2. Innovations in Neutrino Detection Technologies

Exploring Neutrino Properties with Enhanced Quantum-Field Measurement Techniques

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1. Introduction to Neutrino Physics and Quantum Measurement

6. Quantum Technologies and Their Impact on Neutrino Physics

Liquid Argon Detectors (LArTPC)

The Deep Underground Neutrino Experiment (**DUNE**) [6] uses Liquid Argon Time Projection Chambers (**LArTPCs**) [10] for studying neutrino oscillations with high precision. These detectors offer:

- Fine spatial and temporal resolution.
- Crucial for observing neutrino interactions in long-baseline experiments

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- α , β represent neutrino flavors,
- $|U_{\alpha i}|^2$ and $|U_{\beta i}|^2$ are the elements of the PMNS matrix,
- $\Delta m_{ij}^2 = m_i^2 m_i^2$ is the mass squared difference between neutrino states,
- L is the distance traveled,
- E is the energy of the neutrino.

Coherent Elastic Neutrino-Nucleus Scattering (CEvNS)

Recent measurements in CEvNS [6], such as those performed at Oak Ridge National Laboratory (ORNL), have confirmed Standard Model predictions and constrained new types of neutrino interactions .

Super-Kamiokande Upgrades

Entangled neutrinos can be described as: $| \psi \rangle = \alpha | \nu_e \nu_\mu \rangle + \beta | \nu_\mu \nu_e \rangle$

Where α and β are complex coefficients characterizing the entanglement between neutrino pairs.

Incorporating gadolinium into the Super-Kamiokande [1] [2] detector has significantly improved sensitivity, allowing for more precise measurements of low-energy neutrino signals .

Technological Challenges:

Noise reduction techniques are being implemented, including advanced shielding to protect detectors from environmental interference.

5.1 Superconducting Quantum Interference Devices (SQUIDs)

SQUIDs are ultra-sensitive devices capable of detecting the minute **magnetic fields** produced by neutrino interactions. These sensors are instrumental in achieving the **high precision** required for modern neutrino experiments.

5.2 Time Projection Chambers (TPCs) and Quantum-Enhanced Detection

The combination of **Time Projection Chambers** (TPCs) with **quantum-enhanced readout** technologies enables high-precision tracking of neutrino interactions, improving spatial and temporal resolution.

5. Experimental Setups for Neutrino Measurement

6.1 Quantum Entanglement in Neutrino Detection

Quantum entanglement can be harnessed to improve the signal-to-noise ratio. By **correlating measurements of neutrinos**, even when separated by large distances, this method increases measurement accuracy while reducing the amount of data required.

Quantum interferometry allows for the detection of tiny phase shifts induced by neutrino interactions, with the phase shift related to the neutrino's mass. **The phase shift in an interferometric setup** is described by: $\Delta \phi = \frac{2\pi m^2 L}{E} \,.$

6.2 Application in Future Neutrino Detectors

Neutrino oscillations occur when neutrinos, initially in flavor eigenstates, $(\nu_e, \nu_\mu, \nu_\tau)$ evolve into mass eigenstates. The flavor transition probability is described by:

Advances in quantum technologies will facilitate the development of the next generation of neutrino detectors capable of measuring CP violation, mass ordering, and neutrino backgrounds with unprecedented precision.

$$
P(\nu_\alpha \rightarrow \nu_\beta) = \sum_{i=1}^3 |U_{\alpha i}|^2 |U_{\beta i}|^2 \left(1 - \cos \left(\frac{\Delta m_{ij}^2 L}{2E} \right) \right)
$$

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Quantum interferometry promises to revolutionize neutrino physics. By allowing for ultra-precise measurements of neutrino oscillations, it opens up new opportunities to study tiny fluctuations in neutrino energy and mass. This precision could provide insight into previously inaccessible areas of neutrino physics.

7. Quantum Interferometry and Precision Neutrino Measurements

8. Applications in Cosmology and Astrophysics

Quantum measurement techniques are transforming our approach to studying neutrinos. Technologies like quantum sensors, squeezing, and entanglement are improving measurement precision, bringing us closer to resolving key mysteries in particle physics and cosmology. As these technologies advance, they will enable new breakthroughs in both fundamental physics and our understanding of the universe, unlocking exciting possibilities for future neutrino research.

9. Conclusion

Neutrinos play a significant role in understanding cosmological events such as Big Bang nucleosynthesis and supernova explosions. By applying advanced quantum measurement techniques, we can probe the early universe and refine our cosmological models, providing better insight into the formation of the universe's structures

4.1 Quantum Sensing and Signal-to-Noise Ratio Improvement

Quantum squeezing enhances the precision of measurements by reducing uncertainty in one observable. For neutrino detection, this technique helps improve the signal-to-noise ratio, allowing for better energy resolution and spatial accuracy.

The squeezed state uncertainty is given by:

$$
\Delta X_{\text{squeezed}}^2 = \frac{1}{2} \left(\Delta X_{\text{vacuum}}^2 - 2 |\langle \alpha \rangle|^2 \right)
$$

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Where $\langle \alpha \rangle$ is the squeezing amplitude, optimizing the detection of neutrino interactions.

4.2 Quantum Interferometry for Neutrino Detection

Where *m* is the neutrino mass, *L* is the travel distance, and *E* is the neutrino energy. Quantum interferometry enables highly **precise measurements of neutrino properties** in real-time.

4. Quantum Measurement Enhancements in Neutrino Detection

3.1 Quantum Mechanics of Neutrino Oscillations

Where:

This equation allows for precise determination of the **neutrino mass hierarchy**, **mixing angles**, and **CP-violation phase**.

3.2 Experimental Evidence for Neutrino Oscillations

3. Theoretical Framework of Neutrino Oscillations

1.1 Neutrinos: The Elusive Fundamental Particles

Neutrinos are nearly massless particles interacting via the weak nuclear force, making them extremely difficult to detect. They are produced in various astrophysical processes, including supernova explosions, cosmic ray interactions, and nuclear fusion in stars. Despite their elusive nature, neutrinos are critical to understanding phenomena like stellar nucleosynthesis, the solar neutrino problem, and cosmological evolution.

1.2 Quantum-Field Measurement Techniques

Quantum-field measurement techniques, such as **SQUIDs** and optical interferometers, exploit quantum phenomena like coherence and squeezing to enhance neutrino detection precision. These methods have been transformative in experiments like DUNE [8] and T2K [4], advancing our understanding of neutrino properties

> **Figure 3: Unitarity Relations of Neutrino Oscillations** *[9] The flavor transition probabilities PαβP_{\alpha \beta}Pαβ are plotted for an energy range of 0.5 GeV to 10 GeV and baseline length of 1300 km, with the decay parameter* $\gamma=0.1\gamma$ *gamma = 0.1γ*=0.1 *and matter density* $\rho = 3 g/cm3$ *\rho* = 3 *\, \text{g/cm}^3ρ=3g/cm3.*

Figure 1: *The diagram illustrates the signal formation in a Liquid Argon Time Projection Chamber (LArTPC) with three wire planes*

Figure 4: *Schematic layout of the TES/RF-SQUID/resonator system developed by MIT and Lincoln Laboratories [11], illustrating the integration of superconducting technologies for detecting faint signals.*

Figure 2: *Three-flavor neutrino oscillation probabilities* [7] *as evaluated using a quantum computer (squares), a quantum computer simulator (circles), and theoretical predictions (lines). Results highlight agreement across methods, illustrating the potential of quantum computing in neutrino physics*.

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