Majorana phases beyond neutrinoless double beta decay

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Details: A. Dery, S. Gori, Y. Grossman, ZL, JHEP 10 (2024) 100, 2406.18647

The standard model + neutrino mass

- Clearest empirical evidence that the minimal SM is incomplete:
 - Dark matter
 - Baryon asymmetry of the Universe
 - Neutrino mass
 - Inflation in the early universe [have a plausible theoretical picture]
 - Dark energy [cosmological constant? need to know more?]

• We do not even know the Lagrangian that describes the particles we have observed

$$\mathcal{L} = -Y_e^{ij} \overline{L_{Li}^I} \phi e_{Rj}^I - \begin{cases} \frac{Y_{\nu}^{ij}}{\Lambda} L_{Li}^I L_{Lj}^I \phi \phi & \text{violates lepton number} \\ Y_{\nu}^{ij} \overline{L_{Li}^I} \tilde{\phi} \nu_{Rj}^I & \text{requires } \nu_R \text{ fields} \end{cases}$$

Do neutrino mass terms violate lepton number? 10 or 12 parameters in lepton sector?



The mixing of neutrinos

• Fermions with same quantum numbers mix, Yukawas define mass eigenstates:

$$M = \begin{pmatrix} M_{a1} & M_{a2} & M_{a3} \\ M_{b1} & M_{b2} & M_{b3} \\ M_{c1} & M_{c2} & M_{c3} \end{pmatrix} = \begin{pmatrix} 1 \\ c_{23} & s_{23} \\ -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & s_{13}e^{-i\delta} \\ 1 \\ -s_{13}e^{i\delta} & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} \\ -s_{12} & c_{12} \\ 1 \end{pmatrix} \underbrace{\begin{pmatrix} e^{i\eta_1} \\ e^{i\eta_2} \\ 1 \end{pmatrix}}_{\text{if Majorana, 1 if Dirac}}$$

• The additional phases $\eta_{1,2}$ do not affect oscillation experiments, only lepton # violation

• It is often said that we should measure all parameters in the Lagrangian... although...

(Wolfenstein: 'I do not care what the values of the Wolfenstein parameters are, so you should not either; the only question is if their independent determinations give consistent results')

• However, it is $\eta_{1,2}$ and not δ , which is the least known parameter of the PMNS matrix Can we ever hope to measure a different linear combination than what enters $0\nu\beta\beta$?



Neutrino oscillation measurements

Multi-GeV μ-like

cosθ

0.5

150

100

- Three mixing angles have been measured
- Oscillation between two flavors $(\delta m^2 = m_1^2 m_2^2)$

$$P_{\rm osc} = \sin^2(2\theta) \, \sin^2\left(1.27 \, \frac{\delta m^2}{\rm eV^2} \, \frac{L}{\rm km} \, \frac{\rm GeV}{E}\right)$$

- Atmospheric neutrinos: $1 \sim (10^{-3}) \times (10^{1...4}) / (10^{0\pm 1})$ half of up-going ν_{μ} get lost
- Solar neutrinos: $\delta m^2 L/E \gg 1$
- Two mass-squared differences have been measured, but not the absolute mass scale (Short baseline anomalies not easy to fit, even with 4 flavors)







Neutrinos — many unknowns

Are neutrinos = their own antiparticles?

(Different than all other known particles? Theoretically favored, most leptogenesis models)

• What is the absolute mass scale?

Two mass-squared differences measured At least one state $m_{\nu_i}\gtrsim 50\,{\rm meV}$

Cosmology: $\sum m_i < 0.072 \text{ eV}$ [DESI 2024] (n.b.: preference for $\sum m_i < 0$ [e.g., 2407.07878])

- Value of *CP* violating phase δ ?
- Is the mass ordering "normal" (NO) or "inverted" (IO)?

If IO, neutrinoless double beta decay $(0\nu\beta\beta)$ experiments will decide If NO, may or may not see $0\nu\beta\beta$, even in Majorana case









Neutrinos — a history of surprises

- Most theorists' expectations around early 1990's:
 - Solar neutrino problem will go away, we do not understand the Sun
 - If it does not, solution must be small angle MSW, since it's cute
 - Expect $\Delta m^2_{23} \sim 10-100 {\rm eV}^2$, since it's cosmologically interesting (DM)
 - Expect $\theta_{23} \sim V_{cb} \simeq 0.04$, motivated by simple GUT models
 - Atmospheric neutrino anomaly will go away, because it requires large mixing angle — the first that became compelling (⇒ Nobel, 2002)
 - 2000s: tribimaximal mixing ansatz, predicted θ_{13} near zero $\theta_{13} \sim 9^{\circ}$, not too small helps CP violation searches [inspired by H. Murayama]

Experiments crucial, independent of prevailing theoretical "guidance"





Wrong

Wrong

Wrong

Wrong

Wrong

Wrong

Quark vs. lepton mixing

• Mixing matrix parameters, assuming 3-generation unitarity:

U_{PMNS} :	$\sin \theta_{12} = 0.550 \pm 0.011$	$\sin \theta_{13} = 0.148 \pm 0.002$	
	$\sin \theta_{23} = 0.749 \pm 0.010$	$\delta = (177^{+19}_{-20})^\circ~~$ [unconstrained at $_{3\sigma}$]	[$ u$ fit 2024, NO, converted]
$V_{ m CKM}:$	$\sin \theta_{12} = 0.2250 \pm 0.0007$	$\sin \theta_{13} = 0.0037 \pm 0.0001$	
	$\sin\theta_{23} = 0.0418 \pm 0.0008$	$\delta = (65.7 \pm 1.5)^{\circ}$	[PDG 2024]

- Are the origin of quark and lepton masses and mixings related?
- Some lepton processes are especially clean; quark sector much more rich
- Neutrino FCNCs seem impossible to search for; e.g., $\nu_i \rightarrow \nu_j \gamma$, $X \rightarrow \nu_i \bar{\nu}_j(Y)$
- SM flavor puzzle extended: why lepton and quark masses and mixings so different?



Quark vs. lepton mixing (2)







Why don't we know yet if Majorana or Dirac?

- In the $m_{\nu} \rightarrow 0$ limit, the distinction between a Dirac and Majorana fermions disappears Except for oscillation experiments, no consequence of $m_{\nu} \neq 0$ has ever been observed
- Any experiment to probe the nature of neutrino mass involve suppression by $\propto m_{\nu}/E$ FCNC neutrino decays allowed, e.g., $\nu_3 \rightarrow \nu_1 \gamma$, $\nu_3 \rightarrow \nu_2 \nu_1 \bar{\nu}_1$ — rates extremely small
- The smoking gun signature would be the observation of lepton number violation (Majorana neutrinos are their own antiparticles, thus cannot carry any quantum number)
- Neutrinos we can study are always ultrarelativistic

Exception: cosmic neutrino background ($T_{\nu} \sim 2 \text{ K} \sim 2 \times 10^{-4} \text{ eV}$) (In Majorana case, both ν and $\bar{\nu}$ interact with detector)





Experimental challenges detecting LNV

• The rates are suppressed by m_{ν} , need huge statistics Best chance is $0\nu\beta\beta$, rate $\propto m_{ee}^2$ $m_{ee} = \left|\sum_{i=1}^3 m_i U_{ei}^2\right|$

Planned experiments will reach $m_{ee} \sim 0.01 \, \text{eV}$, decisive for IO For NO, m_{ee} can vanish even if neutrinos are Majorana

• Second best: $\mu^- \to e^+$ conversion, search for $pp\mu^- \to nne^+$ Proportional to $m_{\mu e} = \left| \sum_{i=1}^3 m_i U_{ei} U_{\mu i} \right|$

Mu2e and COMET will improve current bound to $\sim\!10^{-16}$

• Expectation from PMNS much smaller; patterns would give powerful constraints





0.100



Invariants and conventions

• Mixing matrix: $U = \begin{pmatrix} 1 \\ c_{23} & s_{23} \\ -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & s_{13}e^{-i\delta} \\ 1 \\ -s_{13}e^{i\delta} & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} \\ -s_{12} & c_{12} \\ 1 \end{pmatrix} \begin{pmatrix} e^{i\eta_1} \\ e^{i\eta_2} \\ 1 \end{pmatrix}$

These quantities are not physical / phase convention independent E.g., the Majorana phases η_1 and η_2 can be shifted to 2nd at 3rd (or 1st and 3rd) entries

- Define: Dirac phase = CP violating phase measurable in LN conserving processes
 Majorana phases = only accessible through LNV processes
- For CKM, it is well known that 4 elements are needed to define a physical CPV quantity: $t_{\alpha i\beta j} = U_{\alpha i}U_{\beta j}U^*_{\alpha j}U^*_{\beta i}$

For Majorana fermions, fewer phases can be absorbed in field redefinitions, hence $s_{\alpha ij} = U_{\alpha i} U^*_{\alpha j}$ are physical





Our choice of parameters

• For Dirac fermions, 3 mixing angles and a phase:

$$\{|t_{e2e3}|, |t_{e3e3}|, |t_{\mu2e3}|, |\Psi_D\}, \qquad \Psi_D = \arg(t_{\mu2e3}) = \arg(c_{12}c_{23}e^{-i\delta} - s_{12}s_{23}s_{13})$$

 $(-|t_{\mu 2e3}| \sin \Psi_D = J)$, the leptonic Jarlskog invariant, the same way as for CKM)

• For Majorana fermions, 2 additional phases:

$$\{\Phi_{12}, \Phi_{23}\} \equiv \{\Phi_{12}^e, \Phi_{23}^e\} \qquad \Phi_{ij}^{\alpha} = \arg(s_{\alpha ij})$$

E.g., $\Phi_{12} = \eta_1 - \eta_2$, etc.







• If interested in other LNV processes besides $0\nu\beta\beta$, it is natural to generalize:

$$m_{\alpha\beta} = \left| \sum_{i=1}^{3} m_i U_{\alpha i} U_{\beta i} \right| \qquad \alpha, \beta \in \{e, \mu, \tau\}$$

Explicit expressions are not too illuminating, simplest is:

$$m_{ee}^{2} = (m_{1}^{2}c_{12}^{4} + m_{2}^{2}s_{12}^{4})c_{13}^{4} + m_{3}^{2}s_{13}^{4} + 2m_{1}m_{2}s_{12}^{2}c_{12}^{2}c_{13}^{4}\cos\left[2(\eta_{1} - \eta_{2})\right] + 2m_{1}m_{3}c_{12}^{2}s_{13}^{2}c_{13}^{2}\cos\left[2(\eta_{1} + \delta)\right] + 2m_{2}m_{3}s_{12}^{2}s_{13}^{2}c_{13}^{2}\cos\left[2(\eta_{2} + \delta)\right]$$

- The rate of a lepton number violating process is proportional to the corresponding $m^2_{\alpha\beta}$

For example, the $\mu^- \rightarrow e^+$ conversion rate $\propto m_{\mu e}^2$

Sensitivity to each Majorana phase scales as the product of corresponding masses, $m_i m_j \Rightarrow$ large suppression



 $\mu^{-}(N,Z) \rightarrow e^{+}(N,Z-2)$ (in suitable nuclei)





Experimental bounds, in a nutshell

Experimental bounds:



- Experimental sensitivity to $0\nu\beta\beta$ searches are by far the best, since macroscopic amounts of nuclei are used, rather than particle beams
- Mu2e & COMET will improve the bound on $m_{\mu e}$ by ~ 4 orders of magnitude Tiny rates: $\sim 3 \times 10^{-22} (m_{\mu e}^2/m_e^2) |M|^2$





The single phase limit

- We define $\{\nu_{\ell}, \nu_2, \nu_o\}$ so that ν_{ℓ} is always the lightest state
- In the $m_{\nu_{\ell}} \rightarrow 0$ limit, the Majorana phase dependence of all $m_{\alpha\beta}$ is the same (instead of $\eta_{1,2}$ only one η)
- Proof: a symmetric rank-2 neutrino Yukawa matrix has 5 real and 5 imaginary parameters (for rank-3: 6 real & 6 imag.)

The charged lepton Yukawas contain 9 + 9 parameters

The global $U(3)_L \times U(3)_E$ symmetry is completely broken, allowing to remove 6 + 12 parameters, leaving 5 masses, 3 mixing angles, and 2 phases (one Dirac and one Majorana) as physical parameters



vertical lines show where single phase limit becomes a good approximation





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Majorana phase from m_{ee} in single phase limit



Insets show that as m_l gets small, measurement of m_{ee} determine Φ_{2o}

• If the lightest m_{ν} gets small, the correlation between m_{ee} and all other $m_{\alpha\beta}$ get stronger





Predicting $m_{\mu e}$ from m_{ee}





rrrr

A no-loose "theorem"

- With current PMNS and Δm^2 central values, m_{ee} and $m_{\mu e}$ cannot simultaneously vanish (even in case of NO)
 - \Rightarrow In principle, LNV is detectable with better $m_{\mu e}$ sensitivity, even for NO
- Tantalizing PMNS values: $m_{ee} + m_{\mu e}$ cannot vanish (barely, at the 2σ level)



[Dery, Gori, Grossman, ZL, 2406.18647]

• Very challenging (impossible?) to reach such sensitivity Neutron stars contain $\sim 10^{55} \mu^{-}$, can it be used somehow?

[Dery, Gori, Grossman, ZL, in progress]





Similar story for $m_{ au i}$



• Experimental prospects to probe these may be even more remote







- Whether neutrino mass terms violate lepton number is the most significant open question about particles we have seen — experimental discovery would be transformational
- 0νββ experiments are sensitive to one linear combination of the two Majorana phases
 (i) may discover LN violation; (ii) may rule it out (IO); (iii) may not find out (NO)
- Can we ever determine both Majorana phases? Seems very challenging
- Single phase limit: as the lightest neutrino mass gets smaller, all $m_{\alpha_{\beta}}$ depend on a single combination of the two Majorana phases
- How far can sensitivity to $m_{\mu e}$ be pushed?
- Nonrelativistic neutrinos? Room for new ideas!





