



# Hobbies/Interests

Physics (who would have thought?)

Astronomy and Stargazing

Art, Music, Coffee







## Overview of projects in Hadrons

#### Week 1, 3&4

- Particle Interactions in Scintillator with Dr. Mifuyu Ukai
- Further guidance by Dr. Yamamoto
   Takeshi, Doctoral students Ryoh Imamoto and Seong
   Chesu

#### Week 2

- Further lectures by Dr. Kyoichiro Ozawa
- Additional simulation tasks on C++/ROOT

#### Weeks 5-7

- Testing of apparatus for E16 experiment
- Semiconductor detectors with Prof. Dr. Kazuya Aoki
- SAMIDARE board (for GEM Trackers) with Doctoral student Nagafusa Shunnosuke

#### Week 8

 Report and Presentation

# Why understand scintillators and interaction of radiation with them?

Motivation: Improve resolution!

- width in signal peaks seen (i.e. standard dev / FWHM)

Investigate the various factors affecting energy resolution of the peak

- are certain materials suitable for strange hadron experiments?

Many applications to improving precision in high-energy physics experiments

Finding combination that optimizes balance of computation time and precision

## Scintillators

Work by luminescence

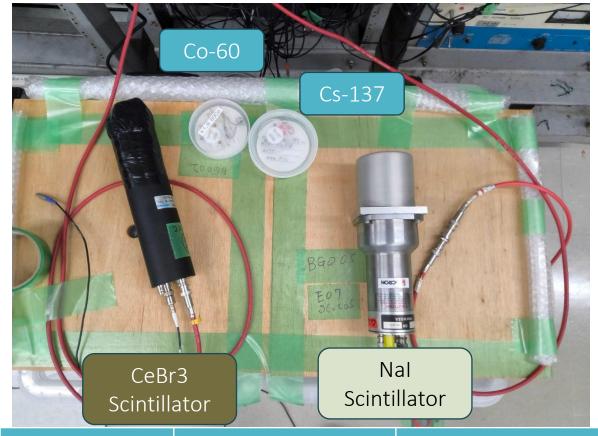
- excitation and de-excitation processes

Energy of incident radiation is detected when radiation loses energy through interactions with the scintillator crystal

Ideal characteristics:

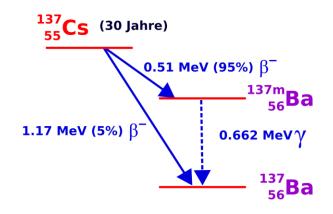
- Linear response
- High light yield
- High Quantum Efficiency

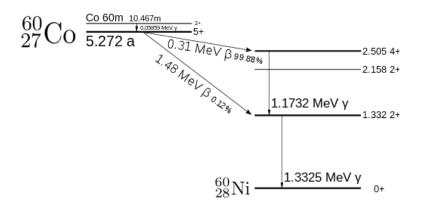
(Many other pros and cons not written here)



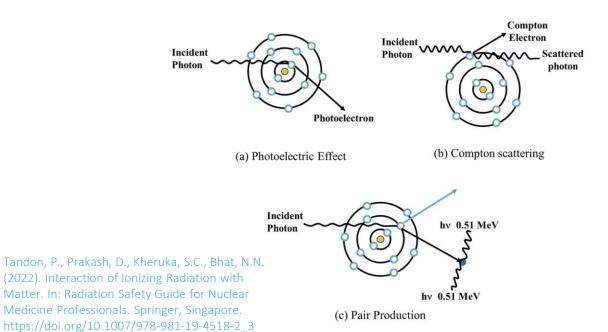
Inorganic Material:	Nal	CeBr3
Length of Cs- 137 signal	1011 +/- 14 ns	136 +/- 1 ns
Decay time (literature)	~250ns	~17 ns
Resolution	Relatively high	Relatively low

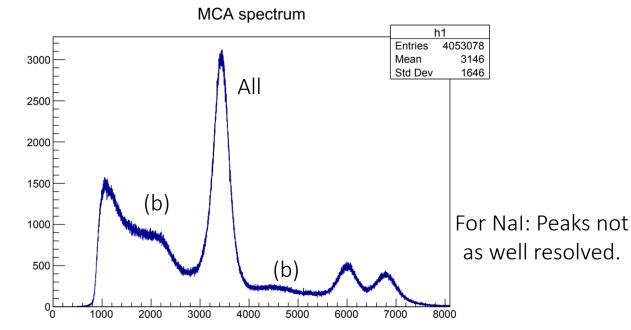
## Decay Energies



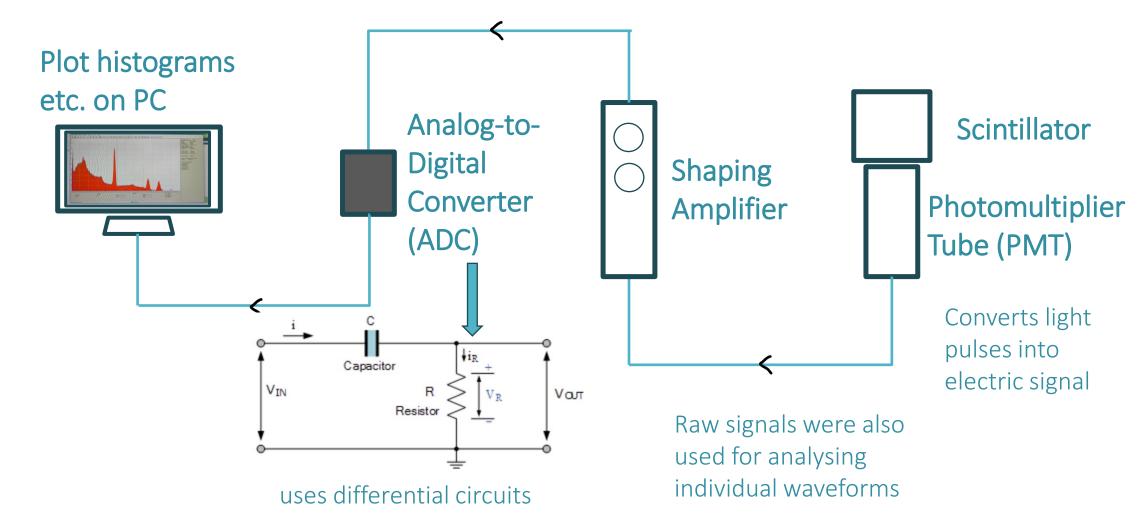


### Interaction of Gamma Rays with matter and Detection





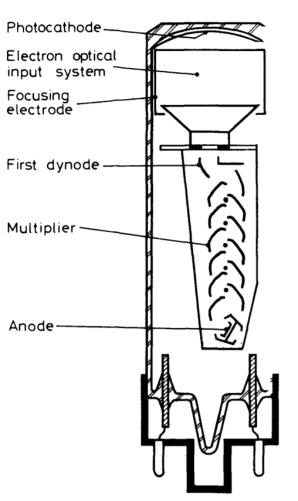
## Data-taking



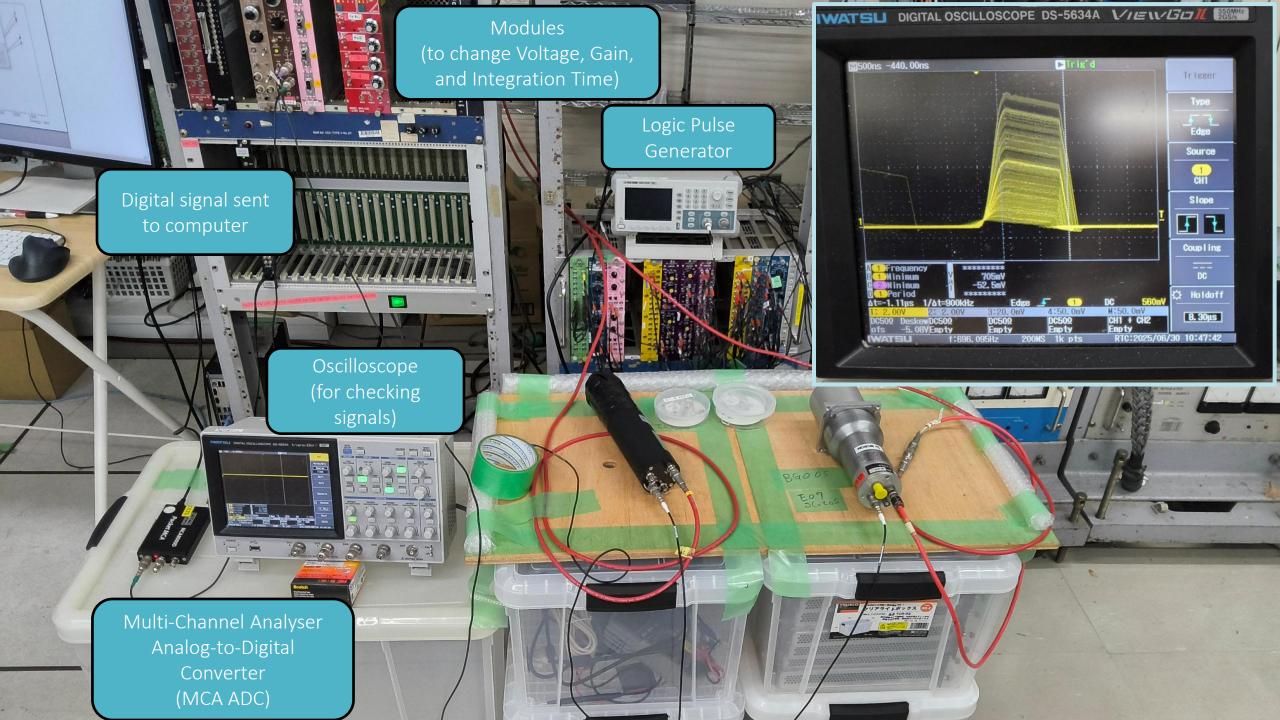


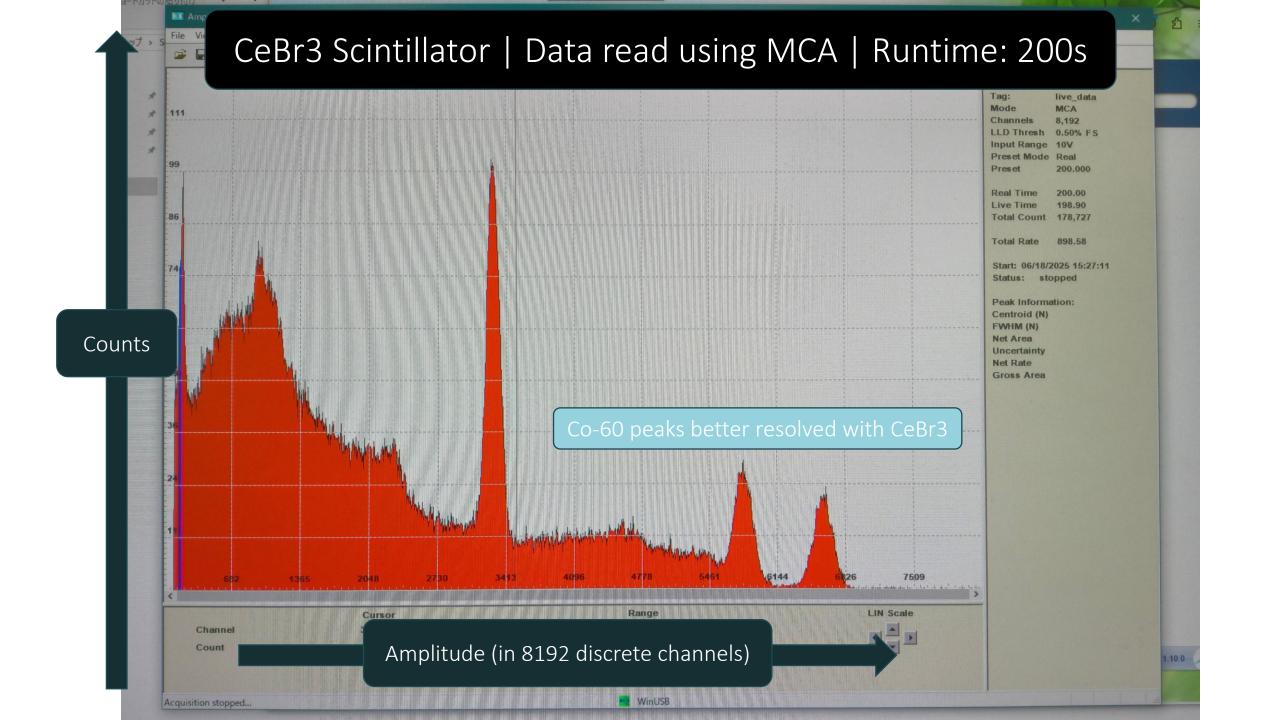
# Incident gamma radiation on scintillator crystal

- 1. Photoelectric Effect
- 2. Compton Scattering
  - 3. Pair Production

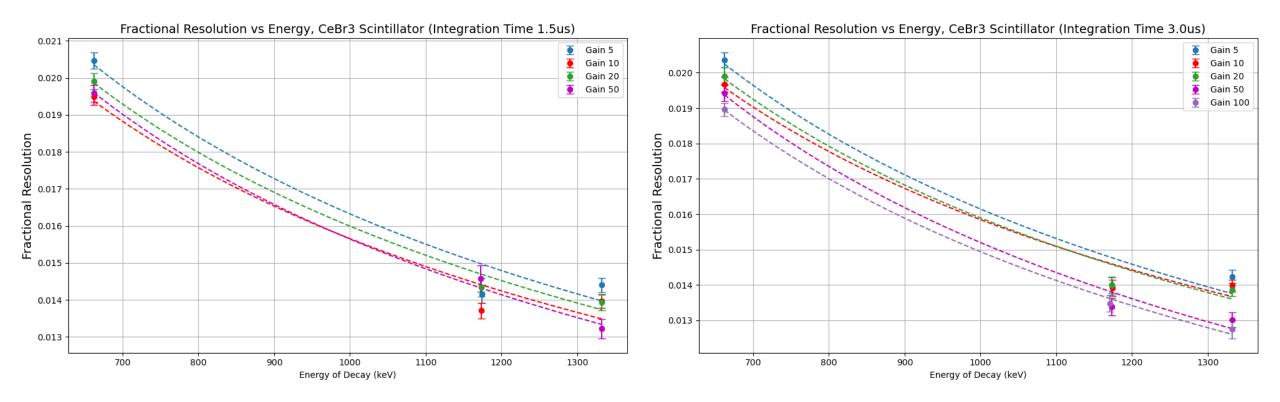


Photomultiplier schematic diagram (From W. R. Leo, 1994)





## Effect of Gain & Integration Time on Resolution



Data presented for CeBr3.

General trend: Improved resolution (i.e. lower fractional resolution) with increasing Gain. Very slight improvement in resolution with higher integration time.

Best fit curve  $A + \frac{B}{\sqrt{E}}$  needs more decay sources for improved significance.

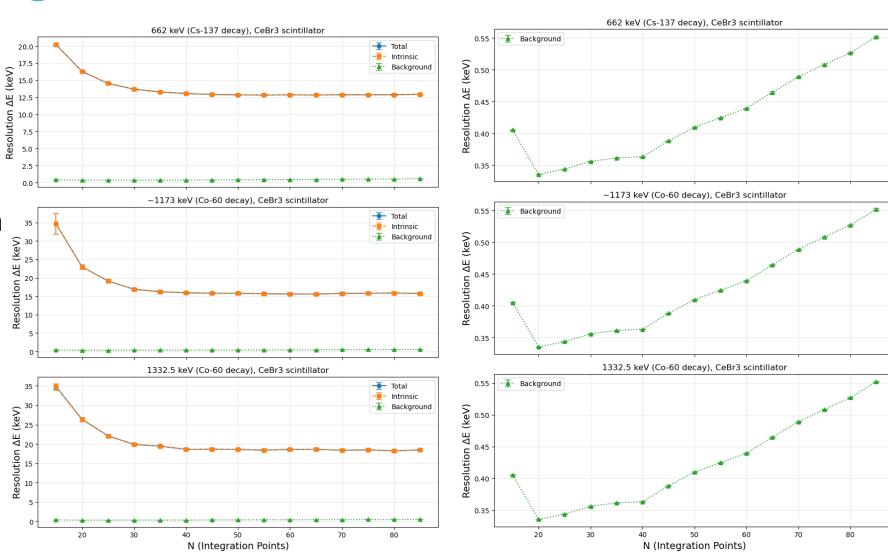
# Integrating raw waveforms – how many points (N) should we integrate? Background Plots

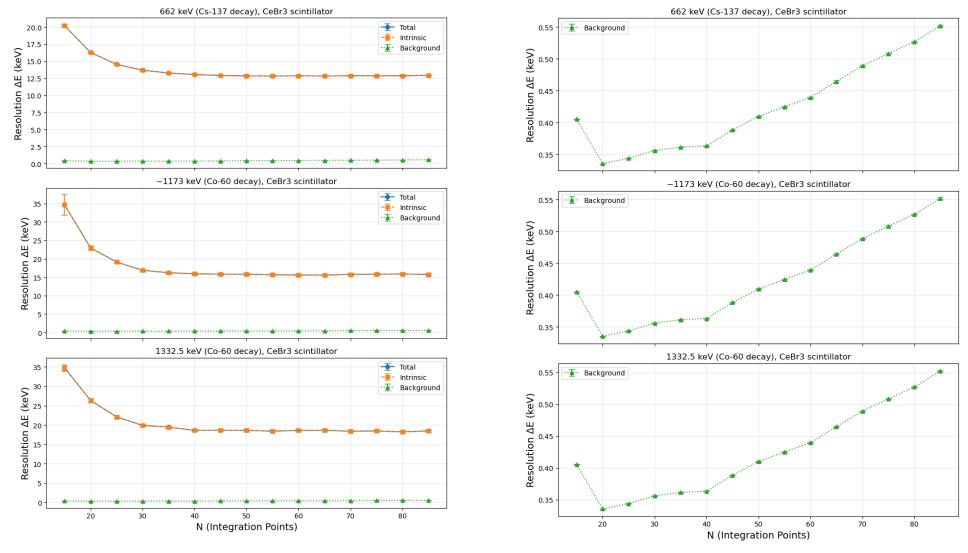
Integrating gives us the signal in terms of "channels"

Make energy calibration functions -> convert channels to energy

Find contribution of intrinsic factors and background to the peak resolutions

$$\sigma_{intrinsic} = \sqrt{\sigma_{total}^2 - \sigma_{background}^2}$$

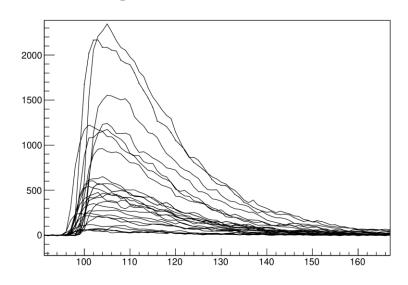




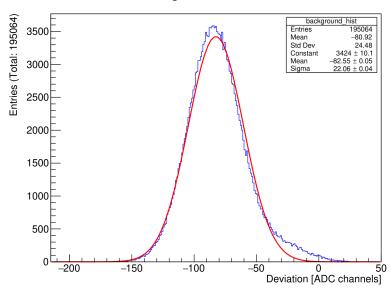
Expect background contribution to resolution to be proportional to sqrt(N) My results show a linear relationship and even a spike at lower N -> Could be due to subtracting baselines at the wrong time

#### Example CeBr3 raw signals

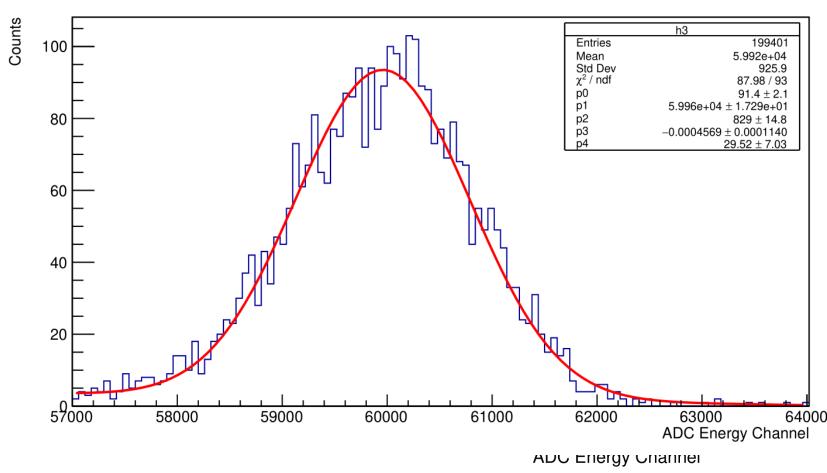
#### - Integrate → Area



#### **Background Distribution**







## Learning points

Deviation likely caused by subtraction of baseline from integrals

- Potentially another reason could exist

A deep physics understanding of particle interactions is needed to detect radiation

Signal processing is a sophisticated, multi-layer process

Every result, every waveform, every anomaly you see in a detector has an explanation.

Important of planning layout early

- Prevent hidden coding errors -> Saves time

### Next course of action

Re-take data with my improved code

Simulate list-wave compression

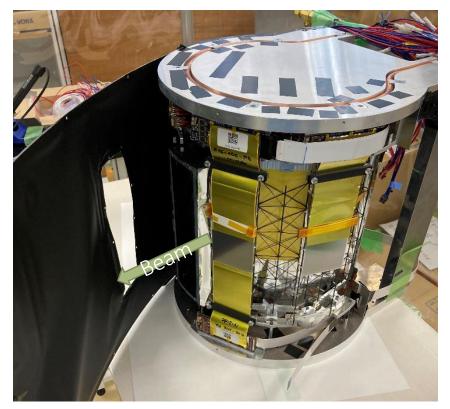
- Investigate optimal sampling rate for data precision and file sizes

Organise all plots (for NaI and CeBr3 detectors) prior to starting the report

# A Brief Summary of E16-related projects

- 1. Silicon Tracking System calibration with Prof. Aoki Kazuya
- 2. SAMIDARE boards with Doctoral student Nagafusa Shunnosuke (Kyoto U.)





## E16 experiment

Spontaneous Chiral Symmetry Breaking

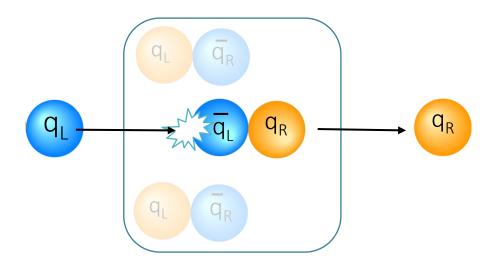
- Origin of mass in hadrons!

Mass loss in  $p + A \rightarrow \rho/\omega/\phi \rightarrow e^+e^-$  reaction

- Temperature and density dependent

$$\psi_R o \psi_R' = \psi_R e^{i lpha}$$
  $\psi_R:$  spin momentum  $\psi_L o \psi_L' = \psi_L e^{-i lpha}$   $\psi_L:$  momentum

If lagrangian unchanged upon chiral exchange -> R-L values conserved



## Progress

### STS calibration experiment

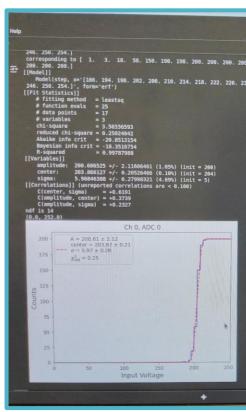
Threshold voltage for each channel-ADC Preliminary data taken

Database of over 4000 detector response plots *Yet to begin* 

#### SAMIDARE testing

Data-taking voltage inputs with triggers Some data taken

Plotting of data readout response curve Yet to begin







## Photo credits

Cover images (UK, Japan, Singapore):

- Mengyang Liu on Pexels
- Tomas Malik on Pexels
- Timo Volz on Pexels

Bishan Park HDB Flats:

- Solaman Soh on Pexels

Cobalt-60 decay diagram:

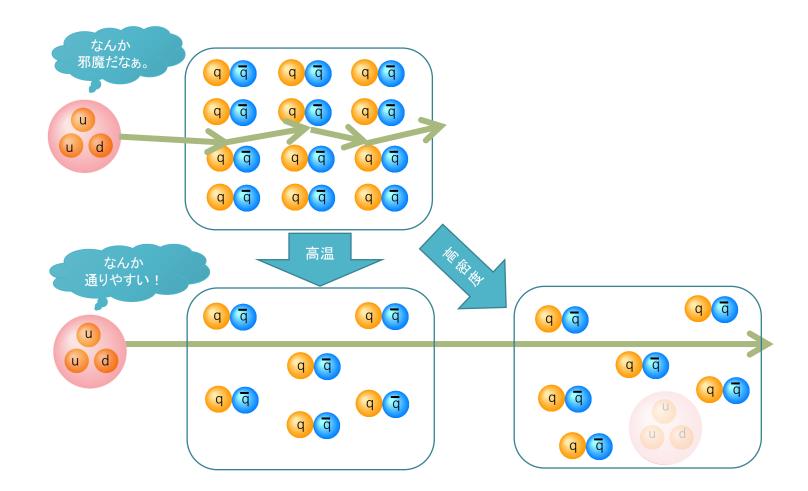
- Nuclear power [dot] com

E16 setup in Hadron Hall, Quark chirality and Interactions:

- Dr. Aoki Kasuya's introduction sldies

# Backup slides / photo bank

# Temperature and Density affects $(\sum invariant\ mass)!$



## Formation of exotic nuclei

Kaon experiments

Lambda baryons...

Iso-spin → Eventually led into QCD

Nucleon-nucleon interactions: n,p are the same!



Two states of same particle!

→ Isospin