

# Nonlinear-Supersymmetric General Relativity Theory(NLSGR)

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## OUTLINE

1. Motivation
2. **Non**linear-**s**upersymmetric **g**eneral **r**elativity theory(**NLSGR**)
3. Evolution of **NLSGR** (**Big Collapse** and **Superon-Graviton Model**)
4. Linearization of NLSUSY and the vacuum of  $L_{SGM}(e, \psi)$  (**NL/L SUSY relation**)
5. Cosmological Implications of **NLSGR**.
6. Summary

# 1. Motivations

@ The success of **GR** and **SM(GWS) model**, still **many unsolved fundamental problems are left**, e.g.,

- Unification of two SMs,
  - Space-time dimension *four*,
  - Three-generations structure of quarks and leptons,
  - C-asymmetry of nature
  - Tiny Neutrino mass  $M_\nu$ , proton stability in GUT
  - Dark Matter, Dark Energy;  $\rho_{D.E.} \sim (M_\nu)^4 \Leftrightarrow \Lambda CDM$ , Inflation
- $\Rightarrow$  **SUSY, SUGRA**,
- Origin of SUSY breaking ?,  $\dots$  etc.

@ As for the **three-generations structure**

among  $SO(N)$  LSUSY sP, only  $N = 10$  gives SM with just 3 generations, where

- The decomposition of 10 supercharges  $Q^I$ , ( $I = 1, 2, \dots, 10$ ) are:

$$\underline{10}_{SO(10)} = \underline{5}_{SU(5)} + \underline{5}^*_{SU(5)}$$

$$\underline{5}_{SU(5)} = [\{\underline{3}^{*c}, \underline{1}^{ew}, (\frac{e}{3}, \frac{e}{3}, \frac{e}{3}) : Q_a(a = 1, 2, 3)\}, \{\underline{1}^c, \underline{2}^{ew}, (0, -e) : Q_m(m = 4, 5)\}].$$

$\Leftrightarrow$  **Supercharge**  $\underline{5}_{SU(5)}$  has the same quantum numbers as  $\underline{5}$  of SU(5) GUT.

- Massless helicity states of **gravity supermultiplet** of SO(10) sP with CPT conjugation are specified by the helicity  $h = (2 - \frac{n}{2})$  and the dimension  $\underline{d}_{[n]} = \frac{10!}{n!(10-n)!}$ :

$$|h\rangle = Q^n Q^{n-1} \dots Q^2 Q^1 |2\rangle, \quad Q^n \quad (n = a, m, a^*, m^*): \text{supercharge}$$

$ h $	3	$\frac{5}{2}$	2	$\frac{3}{2}$	1	$\frac{1}{2}$	0
			$\underline{1}_{[0]}$	$\underline{10}_{[1]}$	$\underline{45}_{[2]}$	$\underline{120}_{[3]}$	$\underline{210}_{[4]}$
$\underline{d}_{[n]}$	$\underline{1}_{[10]}$	$\underline{10}_{[9]}$	$\underline{45}_{[8]}$	$\underline{120}_{[7]}$	$\underline{210}_{[6]}$	$\underline{252}_{[5]}$	$\underline{210}_{[4]}$

- Wishful assumption of a maximal symmetry breaking (superHiggs-like mechanism, i. e., the maximal number of higher helicity-states become massive by absorbing the lower-helicity states in  $SU(3) \times U(1)$  invariant way as many as possible,

© **Dirac particle survivors** after *tentative* superHiggs-like mechanism

$SU(3)$	$Q_e$	$SU(2) \otimes U(1)$
<u><b>1</b></u>	0 -1 -2	$\begin{pmatrix} \nu_e \\ e \end{pmatrix} \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix} \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix} (N)$ $(E)$
<u><b>3</b></u>	5/3 2/3 -1/3 -4/3	$\begin{pmatrix} u \\ d \end{pmatrix} \begin{pmatrix} c \\ s \end{pmatrix} \begin{pmatrix} t \\ b \end{pmatrix} \begin{pmatrix} h \\ o \end{pmatrix} \begin{pmatrix} a \\ f \end{pmatrix} \begin{pmatrix} g \\ m \end{pmatrix} \begin{pmatrix} r \\ i \\ n \end{pmatrix}$
<u><b>6</b></u>	4/3 1/3 -2/3	$\begin{pmatrix} P \\ Q \\ R \end{pmatrix} \begin{pmatrix} X \\ Y \\ Z \end{pmatrix}$
<u><b>8</b></u>	0 -1	$\begin{pmatrix} N_1 \\ E_1 \end{pmatrix} \begin{pmatrix} N_2 \\ E_2 \end{pmatrix}$

- New color-singlet states:  $N^0$  and  $E^{\pm 2}$

@ How to construct N=10 SUSY with gravity  
despite the **No-Go** theorem for  $N > 8$  by the **S-matrix** argument.

- To circumvent the No-Go theorem the degeneracy of space-time is essential.

## A quick review of NLSUSY:

- Take (flat) space-time specified by  $x^a$  for  $SO(1,3)$  and  $\psi_\alpha$  for  $SL(2,C)$ .
- Consider one form  $\omega^a = dx^a + \frac{\kappa^2}{2i}(\bar{\psi}\gamma^a d\psi - d\bar{\psi}\gamma^a\psi)$ ,  
 $\kappa$  is an **arbitrary** constant with the dimension  $l^{+2}$ .

- $\delta\omega^a = 0$  under  $\delta x^a = \frac{i\kappa^2}{2}(\bar{\zeta}\gamma^a\psi - \bar{\psi}\gamma^a\zeta)$  and  $\delta\psi = \zeta$  with a **global** spinor parameter  $\zeta$ .

- An invariant action( $\sim$  invariant volume) is obtained:

$$S = -\frac{1}{2\kappa^2} \int \omega^0 \wedge \omega^1 \wedge \omega^2 \wedge \omega^3 = \int d^4x L_{VA},$$

$L_{VA}$  is **N=1 Volkov-Akulov model of NLSUSY** given by

$$L_{VA} = -\frac{1}{2\kappa^2}|w_{VA}| = -\frac{1}{2\kappa^2} \left[ 1 - t^a_a + \frac{1}{2}(t^a_a t^b_b - t^a_b t^b_a) + \dots \right],$$

$$|w_{VA}| = \det w^a_b = \det(\delta^a_b + t^a_b), \quad t^a_b = -i\kappa^2(\bar{\psi}\gamma^a\partial_b\psi - \partial_b\bar{\psi}\gamma^a\psi),$$

which is invariant under **N=1 NLSUSY** transformation:

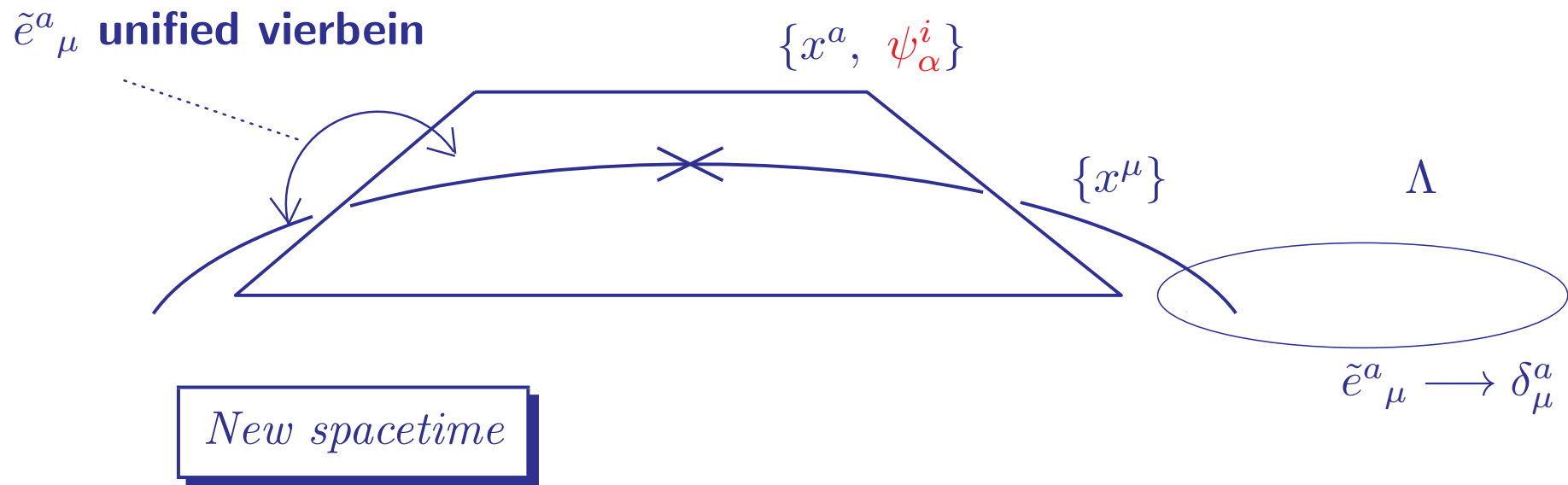
$$\delta_\zeta\psi = \frac{1}{\kappa}\zeta - i\kappa(\bar{\zeta}\gamma^a\psi - \bar{\psi}\gamma^a\zeta)\partial_a\psi, \quad [\delta_1, \delta_2] = \delta_P.$$

- $\psi$  is **Nambu-Goldstone(NG) fermion** for  $\frac{\text{supertranslation}}{\text{translation}}$ .
- $\psi$  is quantized canonically in compatible with **SUSY** algebra.
- Conserved supercurrent, supercharge:

$$J^0 \rightarrow Q = c\psi + (\dots \partial_a\psi \dots)$$

## 2. Nonlinear-supersymmetric general relativity theory(NLSGR)

As an ultimate shape of Nature, considering (unstable) Riemann space-time whose **tangent space possesses NLSUSY structure** specified by the Grassmann coordinates  $\psi_\alpha$  for  $SL(2,C)$  and the ordinary Minkowski coordinates  $x^a$  for  $SO(1,3)$ . we can construct **the unified vierbein**  $\tilde{e}^a_\mu$  which enables the ordinary geometric argument of the general relativity(GR) principle and obtain straightforwardly a new Einstein-Hilbert(EH)-type NLSUSY-invariant general relativity action  $L_{NLSGR}$  equipped with the promising gauge symmetry and the cosmological term  $\Lambda$ .



**@The geometrical arguments of Einstein general relativity (GR) can be extended to new (unstable) space-time.**

• **Unified vierbein  $\tilde{e}^a_\mu(x)$  of new space-time:**

$$\tilde{e}^a_\mu(x) = e^a_\mu + t^a_\mu(\psi) \equiv w^a_b e^b_\mu, \quad w^a_b = \delta^a_b + t^a_b,$$

$$t^a_b(\psi) = \frac{\kappa^2}{2i} (\bar{\psi}^I \gamma^a \partial_b \psi^I - \partial_b \bar{\psi}^I \gamma^a \psi^I), \quad (I = 1, 2, \dots, N)$$

$$\tilde{e}^\mu_a(x) = e^\mu_a - t^\mu_a + t^\mu_\rho t^\rho_a - t^\mu_\sigma t^\sigma_\rho t^\rho_a + t^\mu_\kappa t^\kappa_\sigma t^\sigma_\rho t^\rho_a + \dots + O(t)^4 \equiv e^\mu_b (w^{-1})^b_a,$$

$$\tilde{e}^a_\mu(x) \tilde{e}^\mu_b(x) = \delta^a_b, \quad \tilde{e}^a_\mu(x) \tilde{e}^\nu_a(x) = \delta^\nu_\mu,$$

$$t^a_\mu(\psi) = \frac{\kappa^2}{2i} (\bar{\psi}^I \gamma^a \partial_\mu \psi^I - \partial_\mu \bar{\psi}^I \gamma^a \psi^I), \quad (I = 1, 2, \dots, N)$$

(Note that Grassmann *odd* d.o.f. induces the imaginary part of  $\tilde{e}^a_\mu(x)$  .)



- **$N$ -extended NLSGR action of EH-type for new space-time:**

$$L_{\text{NLSGR}}(\tilde{e}) = -\frac{c^4}{16\pi G}|\tilde{e}|\{R(\tilde{e}) + \Lambda\},$$

$$|\tilde{e}| = \det \tilde{e}^a_\mu = \det(e^a_\mu + t^a_\mu(\psi)),$$

$$t^a_\mu(\psi) = \frac{\kappa^2}{2i}(\bar{\psi}^I \gamma^a \partial_\mu \psi^I - \partial_\mu \bar{\psi}^I \gamma^a \psi^I), (I = 1, 2, \dots, N)$$

- $\tilde{e}^a_\mu(x)$ ,  $\tilde{e}^\mu_a(x)$  : the unified vierbein of new space-time
- $e^a_\mu(x)$  : the ordinary vierbein for the local  $\text{SO}(1,3)$  d.o.f. of GR,
- $t^a_\mu(\psi(x))$  : the mimic vierbein for the local  $\text{SL}(2, \mathbb{C})$  d.o.f. composed of the stress-energy-momentum of NG fermion  $\psi(x)^I$  (called **superons**),
- $R(\tilde{e})$  : Ricci scalar curvature of new space-time computed in terms of  $\tilde{e}^a_\mu(x)$  and  $\tilde{e}^\mu_a(x)$  :
- $\tilde{g}_{\mu\nu} \equiv \tilde{e}^a_\mu \eta_{ab} \tilde{e}^b_\nu$ ,  $\tilde{g}^{\mu\nu}(x) \equiv \tilde{e}^\mu_a(x) \eta^{ab} \tilde{e}^\nu_b(x)$ : metric tensors of new space-time.
- $G$  : the Newton gravitational constant.
- $\Lambda > 0$  : cosmological constant

@NLSGR scenario shows:

- The arbitrary constant  $\kappa^2$  of NLSUSY is fixed at  $\kappa^2 = \left(\frac{c^4 \Lambda}{8\pi G}\right)^{-1}$ ,
- $\Lambda > 0$  in the action  $L_{\text{NLSGR}}$  allows negative dark energy density interpretation of  $\frac{\Lambda}{G}$
- No-go theorem for  $N > 8$  SUSY has been circumvented by the global NLSUSY, i.e. by the degeneracy of the vacuum(flat space).
- Note that  $SO(1, D-1) \cong SL(d, C)$ , i.e.  $\frac{D(D-1)}{2} = 2(d^2 - 1)$  holds for only  $D = 4, d = 2$ .

**NLSGR scenario predicts 4 dimensional space-time.**

- Space-time symmetries ( $\sim sP$ ):

$$[\text{new NLSUSY}] \otimes [\text{local GL}(4, \mathbb{R})] \otimes [\text{local Lorentz}]$$

- Internal symmetries for N-extended NLSUSY GR (N-supers  $\psi^I$  ( $I = 1, 2, \dots, N$ )):

$$[\text{global SO}(N)] \otimes [\text{local U}(1)^N] \otimes [\text{chiral}]. \quad (1)$$

For example:

- Invariance under the new NLSUSY transformation;

$$\delta_\zeta \psi^I = \frac{1}{\kappa} \zeta^I - i\kappa \bar{\zeta}^J \gamma^\rho \psi^J \partial_\rho \psi^I, \quad \delta_\zeta e^a{}_\mu = i\kappa \bar{\zeta}^J \gamma^\rho \psi^J \partial_{[\mu} e^a{}_{\rho]}. \quad (2)$$

induces **GL(4,R) transformations** on  $\tilde{e}^a{}_\mu$  and the unified metric  $\tilde{g}_{\mu\nu}$

$$\delta_\zeta \tilde{e}^a{}_\mu = \xi^\nu \partial_\nu \tilde{e}^a{}_\mu + \partial_\mu \xi^\nu \tilde{e}^a{}_\nu, \quad \delta_\zeta \tilde{g}_{\mu\nu} = \xi^\kappa \partial_\kappa \tilde{g}_{\mu\nu} + \partial_\mu \xi^\kappa \tilde{g}_{\kappa\nu} + \partial_\nu \xi^\kappa \tilde{g}_{\mu\kappa},$$

where  $\zeta$  is a constant spinor parameter,  $\partial_{[\rho} e^a{}_{\mu]} = \partial_\rho e^a{}_\mu - \partial_\mu e^a{}_\rho$  and  $\xi^\rho = -i\kappa \bar{\zeta}^I \gamma^\rho \psi^I$ .

- Commutators of two new NLSUSY transformation on  $\psi^I$  and  $e^a{}_\mu$  close on **GL(4,R)**,

$$[\delta_{\zeta_1}, \delta_{\zeta_2}] \psi^I = \Xi^\mu \partial_\mu \psi^I, \quad [\delta_{\zeta_1}, \delta_{\zeta_2}] e^a{}_\mu = \Xi^\rho \partial_\rho e^a{}_\mu + e^a{}_\rho \partial_\mu \Xi^\rho, \quad (3)$$

where  $\Xi^\mu = 2i\bar{\zeta}_1^I \gamma^\mu \zeta_2^I - \xi_1^\rho \xi_2^\sigma e_a{}^\mu \partial_{[\rho} e^a{}_{\sigma]}$ .

- New NLSUSY is the square-root of GL(4,R);

$$[\delta_1, \delta_2] = \delta_{\text{GL}(4,\text{R})}, \quad i.e. \quad \delta \sim \sqrt{\delta_{\text{GL}(4,\text{R})}}.$$

- Invariance under new local Lorentz transformation;

$$\delta_L \psi^I = -\frac{i}{2} \epsilon_{ab} \sigma^{ab} \psi^I, \quad \delta_L e^a{}_\mu = \epsilon^a{}_b e^b{}_\mu + \frac{\kappa^4}{4} \varepsilon^{abcd} \bar{\psi}^I \gamma_5 \gamma_d \psi^I$$

induce the familiar form of the local Lorentz transformation on  $\tilde{e}^a{}_\mu$ :

$$\delta_L \tilde{e}^a{}_\mu = \epsilon^a{}_b \tilde{e}^b{}_\mu$$

with the local parameter  $\epsilon_{ab} = (1/2) \epsilon_{[ab]}(x)$   
and close the algebra, e. g., **the new form on**  $e^a{}_\mu(x)$

$$[\delta_{L_1}, \delta_{L_2}] e^a{}_\mu = \beta^a{}_b e^b{}_\mu + \frac{\kappa^4}{4} \varepsilon^{abcd} \bar{\psi}^j \gamma_5 \gamma_d \psi^j (\partial_\mu \beta_{bc}),$$

where  $\beta_{ab} = -\beta_{ba}$  is given by  $\beta_{ab} = \epsilon_{2ac} \epsilon_1^c{}_b - \epsilon_{2bc} \epsilon_1^c{}_a$ .

- The ordinary local  $GL(4, \mathbb{R})$  invariance is manifest by the construction.

### 3. Evolution of NLSGR [Cosmology]: Big Collapse and SGM Model

@ Due to the NLSUSY structure of space-time,  $L_{NLSGR}$  would collapse to (called **Big Collapse(BC)**) spontaneously to the vacuum configuration.

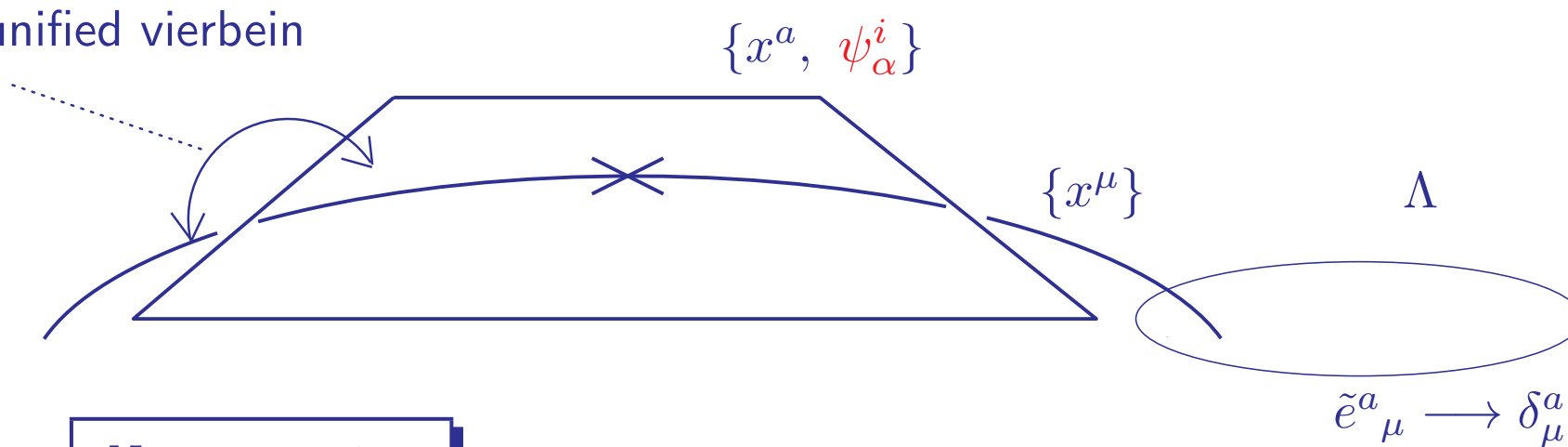
i. e.,

BC to ordinary EH action of graviton  $e^a{}_\mu$ ,  
NLSUSY action of Nambu-Goldstone(NG) fermion  $\psi$  (called **superon**)  
and their gravitational interaction ( called **superon-graviton model(SGM)**) .

$$L_{SGM}(e, \psi) = -\frac{c^4}{16\pi G}|e|\{R(e) + |w(\psi^I)_{VA}|\Lambda + \tilde{T}(e, \psi^I)\}.$$

SGM action  $L_{SGM}(e, \psi)$  describes the vacuum configuration of NLSGR  
for ultimate shape of Nature.

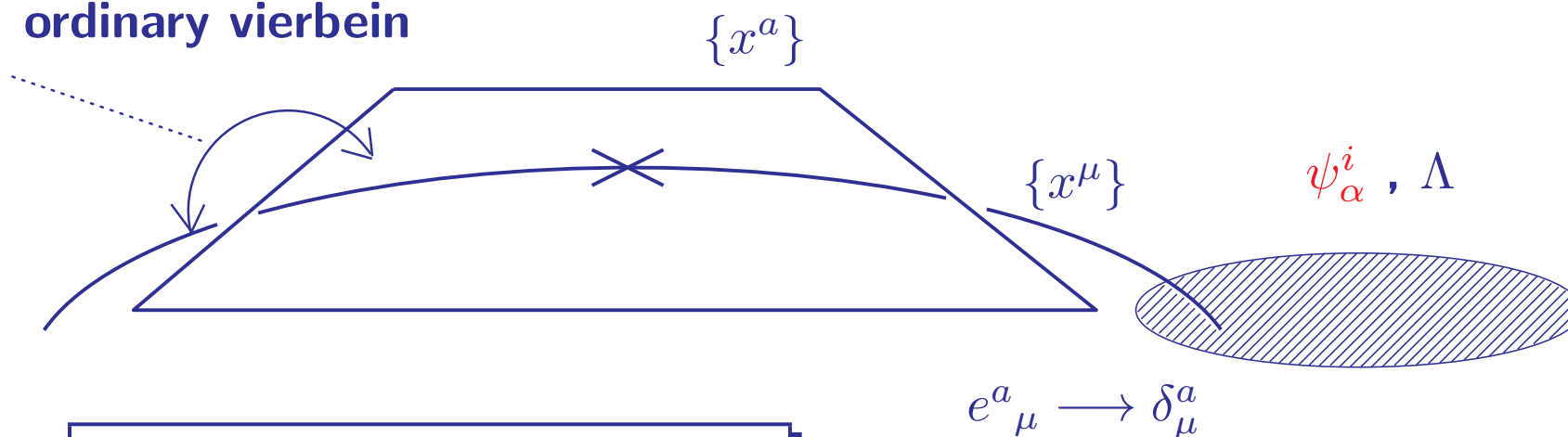
$\tilde{e}^a{}_\mu$  unified vierbein



*New spacetime*

$\Downarrow$  **Big Collapse**

$e^a{}_\mu$  : ordinary vierbein



*Riemann spacetime*  $\oplus$  **matter**

- The variation with respect to  $e^a{}_\mu$  produces Einstein equation of SGM space-time equipting with **negative dark energy density** :

$$R_{\mu\nu}(e) - \frac{1}{2}g_{\mu\nu}R(e) = \frac{8\pi G}{c^4}\{\tilde{T}_{\mu\nu}(e, \psi) - g_{\mu\nu}\frac{c^4\Lambda}{8\pi G}\}. \quad (4)$$

- **Big collapse(BC)** induces 3 dimensional (space-like region) rapid expansion of space-time by Pauli principle(**Quantum Inflation b/a BB**):

$$ds^2 = s_{\mu\nu}(x)dx^\mu dx^\nu = \{g_{\mu\nu} + \underline{\Phi_{\mu\nu}(e, \psi)}\}dx^\mu dx^\nu.$$

$$\begin{aligned} \{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = 0 &\Rightarrow \{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = \delta^{(3)}(\mathbf{x} - \mathbf{y})\delta_{ab} \\ (\{\psi(x)_a, \bar{\psi}(y)_b\} = 0 &\Rightarrow \{\psi(x)_a, \bar{\psi}(y)_b\} = iS(x - y)\delta_{ab} ) \end{aligned}$$

- The conserved supercurrent  $S^{I\mu} = i\sqrt{\frac{c^4\Lambda}{8\pi G}}\gamma^\mu\psi^I + (\cdots\partial\psi, \psi\cdots)$ .

couples the superon(NG fermion) to the vacuum

$$\langle \psi_\beta^J | S_\alpha^{I\mu} | 0 \rangle = i\sqrt{\frac{c^4\Lambda}{8\pi G}}\delta^{IJ}(\gamma^\mu)_{\alpha\beta} \quad \text{with the strength } g_{sv} \equiv \sqrt{\frac{c^4\Lambda}{8\pi G}} = \kappa^{-1},$$

a new universal fundamental constant in Nature characterizing Big Collapse.

**How to extract the (low energy) particle physics from SGM.**

## 4. Linearization of NLSUSY and the NL/L SUSY relation (Equivalence)

@ BC is the ignition of BB and the evolution of the universe:

At BC the universal attractive force graviton would produce all possible gravitational-composites of superons, which span the (massless) irreducible representation of the **LSUSY  $SO(N)$  super-Poincaré(sP) algebra** of nature.

- This can be regarded as **the ignition of the Big Bang** of the universe.
- The vacuum structure of SGM is studied by linearizing NLSUSY in SGM and creating the equivalent **global LSUSY** theory in terms of the **superon-composite global SUSY supermultiplet** (called **NL/L SUSY relation**).
- The **global (NL and L) SUSY algebra** plays an essential role.
- How to find NL/L SUSY relation;: We can adopt ,

**Commutator based linearization and/or Superfield based one.**



As an example of NL/L SUSY relation:

- NL/L SUSY relation(**equivalence**) for N=2 SUSY in  $d = 2$  flat space;

$$L_{\text{N=2LSUSYQED}}(v_a, \lambda^i, A, \phi^i D, \chi, F^i, \dots) \\ = f(\xi, \xi^i) L_{\text{N=2NLSUSY}}(\psi^j) + [\text{suface terms}],$$

$$v^a = -\frac{i}{2} \xi \kappa \epsilon^{ij} \bar{\psi}^i \gamma^a \psi^j |w|, \quad \lambda^i = \xi \psi^i |w|, \quad A = \frac{1}{2} \xi \kappa \bar{\psi}^i \psi^i |w|, \quad D = \frac{\xi}{\kappa} |w| \quad \text{etc.} \\ \chi = \xi^i \left[ \psi^i |w| + \frac{i}{2} \kappa^2 \partial_a \{ \gamma^a \psi^i \bar{\psi}^j \psi^j |w| \} \right], \dots$$

**SUSY compositeness**

- NL/L SUSY relation bridges **the cosmology and the low energy particle physics in NLSGR** scenario. e.g.,  $\text{N=2SUSYQED} : e = \tilde{f}(\xi, \xi^i, G, \Lambda)$

**Broken LSUSY(QED) gauge theory is encoded  
in the vacuum of NLSUSY theory as composites of NG fermion.**

- For linearizing /analyzing the vacuum of NLSGR/SGM  
we need **broken global LSUSY** SUGRA-analogue model which realizes:

**NL/L SUSY relation (*equivalence*):**

$$L_{\text{NLSGR}}(w) = L_{\text{SGM}}(e, \psi) = L_{\text{LSUSY}}(e^a{}_\mu, \psi_\nu(e, \psi), M(e, \psi), N(e, \psi), \dots).$$

## 5. Cosmological implications of SGM scenario

The variation of SGM action  $L_{\text{SGM}}(e, \psi)$  with respect to  $e^a{}_\mu$  yields Einstein equation equipping with matter and cosmological term:

$$R_{\mu\nu}(e) - \frac{1}{2}g_{\mu\nu}R(e) = \frac{8\pi G}{c^4}\{\tilde{T}_{\mu\nu}(e, \psi) - g_{\mu\nu}\frac{c^4\Lambda}{8\pi G}\}. \quad (5)$$

where  $\tilde{T}_{\mu\nu}(e, \psi)$  abbreviates the stress-energy-momentum of superon(NG fermion) including the gravitational interaction.

- Note that the cosmological term  $-\frac{c^4\Lambda}{8\pi G}$  can be interpreted as the negative energy density of space-time, i.e. the dark energy density  $\rho_D$ .

- Big collapse(BC) induces 3 dimensional (space-like region) rapid expansion of space-time by Pauli principle:

$$ds^2 = s_{\mu\nu}(x)dx^\mu dx^\nu = \{g_{\mu\nu} + \underline{\Phi_{\mu\nu}(e, \psi)}\}dx^\mu dx^\nu.$$

$$\{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = 0 \quad \Rightarrow \quad \{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = \delta^{(3)}(\mathbf{x} - \mathbf{y})\delta_{ab}$$

$$( \{ \psi(x)_a, \bar{\psi}(y)_b \} = 0 \quad \Rightarrow \quad \{ \psi(x)_a, \bar{\psi}(y)_b \} = iS(x - y)\delta_{ab} )$$

- BC produces gravitational composite (massless) eigenstates of SO(N) sP algebra, which is the ignition of the Big Bang(BB) SM scenario.

- As shown in the toy model, the vacuum of SGM may explain naturally observed mysterious (numerical) relations:

$$\text{dark energy density } \rho_D \sim O(\kappa^{-2}) \sim m_\nu^4 \sim (10^{-12}GeV)^4 \sim g_{sv}^2,$$

provided  $\lambda_D^0$  is identified with neutrino and  $f\xi \sim O(1)$ .

## 6. Summary

### **NLSGR(SGM)** for unity of nature:

- Ultimate entity; **Unstable**  $d = 4$  **space-time**:  $[x^a, \psi_\alpha^N; x^\mu]$  described by  $[L_{\text{NLSGR}}(w^a_\mu)]$  : Global SUSY NLSGR on New space-time with  $\Lambda > 0$

$\Rightarrow$  **Big Collapse (BC)**

- The creation of ordinary Riemann space-time [graviton  $e^a_\mu$ ] and massless NG fermionic matter [superon  $\psi_\alpha^N$ ], **rapid expansion by Pauli principle**  $[L_{\text{SGM}} = L_{\text{EH}}(e) - \Lambda + T(\psi.e)]$  : **Einstein GR with  $\Lambda > 0$  and  $N$  superon**

- The universal attractive force graviton dictates the evolution of  $[L_{\text{SGM}}$  by forming gravitational composite **global** LSUSY supermultiplet corresponding to (massless) eigenstates of space-time symmetry  $\text{SO}(10)$  sP.

$\Rightarrow$  **Ignition of Big Bang of the universe toward  $\Lambda$ CDM scenario**

- In flat space-time, **broken  $N$ -LSUSY theory** emerges from the  **$N$ -NLSUSY cosmological term of  $L_{\text{SGM}}(e, \psi)$**  [NL/L SUSY relation].  $\longleftrightarrow$  BCS vs GL

**The cosmological term is the origin of everything!**

- NLSGR((SGM) paradigm can bridge naturally the cosmology and the low energy particle physics, which provides new insights into unsolved problems of cosmology and particle physics,, e.g. the space-time dimension *four*, the origin of mass and SUSY breaking, the dark energy and the dark matter, the dark energy density  $\approx (\text{neutrino mass})^4$  the tiny neutrino mass, inflation of space-time, the three-generations structure of quarks and leptons, the rapid expansion of space-time, the magnitude of the bare gauge coupling constant, matter dominance of nature , **the shape and the role of black hole** etc.  
 $NLSGR \rightarrow (BC) \rightarrow SGM \rightarrow (BB) \rightarrow SMs \rightarrow \cdots \rightarrow BH \rightarrow NLSGR \rightarrow \cdots$
- Many important problems are yet to be studied, e.g., global SUSY for higher spins (5/2 3) and SUGRA
- Direct linearization of NLSGR/SGM
- **The revisit of global SUSY superfield fomulation from SGM/SQM. viewpoint.**  $\leftrightarrow$  C-asymmetry, proton decay, etc.

For some details:

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- 3, K. Shima、 Preliminary temporal report: arXiv:2012.01646[hep-th](2020)