar{\nu}$
- Same CKM structure as in the SM
- We have a large interference also for the $K_L \rightarrow \pi^0 \nu \bar{\nu}$ decay

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- Currently, the NA62 limits are not enough to ping down a specific parametric window
- KOTO-II sensitivity at 25% of the measured branching fraction
	- Hypothesis: $\mathcal{B}(K_L \to \pi^0 \nu \bar{\nu})_{\rm exp} = 1.3 \mathcal{B}(K_L \to \pi^0 \nu \bar{\nu})_{\rm SM}$
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• Blue lines: $K^+ \to \pi^+ \nu \bar{\nu}$, red lines: $K_L \to \pi^0 \nu \bar{\nu}$

- $\Delta R_{D(*)}$ fixed to its central value
- filled: $c_{13} = 0 \phi_q = 0$, dashed $c_{13} = 2 \phi_q = \pi$
- "Model" allows for large modifications, NA62 and KOTO-II updates are **essential** to constrain the parameter space 19/28

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Correlation with B**-physics**

If In this setup, the relative shift to fit R_D and R_{D^*} is the same

$$
\Delta R_{D^{(*)}} = \left[\frac{R_{D^{(*)}}}{R_{D^{(*)}}^{\text{SM}}} - 1 \right] = 0.133 \pm 0.036
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- In all scenarios $c_{13} = 0 \phi_q = 0$
- Blue: $\theta_q = 0.6$, red $\theta_q = -1$, green $\theta_a = 0.3$

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KOTO-II is essential to select the parameter space

Fit to di-neutrino modes

• Fit assuming the current central value for $\Delta R_{D(*)}$ and $c_{13} = 0.5$

- \Rightarrow Excellent fit quality
- We predict

$$
\mathcal{B}(K_L \to \pi^0 \nu \bar{\nu}) = (6.7 \pm 1.8) \times 10^{-11}
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testable at 2σ at KOTO-II

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Explicit BSM examples

Which NP particle?

1) Colourless Mediators

 $\bullet \,\, W' + Z'$: tension with high- p_T searches with $\tau_L\tau_L$ or b_Lb_L final states

[Greljo,Isidori,Marzocca,'15]

• Solutions with right-handed neutrino are motivated and help to ease the tension with $b \to c\tau \nu$ data but they are most likely to be excluded from high- p_T [Greljo, Camalich,Ruiz-Álvarez,'18]

2) Leptoquark Mediators

[Angelescu, Bečirević, Faroughy, Sumensari, '18]

Flavour Non-Universal New Physics

The Pati-Salam Leptoquark

 $G_i = SU(4)^i \times SU(2)^i_L \times SU(2)^i_R$

The Pati-Salam Leptoquark

 $G_i = SU(4)^i \times SU(2)^i_L \times SU(2)^i_R$

$G_{12} \times G_3 \Rightarrow SM + U_1 + Z'$

The Pati-Salam Leptoquark

 $G_i = SU(4)^i \times SU(2)^i_L \times SU(2)^i_R$

Large couplings to third generation

+other (heavy) states that depend on the details of the model in the UV

Phenomenology

- The Z' mediates neutral meson mixing at tree-level
	- \Rightarrow Suppressed parametrically
- The U_1 LQ has tree level matching $c_{13} = 1$, no tree-level contribution to $s \to d\nu\bar{\nu}$
- However, there are loop contributions

- The loop doesn't preserve the $U(2)$ CKM structure
- The contribution to $K_L \to \pi \nu \bar{\nu}$ is negligible

The scalar leptoquarks $S_1 + S_3$

- Scalar leptoquarks don't require a UV completion
- They both contribute at tree-level to $s \to d\nu\bar{\nu}$ transitions such as

$$
c_{13}|_{S_1} = \frac{1}{2} \qquad c_{13}|_{S_3} = \frac{1}{2}
$$

• Sweet spot to connect to $B^+ \to K^+ \nu \bar{\nu}$

Conclusions

- Flavour physics has exciting prospects in the search for New Physics
- Understanding the origin of the Flavour Puzzle might hint at what structure new physics couplings have
- Correlations among different observables are crucial to studying the viability of classes of models
	- \Rightarrow Di-neutrino modes are one of the golden modes where to look for new physics signals
	- \Rightarrow The complementarity between measurements $B^+ \to K^+ \nu \bar \nu, K^+ \to \pi^+ \nu \bar \nu$ and $K_L \rightarrow \pi^0 \nu \bar{\nu}$ are essential to ping down specific scenarios

[Appendix](#page-50-0)

Flavour Non-Universal New Physics

Dvali, Shifman, '00 Panico, Pomarol, '16 MB, Cornella, Fuentes-Martin, Isidori '17 Allwicher, Isidori, Thomsen '20 Barbieri, Cornella, Isidori, '21 Davighi, Isidori '21

Basic idea:

- 1st and 2nd have small masses and small couplings to NP because they are generated by dynamics at a heavier scale
- 3rd generation is linked to dynamics at lower scales and has stronger couplings

Flavour deconstruction:

fermion families interact with different gauge groups and flavour hierarchies emerge as accidental symmetries