

# Searches for Nucleon Decay and Baryon Number Violation with Super-Kamiokande Experiment

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*for the Super-Kamiokande Collaboration*

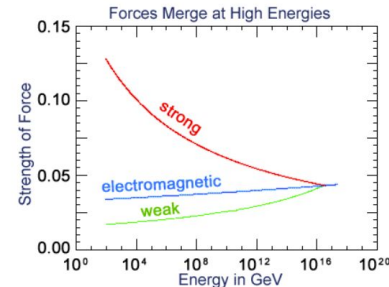


# Do protons decay?

- Lightest baryon (proton) stable
- **B-number** proposed to explain matter stability [Weyl, 1929; Stueckelberg, 1939; Wigner, 1949]
- In Standard Model (SM) **B** not fundamental, accidental global symmetry  
... long history of such symmetries later found violated (e.g.  $C$ ,  $P$ )
- Already in SM **B** is violated by non-perturbative effects [t'Hooft, 1976]
- SM incomplete, **B** viol. in many theories (baryogenesis, GUTs, SUSY, extra-dim...)
- Global symmetries expected to be violated by quantum gravity e.g. [Harlow, Ooguri, 2018]
- Proton decay can be essential for fate of astrophysical objects [Adams, Laughlin, 1997]

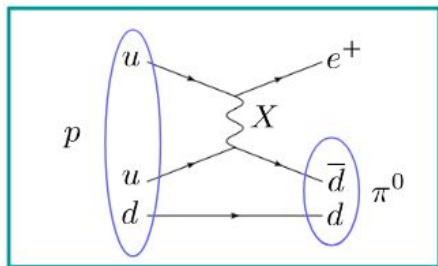
# Grand Unified Theories (GUTs)

- Unify SM gauge groups [Georgi, Glashow, 1974; Fritzsch, Minkowski, 1975]  
→ charge quantization, coupling unification, anomaly cancellation...
- **Leptons** ↔ **quarks** interact → **nucleon decay**



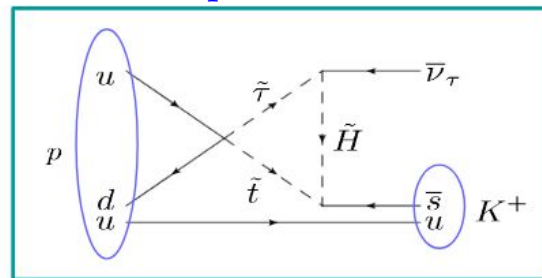
typically dominant:  $p \rightarrow e^+ \pi^0$   $SU(5)$

$$\tau = \frac{1}{\Gamma} \propto \left[ \frac{M_X^2}{\alpha^2} \right]^2$$



→ prediction  $\tau \sim 10^{29-36}$  yrs  
→ minimal model ruled out  
(IMB-3, Kamiokande, Super-K)

$p \rightarrow \bar{\nu} K^+$

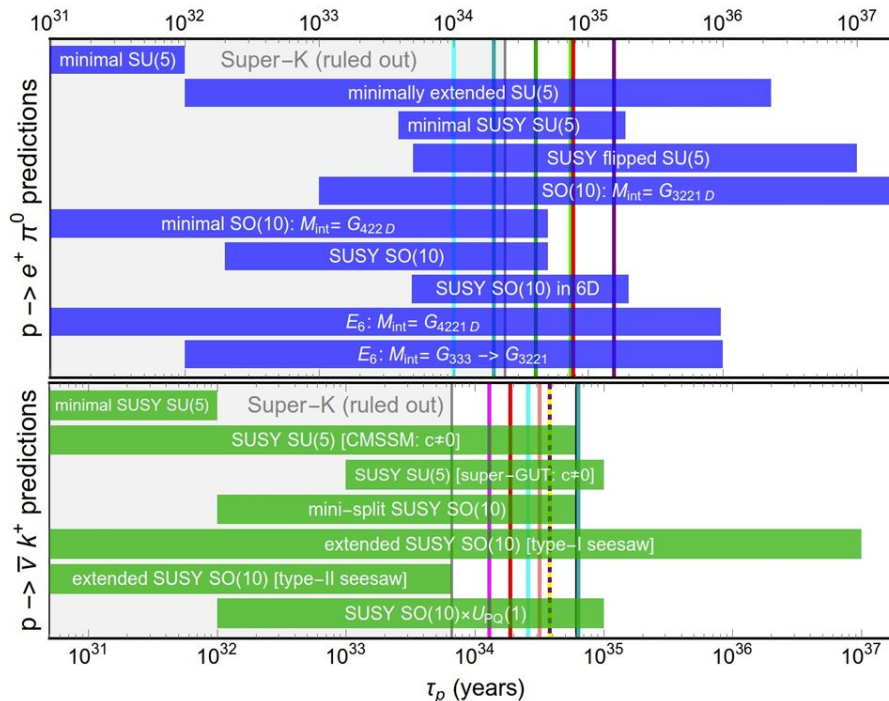


→ prediction  $\tau \sim 10^{29-36}$  yrs  
→ minimal (TeV-)SUSY model ruled out  
by Super-K [Kobayashi+ (SK), PRD, (2005)]

**SUSY**  
 $SU(5)$

$$\tau = \frac{1}{\Gamma} \propto \left[ \frac{M_s M_T}{\alpha^2} \right]^2$$

# Proton Lifetime Predictions



Model	Ref.	Modes	$\tau_N$ (years)
Minimal $SU(5)$	Georgi, Glashow [2]	$p \rightarrow e^+ \pi^0$	$10^{30} - 10^{31}$
Minimal SUSY $SU(5)$	Dimopoulos, Georgi [11], Sakai [12] Lifetime Calculations: Hisano, Murayama, Yanagida [13]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$	$10^{28} - 10^{32}$
SUGRA $SU(5)$	Nath, Arnowitt [14, 15]	$p \rightarrow \bar{\nu} K^+$	$10^{32} - 10^{34}$
SUSY $SO(10)$ with anomalous flavor $U(1)$	Shafi, Tavartkiladze [16]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$ $p \rightarrow \mu^+ K^0$	$10^{32} - 10^{35}$
SUSY $SO(10)$ MSSM (std. $d = 5$ )	Lucas, Raby [17], Pati [18]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$	$10^{33} - 10^{34}$ $10^{32} - 10^{33}$
SUSY $SO(10)$ ESSM (std. $d = 5$ )	Pati [18]	$p \rightarrow \bar{\nu} K^+$	$10^{33} - 10^{34}$ $\lesssim 10^{35}$
SUSY $SO(10)/G(224)$ MSSM or ESSM (new $d = 5$ )	Babu, Pati, Wilczek [19, 20, 21], Pati [18]	$p \rightarrow \bar{\nu} K^+$ $p \rightarrow \mu^+ K^0$	$\lesssim 2 \cdot 10^{34}$ $B \sim (1 - 50)\%$
SUSY $SU(5)$ or $SO(10)$ MSSM ( $d = 6$ )	Pati [18]	$p \rightarrow e^+ \pi^0$	$\sim 10^{34.9 \pm 1}$
Flipped $SU(5)$ in CMSSM	Ellis, Nanopoulos and Wlaker[22]	$p \rightarrow e/\mu^+ \pi^0$	$10^{35} - 10^{36}$
Split $SU(5)$ SUSY	Arkani-Hamed, <i>et. al.</i> [23]	$p \rightarrow e^+ \pi^0$	$10^{35} - 10^{37}$
$SU(5)$ in 5 dimensions	Hebecker, March-Russell[24]	$p \rightarrow \mu^+ K^0$ $p \rightarrow e^+ \pi^0$	$10^{34} - 10^{35}$
$SU(5)$ in 5 dimensions option II	Alciati <i>et.al.</i> [25]	$p \rightarrow \bar{\nu} K^+$	$10^{36} - 10^{37}$
GUT-like models from Type IIA string with D6-branes	Klebanov, Witten[26]	$p \rightarrow e^+ \pi^0$	$\sim 10^{36}$

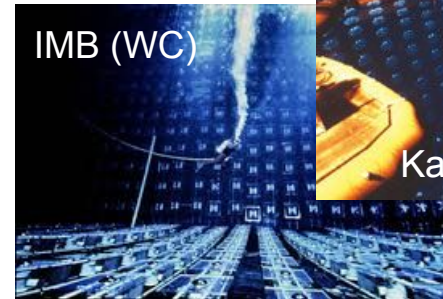
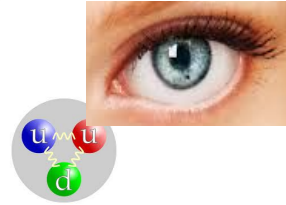
Snowmass-2021 Whitepaper [Dev et.al., *J.Phys.G*, 2024]

[Bueno+, *JHEP*, 2007]

- Significant uncertainties in model predictions, many viable models within experiment reach

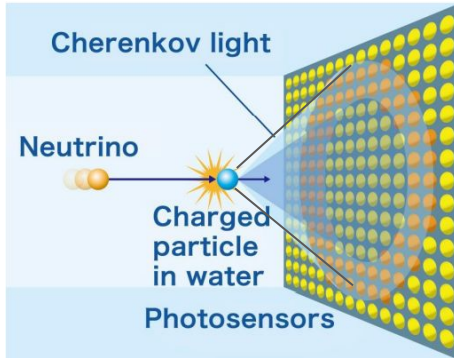
# How to search?

- Already in 1950s, Goldhaber argued on general grounds  $\tau > 10^{18}$  yrs  
→ Many searches with different methods since
- Proton lives  $\tau > 10^{30}$  yrs, how to test?
  - Look at 1 proton VERY long time ✗
  - Look at many protons for few years ✓
- **Large-scale water Cherenkov (WC) detectors - excellent targets**
  - Cheap, easily scalable (e.g. iron calorimeters are difficult)
  - Proven technology
  - $\text{H}_2\text{O}$  includes 2 hydrogen “free protons”  
(high selection efficiency, low uncertainty)



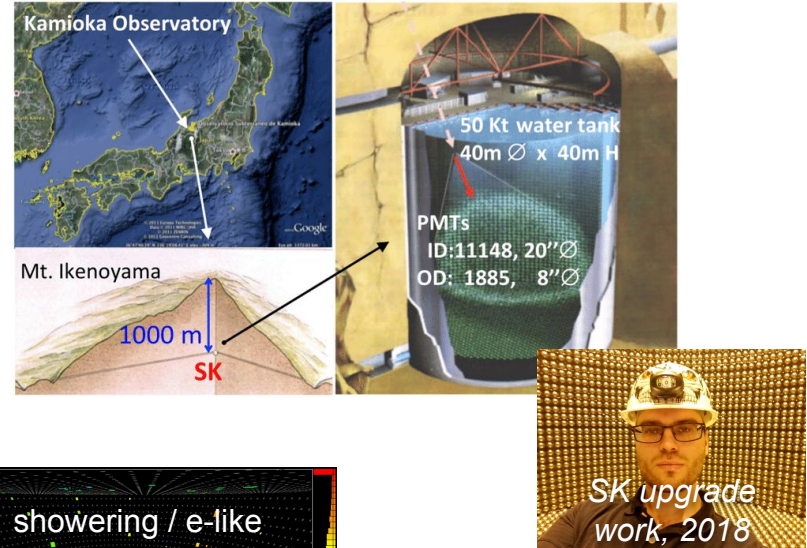


# Water Cherenkov Ring Imaging and State-of-the-Art

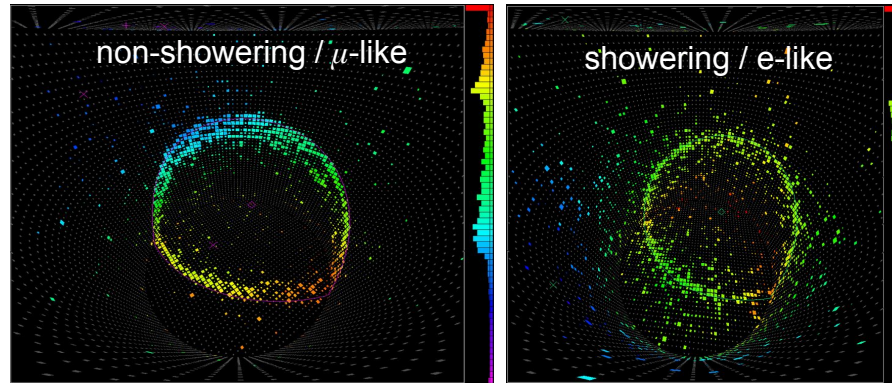


state-of-the-art

**Super-  
Kamiokande  
(SK, Super-K)**

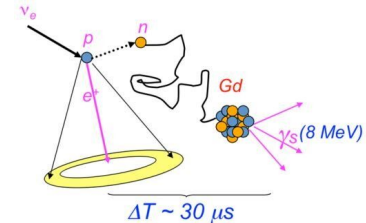
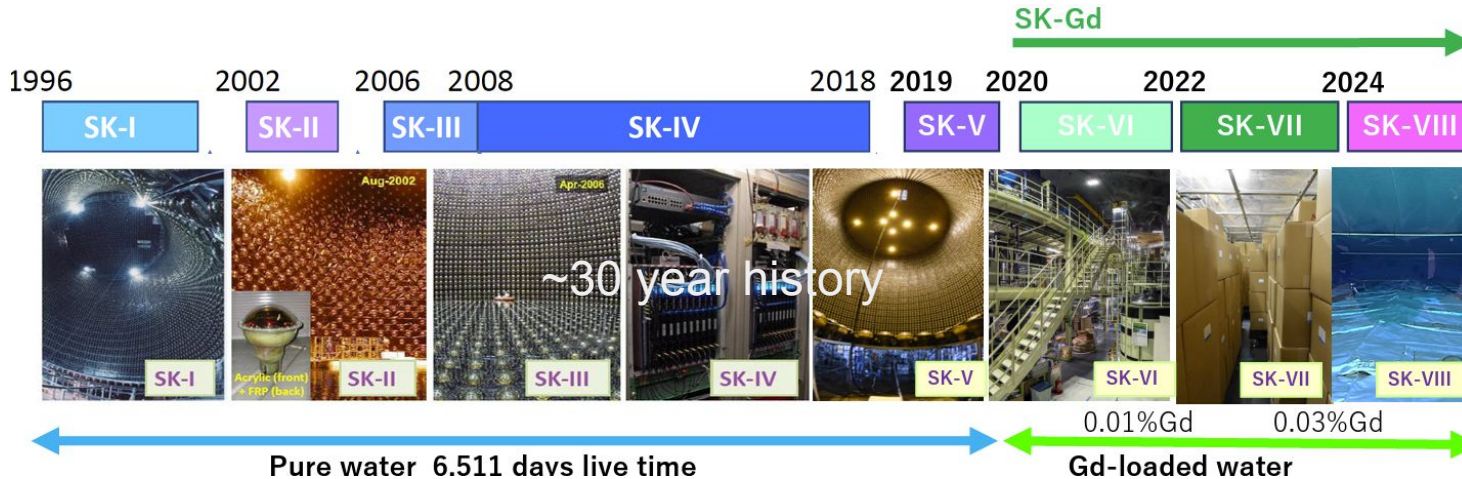
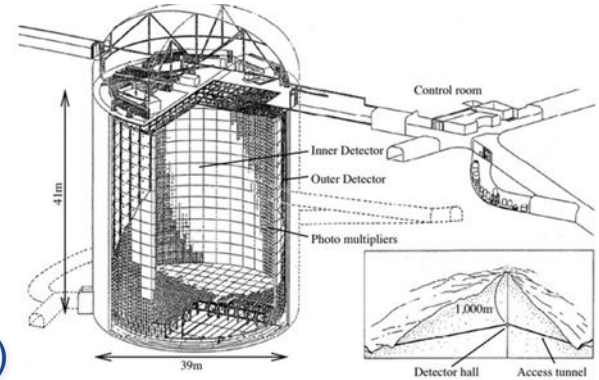


*real SK data (1998)*



# The Super-Kamiokande (SK) Experiment

- ~50 kton WC experiment
  - ~22.5 kton target fiducial volume
- 1 km underground (2700 m.w.e.) in Kamioka mine
- Inner (outer) detector: ~11,000 (~2,000) PMTs
- Since ~2020 doped with Gadolinium(Gd), ~0.03 w%
  - SKGd neutron capture efficiency ~75% (water: ~20%)



# A Multipurpose Discovery Experiment

- Outstanding detector performance

- Excellent ring-imaging capabilities; proven over 25+ years
- Particle ID > 99%, energy resolution at 1 GeV ~ percent-level

*discovery of neutrino oscillations*



2015

- Exceptional physics capabilities (energy range: ~ few -  $10^4$  MeV)

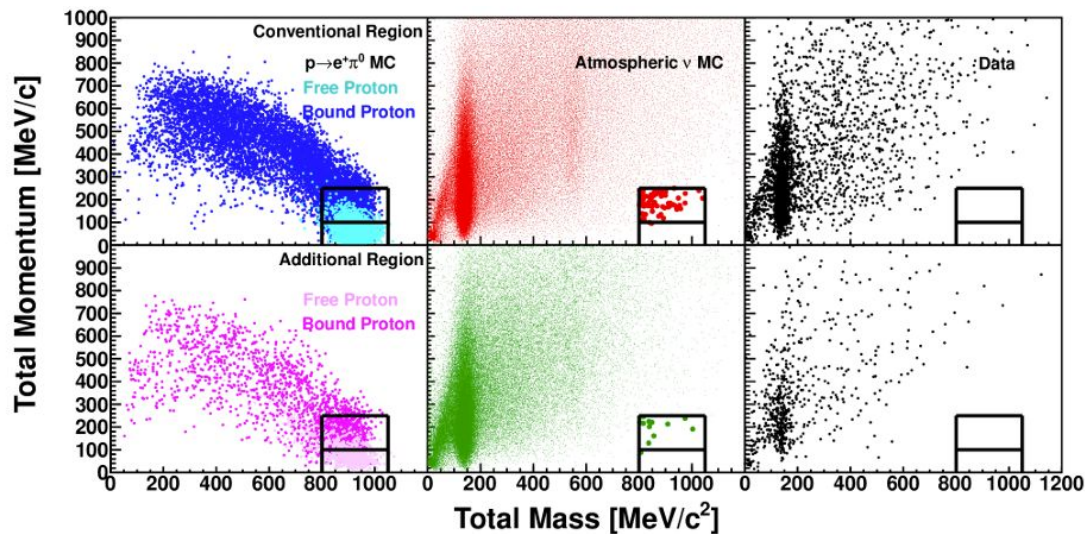
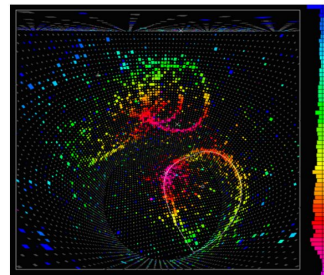
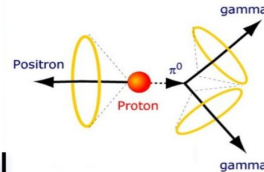
- Precision fundamental  $\nu$ -oscillation studies (atmospheric, accelerator, Lorentz inv., NSI, sterile)
- Precision solar neutrino physics (fluxes, spectral distortions, day/night effect, NSI, g-modes)
- Neutrino astronomy, supernova and beyond (multimessenger, burst, diffuse background, etc.)
- World leading sensitivity to B-violating processes (proton decay, di-nucleon decays, etc.)
- Dark matter searches: indirect (Sun, Earth, Galactic Center), direct (via cosmic-ray boost, etc.)
- Exotic targets (monopoles, Q-balls, millicharge particles, etc.)



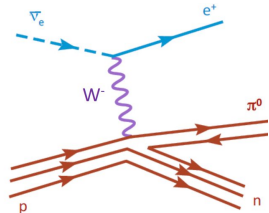


# $p \rightarrow e^+ \pi^0$ Search

- **Motivation:** Typically dominant GUT channel
- Final state (positron + 2 gammas) fully visible  
→ can reconstruct proton mass/momentum, clean channel



*typical atm.-v back.  
(benefit from n-tag)*



- Exposure: 450 kton\*yr
- Signal efficiency: 38.6% (SK I-IV)
- Expected background:  
0.05 ev. in  $P_{\text{tot}} < 100$  MeV  
0.58 ev. in  $100 \leq P_{\text{tot}} < 250$  MeV

**No candidates, lifetime limit (90% CL)**

$$> 2.4 \times 10^{34} \text{ yrs.}$$

[Abe+ (SK), PRD, 2020]

# $p \rightarrow \nu K^+$ Search

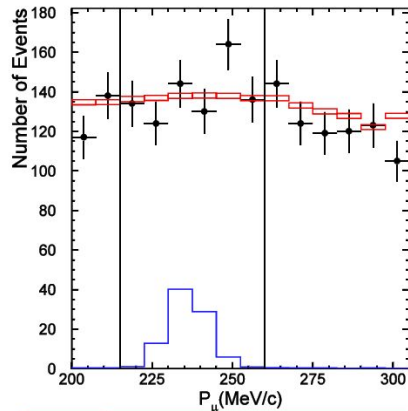
- **Motivation:** Typically dominant SUSY GUT channel
- Invisible  $\nu$ , can't reconstruct proton mass/momentum
- Momentum of  $K^+ \sim 339$  MeV below Cherenkov threshold (invisible)  $\rightarrow$  analyze decays

$$K^+ \rightarrow \nu \mu^+ : 64 \%$$

$$K^+ \rightarrow \pi^+ \pi^0 : 21 \%$$

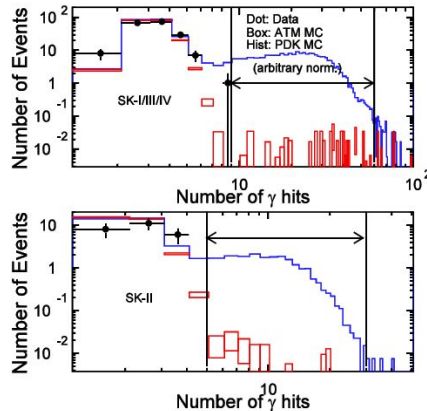
A ( $K^+ \rightarrow \mu^+ \nu$ )

$\mu$  momentum fit



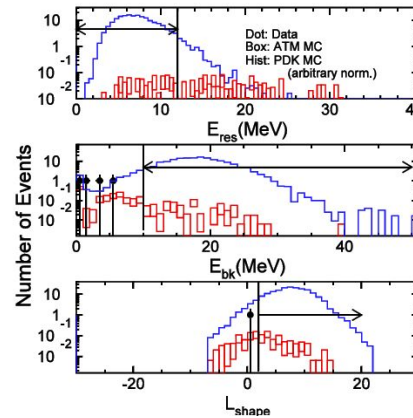
B ( $K^+ \rightarrow \mu^+ \nu$ )

nuclear de-excitation  $\gamma$  tag



C ( $K^+ \rightarrow \pi^0 \pi^+$ )

reconstruct pions



*combine analyses*

- Exposure: 365 kton\*yr
- Signal eff.:  $\sim 10\%$  (SK I-IV)

**No excess,  
lifetime limit (90% CL)**

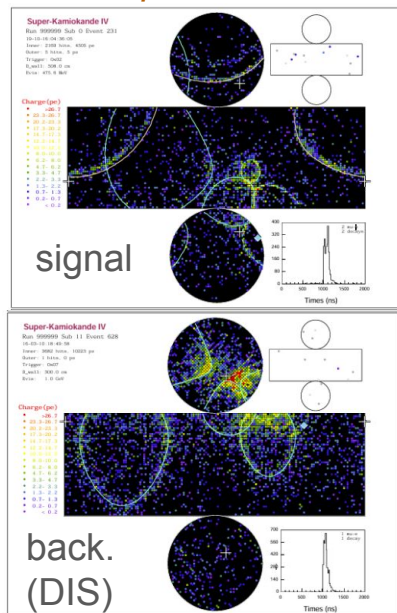
$$> 8.2 \times 10^{33} \text{ yrs.}$$

[Abe+ (SK), PRD, 2014] (Updated)

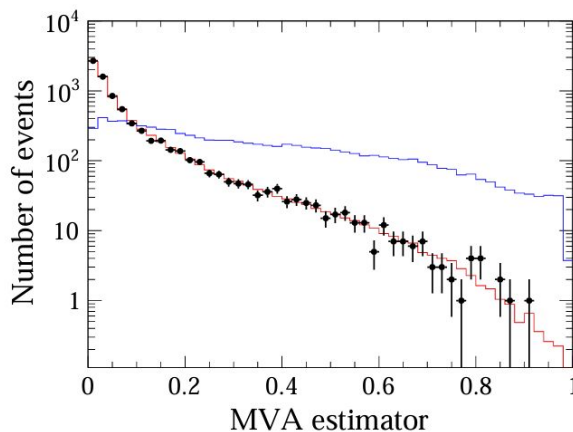
# Neutron-Antineutron Oscillations

- **Motivation:**  $\Delta B = \Delta(B - L) = 2$ , neutrino mass connections, new intermediate scales, etc.
- Capture of  $\bar{n}$  by  $^{16}\text{O}$  nucleons (n, p) produces mesons ( $\sim$ pions)  $\rightarrow$  reconstruct initial state
- Complicated final-state interactions with multi-ring signatures

## example simulations



*Multivariate analysis allows ~2x sensitivity*



- Exposure: 370 kton\*yr
- Signal efficiency: ~4% (SK I-IV, MVA)
- Expected background:  $9.3 \pm 2.7$  ev.
- Observed: 11 ev.

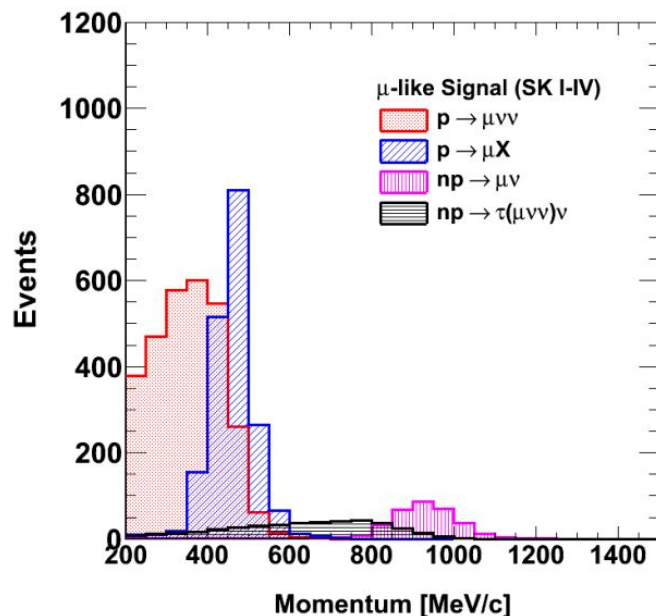
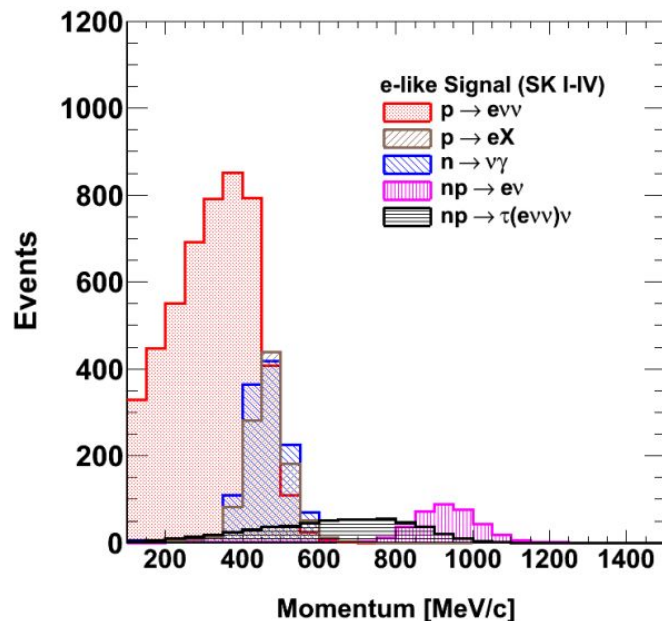
## No excess, lifetime limit (90% CL)

$T_{n-\bar{n}}(10^{32} \text{ years})$	$R (10^{23}/\text{s})$	$\tau_{n\rightarrow\bar{n}}(10^8 \text{ s})$
3.6	0.517	4.7

[Abe+ (SK), PRD, 2021]

$$T_{n-\bar{n}}^{\text{bound}} = R \cdot \tau_{n \rightarrow \bar{n}}^2$$

# Spectral Searches with Missing Energy



**No excess,  
lifetime limits  
(90% CL)**

$\geq \text{few} \times 10^{32} \text{ yrs.}$

[Takhistov+ (SK), *PRL*, (2014); Takhistov+ (SK), *PRL*, (2015); + ongoing analyses]

- Look for visible energy-momentum, distribution important (e.g.  $\nu$ , massless invisible X), highlighted in [Chen, Takhistov, *PRD*, (2014)]

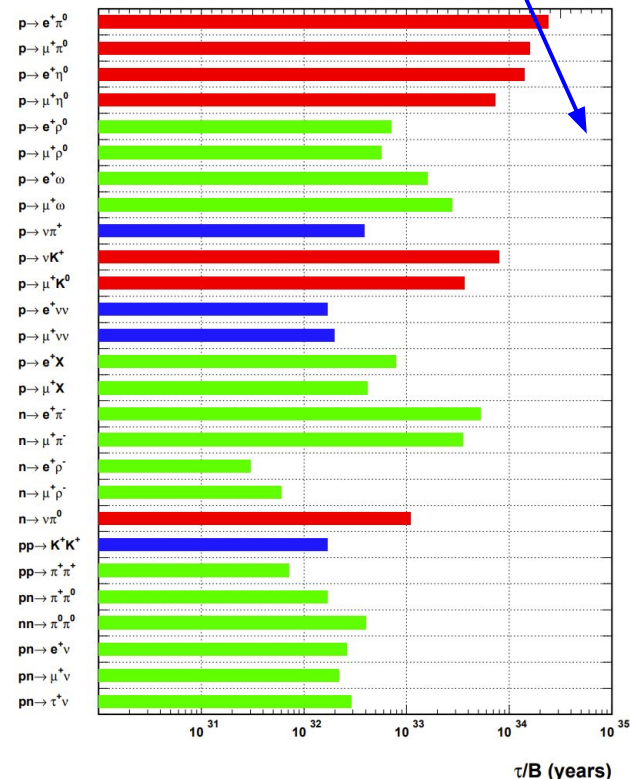


# SK Broadly Set Most Sensitive Limits

$$\tau \sim 10^{35} \text{ yrs}$$

- SK already broadly probed ~50 channels (including dinucleon decays, decays to >2 final states...)
- Analysis strategy depends on channel specifics  
→ visible Cherenkov rings, backgrounds, etc.
- Occasional candidates consistent with background, no evidence of nucleon decays  
→ leading limits on nucleon decays

*... are we done?  
No, unexplored frontier ahead !*

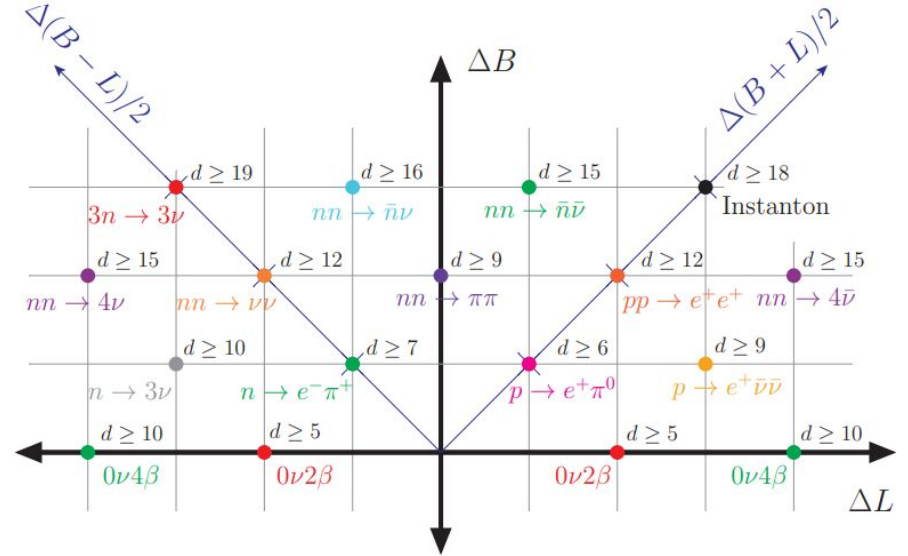


# Very Many Channels Never Searched

Channel	$ \Delta(B-L) $	$\frac{\Gamma^{-1}}{10^{30} \text{ yr}}$
$p \rightarrow e^- + e^+ + e^+$	0	793 [65]
$p \rightarrow e^- + e^+ + \mu^+$	0	529 [65]
$p \rightarrow e^+ + e^+ + \mu^-$	0	529* [65]
$p \rightarrow e^- + \mu^+ + \mu^+$	0	6 [64] (359)* [65]
$p \rightarrow e^+ + \mu^+ + \mu^+$	0	359 [65]
$p \rightarrow \mu^- + \mu^+ + \mu^+$	0	675 [65]
$p \rightarrow e^+ + 2\nu$	0.2	170 [81]
$p \rightarrow \mu^+ + 2\nu$	0.2	220 [81]
$p \rightarrow e^- + 2\pi^+$	2	30 [62] (82)* [65]
$p \rightarrow e^- + \pi^+ + \rho^+$	2	
$p \rightarrow e^- + K^+ + \pi^+$	2	75 [65]
$p \rightarrow e^+ + 2\gamma$	0	100 [82] (793)* [65]
$p \rightarrow e^+ + \rho^- + \pi^+$	0	82 [65]
$p \rightarrow e^+ + \rho^- + \pi^+$	0	
$p \rightarrow e^+ + \pi^- + \rho^+$	0	75* [65]
$p \rightarrow e^+ + \pi^- + \rho^+$	0	
$p \rightarrow e^+ + \pi^- + K^+$	0	75* [65]
$p \rightarrow e^+ + 2\pi^0$	0	140 [65]
$p \rightarrow e^+ + \pi^0 + \eta$	0	
$p \rightarrow e^+ + \pi^0 + \rho^0$	0	
$p \rightarrow e^+ + \pi^0 + \omega$	0	
$p \rightarrow e^+ + \pi^0 + K^0$	0	
$p \rightarrow \mu^+ + 2\pi^+$	2	17 [62] (133)* [65]
$p \rightarrow \mu^+ + K^+ + \pi^+$	2	245 [65]
$p \rightarrow \mu^+ + 2\gamma$	0	529* [65]
$p \rightarrow \mu^+ + \pi^- + \pi^+$	0	133 [65]
$p \rightarrow \mu^+ + K^- + \pi^+$	0	245* [65]
$p \rightarrow \mu^+ + \pi^- + K^+$	0	245* [65]
$p \rightarrow \mu^+ + 2\pi^0$	0	101 [65]
$p \rightarrow \mu^+ + \pi^0 + \eta$	0	
$p \rightarrow \mu^+ + \pi^0 + K^0$	0	
$p \rightarrow \nu + \pi^+ + \pi^0$	0.2	
$p \rightarrow \nu + \pi^+ + \eta$	0.2	
$p \rightarrow \nu + \pi^+ + \rho^0$	0.2	
$p \rightarrow \nu + \pi^+ + \omega$	0.2	
$p \rightarrow \nu + \pi^+ + K^0$	0.2	
$p \rightarrow \nu + \rho^+ + \pi^0$	0.2	
$p \rightarrow \nu + K^+ + \pi^0$	0.2	

Channel	$ \Delta(B-L) $	$\frac{\Gamma^{-1}}{10^{30} \text{ yr}}$
$nn \rightarrow \pi^0 + \phi$	2	
$nn \rightarrow 2\eta$	2	
$nn \rightarrow \eta + \rho^0$	2	
$nn \rightarrow \eta + \omega$	2	
$nn \rightarrow \eta + \eta'$	2	
$nn \rightarrow \eta + K^0$	2	
$nn \rightarrow \eta + K^{*,0}$	2	
$nn \rightarrow \eta + \phi$	2	
$nn \rightarrow 2\rho^0$	2	
$nn \rightarrow \rho^0 + \omega$	2	
$nn \rightarrow \eta' + \rho^0$	2	
$nn \rightarrow K^0 + \rho^0$	2	
$nn \rightarrow K^{*,0} + \rho^0$	2	
$nn \rightarrow \rho^0 + \phi$	2	
$nn \rightarrow \rho^- + \rho^+$	2	
$nn \rightarrow K^{*+} + \rho^-$	2	
$nn \rightarrow K^{*-} + \rho^+$	2	
$nn \rightarrow K^{*0} + \rho^0$	2	
$nn \rightarrow K^{*+} + \rho^-$	2	
$nn \rightarrow 2\omega$	2	
$nn \rightarrow \eta' + \omega$	2	
$nn \rightarrow K^0 + \omega$	2	
$nn \rightarrow K^{*,0} + \omega$	2	
$nn \rightarrow \omega + \phi$	2	
$nn \rightarrow \eta' + K^0$	2	
$nn \rightarrow \eta' + K^{*,0}$	2	
$nn \rightarrow K^- + K^+$	2	170* [116]
$nn \rightarrow K^+ + K^{*-}$	2	
$nn \rightarrow K^- + K^{*+}$	2	
$nn \rightarrow 2K^0$	2	
$nn \rightarrow K^{*,0} + K^0$	2	
$nn \rightarrow K^0 + \phi$	2	
$nn \rightarrow 2K^{*,0}$	2	
$nn \rightarrow K^{*-} + K^{*+}$	2	

Channel	$ \Delta(B-L) $	$\frac{\Gamma^{-1}}{10^{30} \text{ yr}}$
$n \rightarrow \nu + e^- + e^+$	0.2	257 [65]
$n \rightarrow \nu + e^- + \mu^+$	0.2	83 [65]
$n \rightarrow \nu + e^+ + \mu^-$	0.2	83* [65]
$n \rightarrow \nu + \mu^- + \mu^+$	0.2	79 [65]
$n \rightarrow 3\nu$	0.2, 4	0.58 [83]
$n \rightarrow e^- + \pi^+ + \pi^0$	2	29 [62] (52)* [65]
$n \rightarrow e^- + \pi^+ + \eta$	2	
$n \rightarrow e^- + \pi^+ + \rho^0$	2	
$n \rightarrow e^- + \pi^+ + \omega$	2	
$n \rightarrow e^- + \pi^+ + K^0$	2	
$n \rightarrow e^- + \rho^+ + \pi^0$	2	
$n \rightarrow e^- + K^+ + \pi^0$	2	
$n \rightarrow e^+ + \pi^- + \pi^0$	0	52 [65]
$n \rightarrow e^+ + \pi^- + \eta$	0	
$n \rightarrow e^+ + \pi^- + \rho^0$	0	
$n \rightarrow e^+ + \pi^- + \omega$	0	
$n \rightarrow e^+ + \pi^- + K^0$	0	18 [62]
$n \rightarrow e^+ + \rho^- + \pi^0$	0	
$n \rightarrow e^+ + K^- + \pi^0$	0	
$n \rightarrow \mu^- + \pi^+ + \pi^0$	2	24 [62] (14)* [65]
$n \rightarrow \mu^- + \pi^+ + \eta$	2	
$n \rightarrow \mu^- + \pi^+ + K^0$	2	
$n \rightarrow \mu^- + K^+ + \pi^0$	2	
$n \rightarrow \mu^+ + \pi^- + \pi^0$	2	44 [65]
$n \rightarrow \mu^+ + \pi^- + \eta$	0	
$n \rightarrow \mu^+ + \pi^- + K^0$	0	
$n \rightarrow \mu^+ + K^- + \pi^0$	0	
$n \rightarrow \nu + 2\gamma$	0.2	219 [65]
$n \rightarrow \nu + \pi^- + \pi^+$	0.2	
$n \rightarrow \nu + \rho^- + \pi^+$	0.2	
$n \rightarrow \nu + K^- + \pi^+$	0.2	
$n \rightarrow \nu + \pi^- + \rho^+$	0.2	
$n \rightarrow \nu + \pi^- + K^+$	0.2	
$n \rightarrow \nu + 2\pi^0$	0.2	
$n \rightarrow \nu + \pi^0 + \eta$	0.2	
$n \rightarrow \nu + \pi^0 + K^0$	0.2	
$n \rightarrow \nu + \pi^0 + \rho^0$	0.2	
$n \rightarrow \nu + \pi^0 + \omega$	0.2	
$n \rightarrow \nu + \pi^0 + K^0$	0.2	



Not exhaustive list

[Heeck, Takhistov, PRD, (2019)]

# New Broad Program: Non-Canonical Nucleon Decays

*Novel probes of light new physics: axion-like particles, dark photons, sterile  $\nu$ 's...*

$\mathcal{O}$	Operator	$(\Delta B, \Delta L)$	Dim	Decay modes	New Field(s)
$\mathcal{O}_{d^2uN}$	$\epsilon^{abc} (d_a N) (d_b u_c)$	(1, 1)	6	$p(n) \rightarrow \pi^{+(0)} N$	sterile neutrinos
$\mathcal{O}_{Dd^2u\tilde{N}}$	$\epsilon^{abc} (d_a \sigma_\mu N^\dagger) (d_b D^\mu u_c)$	(1, -1)	7	$n \rightarrow N \gamma$ $p(n) \rightarrow \pi^{+(0)} N \gamma$	sterile neutrinos
$\mathcal{O}_{du^2e\phi}$	$\epsilon^{abc} (d_a u_b) (e u_c) \phi$	(1, 1)	7	$p \rightarrow e^+ \phi$ $p(n) \rightarrow e^+ \pi^{0(-)} \phi$	dark scalars, majorons
$\mathcal{O}_{d^2QLX}$	$\epsilon^{abc} (Q_a^i \sigma^\mu d_b) (d_c L_i^\dagger) X_\mu$	(1, 1)	7	$n \rightarrow \nu X$ $p \rightarrow \nu \pi^{+(0)} X$ $p(n) \rightarrow e^+ \pi^{0(-)} X$	dark photons
$\mathcal{O}_{dQ^2LH\phi}$	$\epsilon^{abc} (Q_a^i Q_b^j) (L_i^\dagger d_c) H_j^\dagger \phi$	(1, 1)	8	$p \rightarrow e^+ \phi$ $p(n) \rightarrow e^+ \pi^{0(-)} \phi$	dark scalars, majorons
$\mathcal{O}_{Dd^2QLa}$	$\epsilon^{abc} (\partial^\mu a) (Q_a^i \sigma^\mu d_b) (d_c L_i^\dagger)$	(1, 1)	8	$p(n) \rightarrow \nu \pi^{+(0)} a$ $p(n) \rightarrow e^+ \pi^{0(-)} a$	axion-like particles
$\mathcal{O}_{Dd^2uNa}$	$\epsilon^{abc} (\partial^\mu a) (d_a \sigma_\mu N^\dagger) (d_b u_c)$	(1, -1)	8	$n \rightarrow Na$ $p(n) \rightarrow \pi^{+(0)} Na$	axion-like particles with sterile neutrinos
$\mathcal{O}_{duQeLN}$	$\epsilon^{abc} (e u_a) (Q_b^i \sigma_\mu d_c) (N^\dagger \sigma^\mu L_i^\dagger)$	(1, 1)	9	$p \rightarrow e^+ \nu N$ $n \rightarrow e^+ e^- N$	sterile neutrinos
$\mathcal{O}_{du^2eN^2}$	$\epsilon^{abc} (d_a u_b) (e u_c) (NN)$	(1, 3)	9	$p \rightarrow e^+ NN$	sterile neutrino

- Many novel possibilities mediated by motivated BSM particles
- Multi-nucleon decays possible (dinucleon, trinucleon...)
- Mixed scenarios possible
- Tests of light new physics over decades in mass below  $\sim$  few GeV

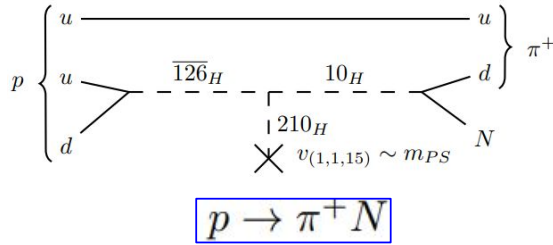
[Fridell, Hati, Takhistov, *PRD Lett.* (2024); + in preparation & others]

→ **SK uniquely positioned to lead this**

# Non-Canonical and Dark Matter-Induced Decay Signatures

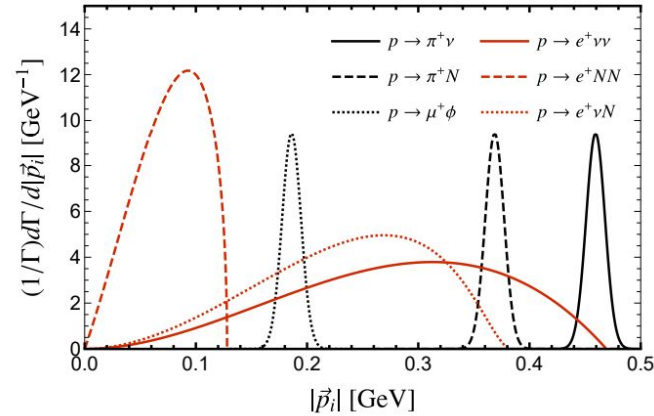
Ex.)

heavy sterile  $\nu$ 's ( $N$ ) in SO(10)  
GUT with Pati-Salam



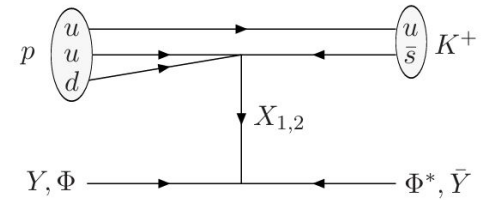
Predicted lifetime can be  
 $\sim 10^{34} - 10^{35}$  yrs  
 $\sim$ testable in SK

Ex.)



$$m_N = 400 \text{ MeV} \quad m_\phi = 700 \text{ MeV}.$$

Complementary process class:  
“induced decays” by dark matter  
with (anti)baryon charge



$$m_\Phi, m_Y \sim \text{few} \times \text{GeV}$$

[Davoudiasl+, PRL, 2010; + others]

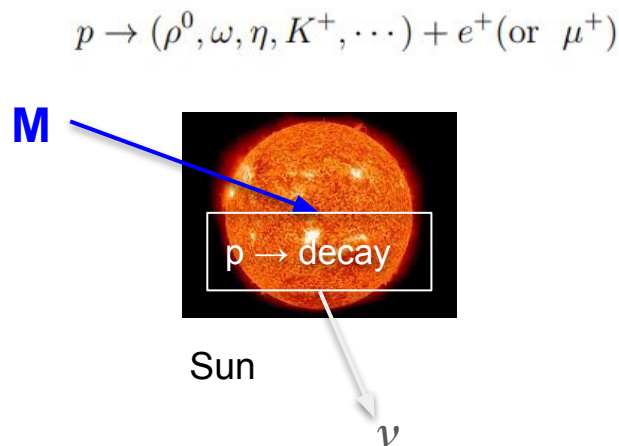
- Conventional searches can misinterpret or completely miss such channels
- Inclusive searches can probe many modes simultaneously [Heeck, Takhistov, PRD (2019)]

[Fridell, Hati, Takhistov, PRD Lett. (2024)]

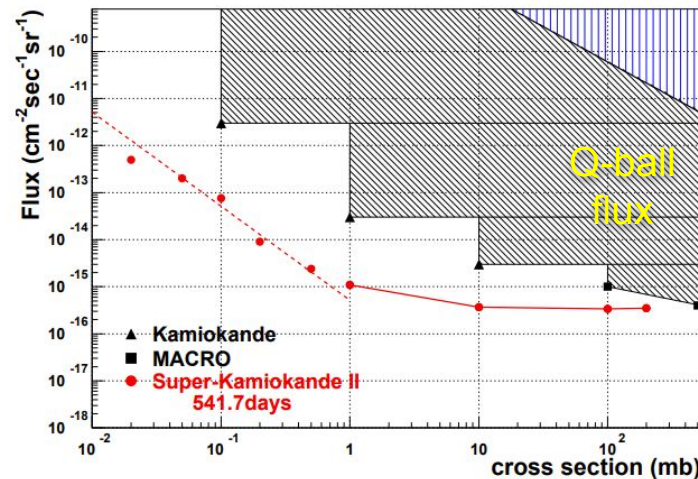
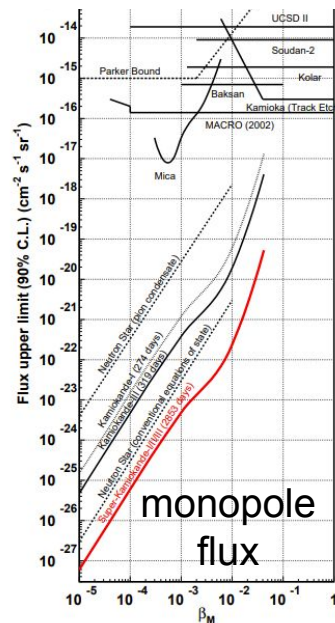


# Decay Catalysis by Extended Objects: Monopoles, Q-Balls

- Heavy monopoles from unified theories can be captured by Sun and catalyze proton decay via Callan-Rubakov process producing neutrinos
- Q-balls naturally appear from field instabilities in e.g. supersymmetric theories, dark matter Q-balls can catalyze nucleon decay in SK when carrying baryon number



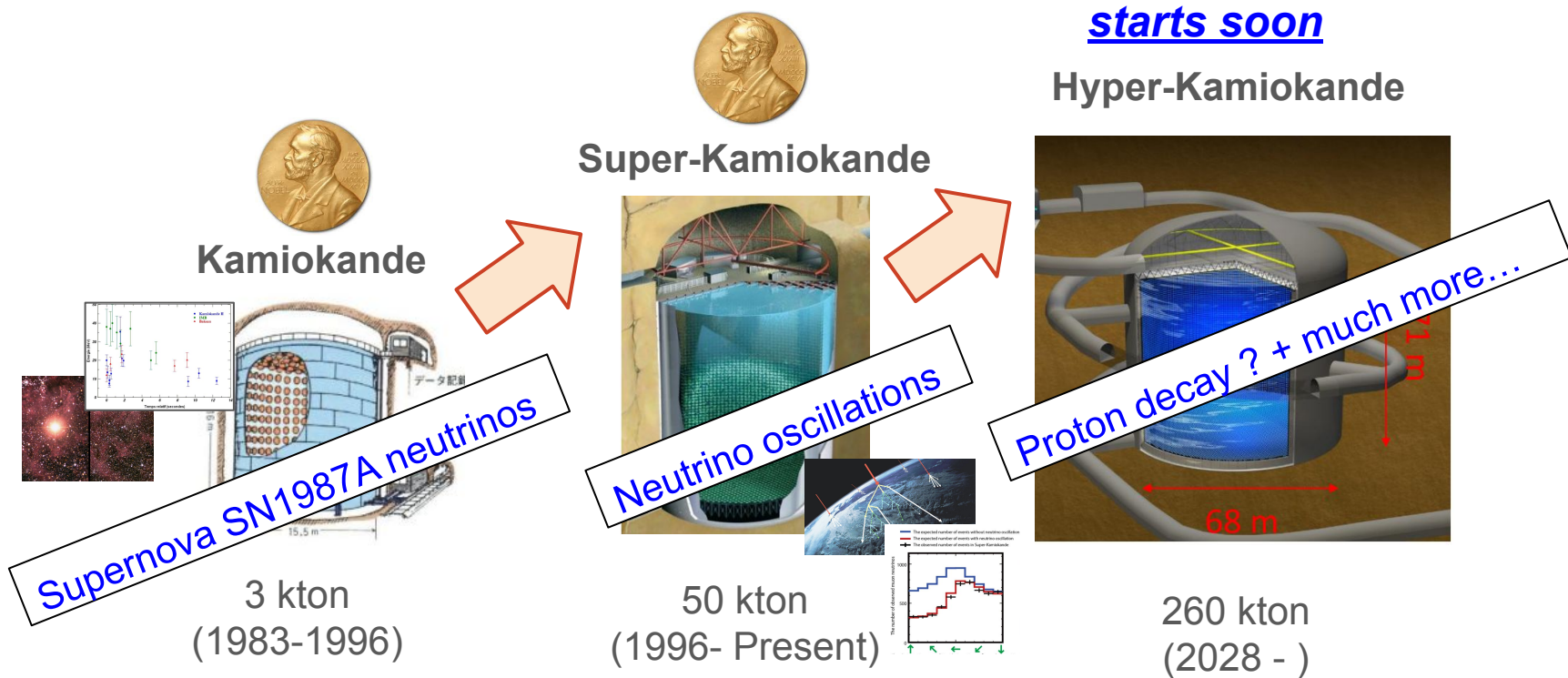
[Ueno+ (SK), 2012;  
+ Feng (SK), M.S. thesis, 2021]



[Takenaga+ (SK), 2007]

Q-ball dark matter flux  $\sim \Phi_Q \sim \frac{\rho_{\text{DM}}}{M_Q} v$

# Into the Future with Hyper-Kamiokande (Hyper-K, HK)



# The Hyper-Kamiokande Experiment

- **Next generation flagship neutrino observatory in Japan**

- $\sim 8.4 \times$  SK fiducial mass

- **Big construction progress**

- cavern excavation  $\rightarrow$  *complete* (2025) ✓

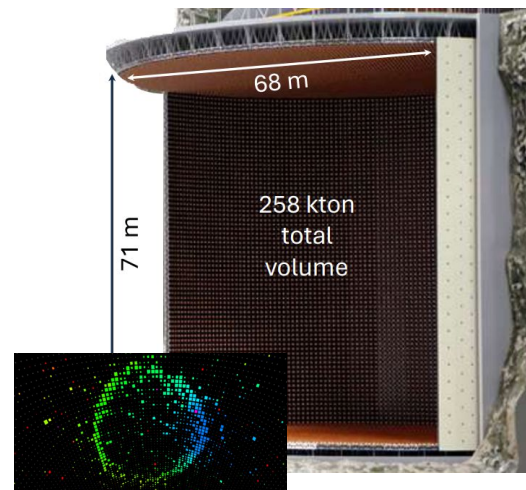
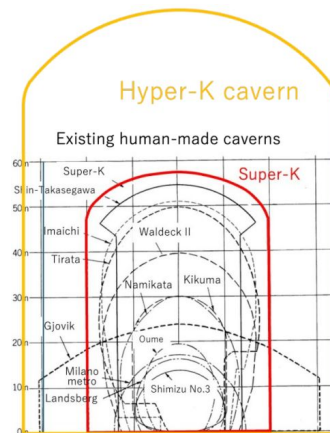
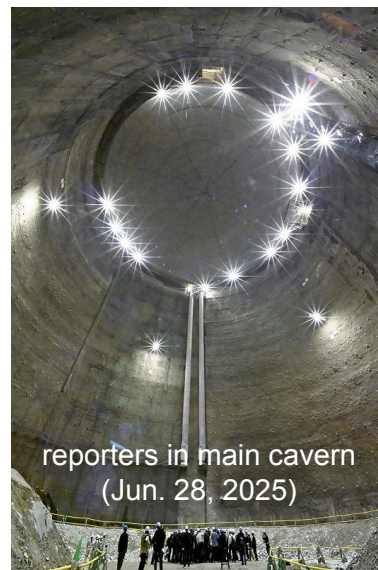
- **Engineering marvel**

$\rightarrow$  largest man-made cavern in world

- **Operation start 2028**

- R&D of all stages finalizing
- PMT mass production continues

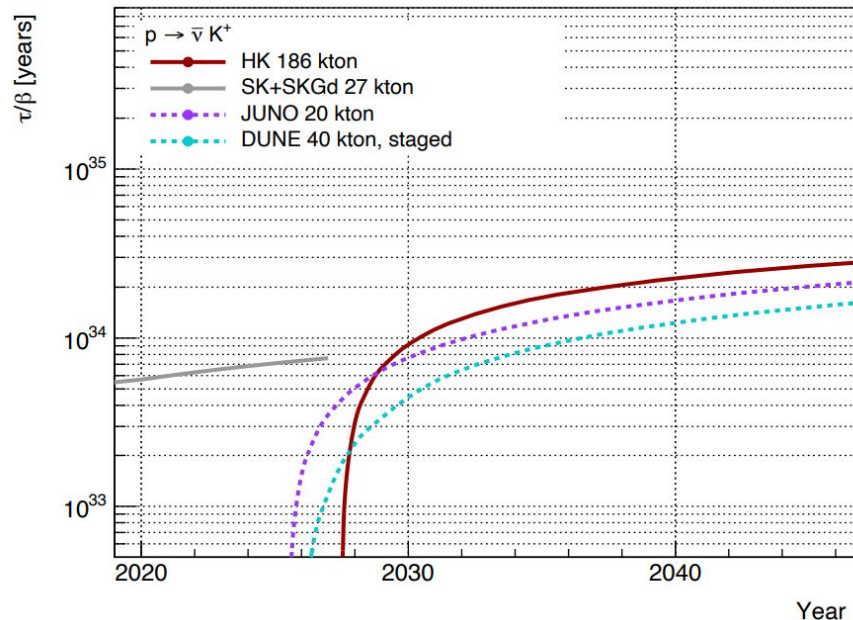
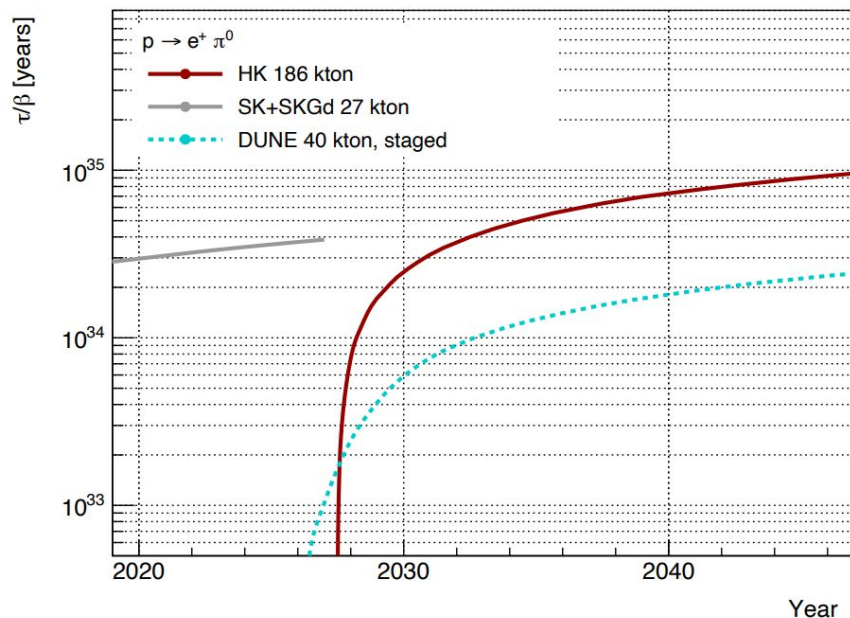
	Super-K	Hyper-K
Site	Mozumi	Tochibora
Overburden	2780 m.w.e.	1700 m.w.e.
Number of ID PMTs	11129	20000
Photo-coverage	40%	20% ( $\times 2$ efficiency)
Mass / Fiducial Mass	50 kton / 22.5 kton	258 kton / 186 kton





# Going Beyond Existing Reach

- Uniquely positioned to broadly lead nucleon decay searches
- Only approved experiment capable of reaching  $\tau \sim 10^{35}$  yrs



[HK Collab., (2025) arXiv: 2506.16641]



# Conclusions

- Super-Kamiokande is an exceptional multipurpose experiment and has established the world's most sensitive limits on a wide range of nucleon decay and baryon number violation processes
- Wide range of searches are pursued, covering conventional channels, multi-nucleon decays,  $n-\bar{n}$  oscillations and other modes
- No evidence for nucleon decay has been observed, with lifetime limits reaching  $\sim \text{few} \times 10^{34}$  yrs  $\rightarrow$  SK already excluded variety of theoretical models
- SK has broad search program ahead. Ongoing analyses and SK-Gd upgrade will further expand sensitivity and discovery potential
- Hyper-Kamiokande will build on this foundation with a broad nucleon decay program, extending searches to new frontiers and capabilities of uniquely probing lifetimes approaching  $\sim 10^{35}$  yrs lifetime