

Nonlinear-Supersymmetric General Relativity Theory(NLSGR)

Kazunari Shima
Saitama Institute of Technology

OUTLINE

1. Motivation
2. Nonlinear-supersymmetric general relativity theory(NLSGR)
3. Evolution of NLSGR (Big Collapse and Superon-Graviton Model)
4. Linearization of NLSUSY and the vacuum of $L_{SGM}(e, \psi)$ (NL/L SUSY relation)
5. Cosmological Implications of NLSGR.
6. Summary

1. Motivations

① The success of **GR** and **SM(GWS)** model,
still many unsolved fundamental problems are left, e.g.,

- Unification of two SMs,
- Space-time dimension *four*,
- Three-generations structure of quarks and leptons,
- C-asymmetry of nature
- Tiny Neutrino mass M_ν , proton stability in GUT
- Dark Matter, Dark Energy; $\rho_{D.E.} \sim (M_\nu)^4 \Leftrightarrow \Lambda CDM$, Inflation
⇒ **SUSY, SUGRA**,
- Origin of SUSY breaking ?, etc.

② As for the three-generations structure

among $SO(N)$ LSUSY sP, only $N = 10$ gives SM with just 3 generations:,

where

- The decomposition of 10 supercharges Q^I , ($I = 1, 2, \dots, 10$) are:

$$\underline{\mathbf{10}}_{SO(10)} = \underline{\mathbf{5}}_{SU(5)} + \underline{\mathbf{5}}^*_{SU(5)}$$

$$\underline{\mathbf{5}}_{SU(5)} = [\{\underline{\mathbf{3}}^{*c}, \underline{\mathbf{1}}^{ew}, (\frac{e}{3}, \frac{e}{3}, \frac{e}{3}) : Q_a (a = 1, 2, 3)\}, \{\underline{\mathbf{1}}^c, \underline{\mathbf{2}}^{ew}, (0, -e) : Q_m (m = 4, 5)\}].$$

⇒ **Supercharge** $\underline{\mathbf{5}}_{SU(5)}$ has the same quantum numbers as $\underline{\mathbf{5}}$ of $SU(5)$ GUT.

- Massless helicity states of **gravity supermultiplet** of $SO(10)$ sP with CPT conjugation are specified by the helicity $h = (2 - \frac{n}{2})$ and the dimension $\underline{d}_{[n]} = \frac{10!}{n!(10-n)!}$:

$|h\rangle = Q^n Q^{n-1} \dots Q^2 Q^1 |2\rangle$, Q^n ($n = a, m, a^*, m^*$): supercharge

$ h $	3	$\frac{5}{2}$	2	$\frac{3}{2}$	1	$\frac{1}{2}$	0
$\underline{d}_{[n]}$	$\underline{1}_{[10]}$	$\underline{10}_{[9]}$	$\underline{1}_{[0]}$	$\underline{10}_{[1]}$	$\underline{45}_{[2]}$	$\underline{120}_{[3]}$	$\underline{210}_{[4]}$
			$\underline{45}_{[8]}$	$\underline{120}_{[7]}$	$\underline{210}_{[6]}$	$\underline{252}_{[5]}$	$\underline{210}_{[4]}$

- Wishful assumption of a maximal symmetry breaking (superHiggs-like mechanism, i. e., the maximal number of higher helicity-states become massive by absorbing the lower-helicity states in $SU(3) \times U(1)$ invariant way as many as possible,

© Dirac particle survivors after *tentative superHiggs-like mechanism*

$SU(3)$	Q_e	$SU(2) \otimes U(1)$
$\underline{\mathbf{1}}$	0	$\begin{pmatrix} \nu_e \\ e \end{pmatrix} \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix} \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix} (N)$
	-1	
	-2	(E)
$\underline{\mathbf{3}}$	$5/3$	
	$2/3$	
	$-1/3$	$\begin{pmatrix} u \\ d \end{pmatrix} \begin{pmatrix} c \\ s \end{pmatrix} \begin{pmatrix} t \\ b \end{pmatrix} \begin{pmatrix} h \\ o \end{pmatrix} \begin{pmatrix} a \\ f \end{pmatrix} \begin{pmatrix} g \\ m \end{pmatrix} \begin{pmatrix} r \\ i \end{pmatrix} (n)$
	$-4/3$	
$\underline{\mathbf{6}}$	$4/3$	
	$1/3$	
	$-2/3$	$\begin{pmatrix} P \\ Q \\ R \end{pmatrix} \begin{pmatrix} X \\ Y \\ Z \end{pmatrix}$
$\underline{\mathbf{8}}$	0	
	-1	$\begin{pmatrix} N_1 \\ E_1 \end{pmatrix} \begin{pmatrix} N_2 \\ E_2 \end{pmatrix}$

- New color-singlet states: N^0 and $E^{\pm 2}$

@ How to construct $N=10$ SUSY with gravity
despite the **No-Go** theorem for $N > 8$ by the **S-matrix** argument.

- To circumvent the No-Go theorem the degeneracy of space-time is essential.

A quick review of NLSUSY:

- Take (flat) space-time specified by x^a for $\text{SO}(1,3)$ and ψ_α for $\text{SL}(2,\mathbb{C})$.
- Consider one form $\omega^a = dx^a + \frac{\kappa^2}{2i}(\bar{\psi}\gamma^a d\psi - d\bar{\psi}\gamma^a \psi)$,
 κ is an arbitrary constant with the dimension l^{+2} .
- $\delta\omega^a = 0$ under $\delta x^a = \frac{i\kappa^2}{2}(\bar{\zeta}\gamma^a \psi - \bar{\psi}\gamma^a \zeta)$ and $\delta\psi = \zeta$
with a global spinor parameter ζ .
- An invariant action(\sim invariant volume) is obtained:

$$S = -\frac{1}{2\kappa^2} \int \omega^0 \wedge \omega^1 \wedge \omega^2 \wedge \omega^3 = \int d^4x L_{VA},$$

L_{VA} is N=1 Volkov-Akulov model of NLSUSY given by

$$L_{VA} = -\frac{1}{2\kappa^2} |w_{VA}| = -\frac{1}{2\kappa^2} [1 - t^a{}_a + \frac{1}{2}(t^a{}_a t^b{}_b - t^a{}_b t^b{}_a) + \dots],$$

$$|w_{VA}| = \det w^a{}_b = \det(\delta^a_b + t^a{}_b), \quad t^a{}_b = -i\kappa^2(\bar{\psi}\gamma^a \partial_b \psi - \partial_b \bar{\psi}\gamma^a \psi),$$

which is invariant under N=1 NLSUSY transformation:

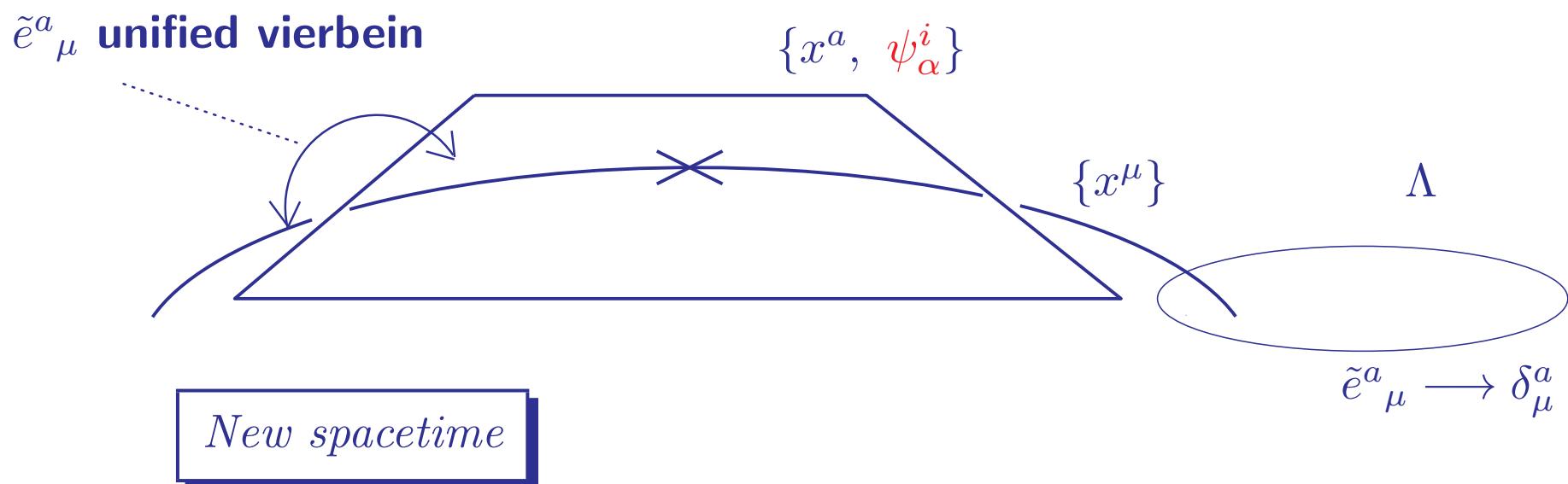
$$\delta_\zeta \psi = \frac{1}{\kappa} \zeta - i\kappa(\bar{\zeta}\gamma^a \psi - \bar{\psi}\gamma^a \zeta) \partial_a \psi, \quad [\delta_1, \delta_2] = \delta_{\mathbf{P}}.$$

- ψ is Nambu-Goldstone(NG) fermion for $\frac{\text{supertranslation}}{\text{translation}}$.
- ψ is quantized canonically in compatible with SUSY algebra.
- Conserved supercurrent, supercharge:

$$J^0 \rightarrow Q = c\psi + (\dots \partial_a \psi \dots)$$

2. Nonlinear-supersymmetric general relativity theory(NLSGR)

As an ultimate shape of Nature, considering (unstable) Riemann space-time whose tangent space possesses **NLSUSY structure** specified by the Grassmann coordinates ψ_α for $SL(2, \mathbb{C})$ and the ordinary Minkowski coordinates x^a for $SO(1, 3)$. we can construct the **unified vierbein** \tilde{e}_μ^a which enables the ordinary geometric argument of the general relativity(GR) principle and obtain straightforwardly a new Einstein-Hilbert(EH)-type NLSUSY-invariant general relativity action L_{NLSGR} equipped with the promising gauge symmetry and the cosmological term Λ .



@The geometrical arguments of Einstein general relativity(GR) can be extended to new (unstable) space-time.

- Unified vierbein $\tilde{e}^a{}_\mu(x)$ of new space-time:

$$\tilde{e}^a{}_\mu(x) = e^a{}_\mu + t^a{}_\mu(\psi) \equiv w^a{}_b e^b{}_\mu, \quad w^a{}_b = \delta^a{}_b + t^a{}_b,$$

$$t^a{}_b(\psi) = \frac{\kappa^2}{2i} (\bar{\psi}^I \gamma^a \partial_b \psi^I - \partial_b \bar{\psi}^I \gamma^a \psi^I), \quad (I = 1, 2, \dots, N)$$

$$\tilde{e}^\mu{}_a(x) = e^\mu{}_a - t^\mu{}_a + t^\mu{}_\rho t^\rho{}_a - t^\mu{}_\sigma t^\sigma{}_\rho t^\rho{}_a + t^\mu{}_\kappa t^\kappa{}_\sigma t^\sigma{}_\rho t^\rho{}_a + \dots + O(t)^4 \equiv e^\mu{}_b (w^{-1})^b{}_a,$$

$$\tilde{e}^a{}_\mu(x) \tilde{e}^\mu{}_b(x) = \delta^a{}_b, \quad \tilde{e}^a{}_\mu(x) \tilde{e}^\nu{}_a(x) = \delta^\nu{}_\mu,$$

$$t^a{}_\mu(\psi) = \frac{\kappa^2}{2i} (\bar{\psi}^I \gamma^a \partial_\mu \psi^I - \partial_\mu \bar{\psi}^I \gamma^a \psi^I), \quad (I = 1, 2, \dots, N)$$

(Note that Grassmann *odd* d.o.f. induces the imaginary part of $\tilde{e}^a{}_\mu(x)$.)

- ***N*-extended NLSGR action of EH-type for new space-time:**

$$L_{\text{NLSGR}}(\tilde{e}) = -\frac{c^4}{16\pi G} |\tilde{e}| \{R(\tilde{e}) + \Lambda\},$$

$$|\tilde{e}| = \det \tilde{e}_\mu^a = \det(e^a{}_\mu + t^a{}_\mu(\psi)),$$

$$t^a{}_\mu(\psi) = \frac{\kappa^2}{2i} (\bar{\psi}^I \gamma^a \partial_\mu \psi^I - \partial_\mu \bar{\psi}^I \gamma^a \psi^I), (I = 1, 2, \dots, N)$$

- $\tilde{e}^a{}_\mu(x)$, $\tilde{e}^\mu{}_a(x)$: **the unified vierbein of new space-time**
- $e^a{}_\mu(x)$: **the ordinary vierbein for the local $\text{SO}(1,3)$ d.o.f. of GR,**
- $t^a{}_\mu(\psi(x))$: **the mimic vierbein for the local $\text{SL}(2,\mathbb{C})$ d.o.f. composed of the stress-energy-momentum of NG fermion $\psi(x)^I$ (called **superons**),**
- $R(\tilde{e})$: **Ricci scalar curvature of new space-time**
computed in terms of $\tilde{e}^a{}_\mu(x)$ and $\tilde{e}^\mu{}_a(x)$:
- $\tilde{g}_{\mu\nu} \equiv \tilde{e}^a{}_\mu \eta_{ab} \tilde{e}^b{}_\nu$, $\tilde{g}^{\mu\nu}(x) \equiv \tilde{e}^\mu{}_a(x) \eta^{ab} \tilde{e}^\nu{}_a(x)$: **metric tensors of new space-time.**
- G : **the Newton gravitational constant.**
- $\Lambda > 0$: **cosmological constant**

@NLSGR scenario shows:

- The arbitrary constant κ^2 of NLSUSY is fixed at $\kappa^2 = (\frac{c^4 \Lambda}{8\pi G})^{-1}$,
- $\Lambda > 0$ in the action L_{NLSGR} allows negative dark energy density interpretation of $\frac{\Lambda}{G}$
- No-go theorem for $N > 8$ SUSY has been circumvented by the global NLSUSY, i.e. by the degeneracy of the vacuum (flat space).
- Note that $SO(1, D - 1) \cong SL(d, C)$, i.e. $\frac{D(D-1)}{2} = 2(d^2 - 1)$ holds for only $D = 4, d = 2$.

NLSGR scenario predicts 4 dimensional space-time.

- Space-time symmetries ($\sim sP$):

[new NLSUSY] \otimes [local GL(4, R)] \otimes [local Lorentz]

- Internal symmetries for N-extended NLSUSY GR (N-superons ψ^I ($I = 1, 2, \dots, N$)):

[global SO(N)] \otimes [local U(1)^N] \otimes [chiral]. (1)

For example:

- Invariance under the new NLSUSY transformation;

$$\delta_\zeta \psi^I = \frac{1}{\kappa} \zeta^I - i\kappa \bar{\zeta}^J \gamma^\rho \psi^J \partial_\rho \psi^I, \quad \delta_\zeta e^a{}_\mu = i\kappa \bar{\zeta}^J \gamma^\rho \psi^J \partial_{[\mu} e^a{}_{\rho]}. \quad (2)$$

induces **GL(4,R) transformations** on \tilde{e}_μ^a and the unified metric $\tilde{g}_{\mu\nu}$

$$\delta_\zeta \tilde{e}_\mu^a = \xi^\nu \partial_\nu \tilde{e}_\mu^a + \partial_\mu \xi^\nu \tilde{e}_\nu^a, \quad \delta_\zeta \tilde{g}_{\mu\nu} = \xi^\kappa \partial_\kappa \tilde{g}_{\mu\nu} + \partial_\mu \xi^\kappa \tilde{g}_{\kappa\nu} + \partial_\nu \xi^\kappa \tilde{g}_{\mu\kappa},$$

where ζ is a constant spinor parameter, $\partial_{[\rho} e^a{}_{\mu]} = \partial_\rho e^a{}_\mu - \partial_\mu e^a{}_\rho$ and $\xi^\rho = -i\kappa \bar{\zeta}^I \gamma^\rho \psi^I$.

- Commutators of two new NLSUSY transformation on ψ^I and $e^a{}_\mu$ close on **GL(4,R)**,

$$[\delta_{\zeta_1}, \delta_{\zeta_2}] \psi^I = \Xi^\mu \partial_\mu \psi^I, \quad [\delta_{\zeta_1}, \delta_{\zeta_2}] e^a{}_\mu = \Xi^\rho \partial_\rho e^a{}_\mu + e^a{}_\rho \partial_\mu \Xi^\rho, \quad (3)$$

where $\Xi^\mu = 2i\bar{\zeta}^I{}_1 \gamma^\mu \zeta^I{}_2 - \xi_1^\rho \xi_2^\sigma e_a{}^\mu \partial_{[\rho} e^a{}_{\sigma]}$.

- **New NLSUSY is the square-root of GL(4,R);**

$$[\delta_1, \delta_2] = \delta_{\text{GL}(4,\text{R})}, \quad \text{i.e.} \quad \delta \sim \sqrt{\delta_{\text{GL}(4,\text{R})}}.$$

- Invariance under new local Lorentz transformation;

$$\delta_L \psi^I = -\frac{i}{2} \epsilon_{ab} \sigma^{ab} \psi^I, \quad \delta_L e^a{}_\mu = \epsilon^a{}_b e^b{}_\mu + \frac{\kappa^4}{4} \varepsilon^{abcd} \bar{\psi}^I \gamma_5 \gamma_d \psi^I$$

induce the familiaform form of the local Lorentz transformation on $\tilde{e}^a{}_\mu$:

$$\delta_L \tilde{e}^a{}_\mu = \epsilon^a{}_b \tilde{e}^b{}_\mu$$

with the local parameter $\epsilon_{ab} = (1/2) \epsilon_{[ab]}(x)$
and close the algebra, e. g., the new form on $e^a{}_\mu(x)$

$$[\delta_{L_1}, \delta_{L_2}] e^a{}_\mu = \beta^a{}_b e^b{}_\mu + \frac{\kappa^4}{4} \varepsilon^{abcd} \bar{\psi}^j \gamma_5 \gamma_d \psi^j (\partial_\mu \beta_{bc}),$$

where $\beta_{ab} = -\beta_{ba}$ is given by $\beta_{ab} = \epsilon_{2ac} \epsilon_1{}^c{}_b - \epsilon_{2bc} \epsilon_1{}^c{}_a$.

- The ordinary local $GL(4, \mathbb{R})$ invariance is manifest by the construction.

3. Evolution of NLSGR [Cosmology]: Big Collapse and SGM Model

⑥ Due to the NLSUSY structure of space-time, L_{NLSGR} would collapse to (called **Big Collapse(BC)**) spontaneously to the vacuum configuration.

i. e.,

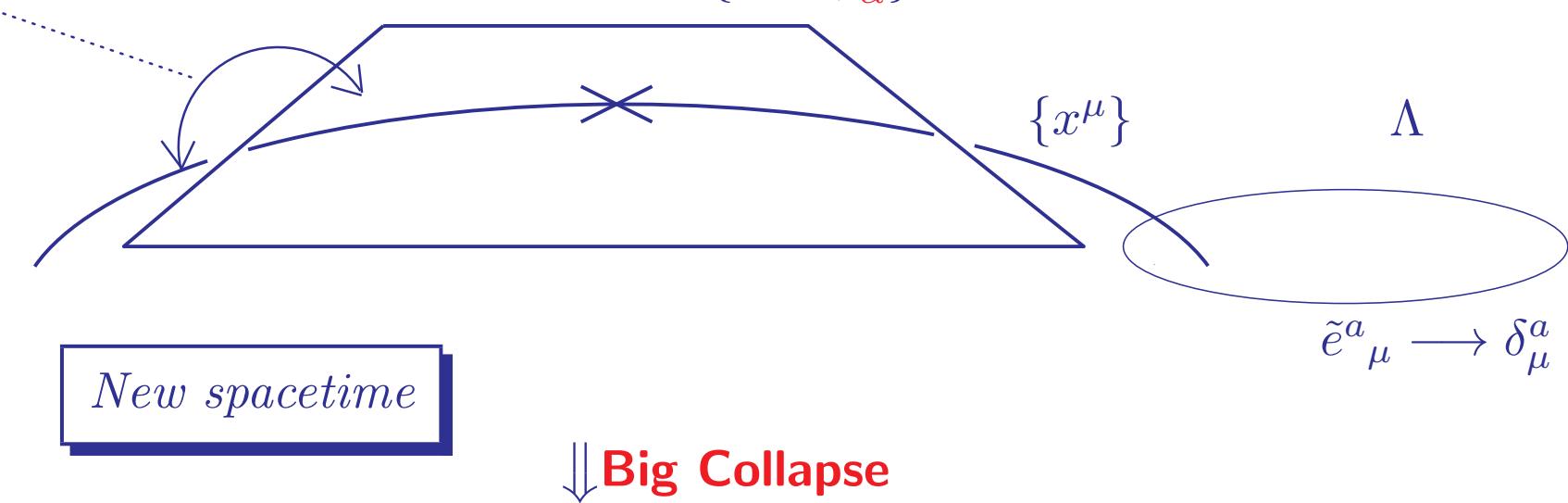
BC to ordinary EH action of graviton e^a_μ ,
NLSUSY action of Nambu-Goldstone(NG) fermion ψ (called superon)
and their gravitational interaction (called superon-graviton model(SGM)) .

$$L_{SGM}(e, \psi) = -\frac{c^4}{16\pi G} |e| \{ R(e) + |w(\psi^I)_{VA}| \Lambda + \tilde{T}(e, \psi^I) \}.$$

SGM action $L_{SGM}(e, \psi)$ describes the vacuum configuration of NLSGR
for ultimate shape of Nature.

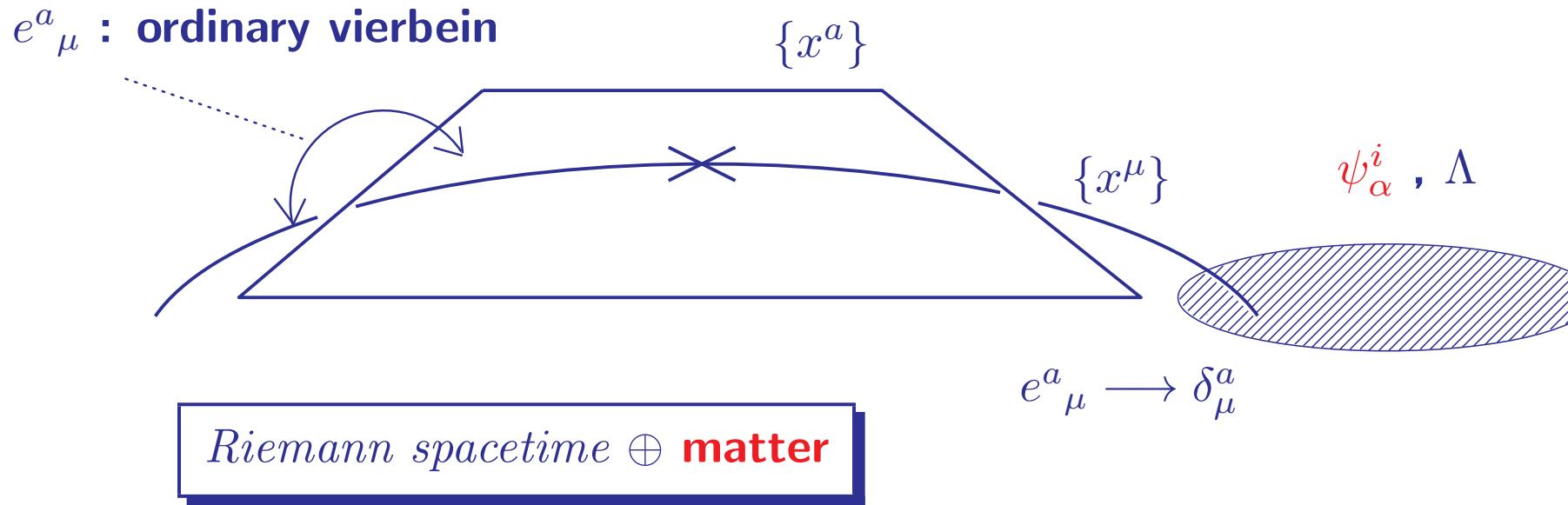
\tilde{e}^a_μ unified vierbein

$\{x^a, \psi_\alpha^i\}$



e^a_μ : ordinary vierbein

$\{x^a\}$



- The variation with respect to e^a_μ produces Einstein equation of SGM space-time equipting with **negative dark energy density** :

$$R_{\mu\nu}(e) - \frac{1}{2}g_{\mu\nu}R(e) = \frac{8\pi G}{c^4}\{\tilde{T}_{\mu\nu}(e, \psi) - g_{\mu\nu}\frac{c^4\Lambda}{8\pi G}\}. \quad (4)$$

- **Big collapse(BC)** induces 3 dimensional (space-like region) rapid expansion of space-time by Pauli principle(**Quantum Inflation b/a BB**):

$$ds^2 = s_{\mu\nu}(x)dx^\mu dx^\nu = \{g_{\mu\nu} + \underline{\Phi_{\mu\nu}(e, \psi)}\}dx^\mu dx^\nu.$$

$$\{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = 0 \quad \Rightarrow \quad \{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = \delta^{(3)}(\mathbf{x} - \mathbf{y})\delta_{ab}$$

$$(\{\psi(x)_a, \bar{\psi}(y)_b\} = 0 \quad \Rightarrow \quad \{\psi(x)_a, \bar{\psi}(y)_b\} = iS(x - y)\delta_{ab})$$

- The conserved supercurrent $S^{I\mu} = i\sqrt{\frac{c^4\Lambda}{8\pi G}}\gamma^\mu\psi^I + (\cdots \partial\psi, \psi \cdots)$.

couples the superon(NG fermion) to the vacuum

$\langle\psi_\beta^J|S_\alpha^{I\mu}|0\rangle = i\sqrt{\frac{c^4\Lambda}{8\pi G}}\delta^{IJ}(\gamma^\mu)_{\alpha\beta}$ with the strength $g_{sv} \equiv \sqrt{\frac{c^4\Lambda}{8\pi G}} = \kappa^{-1}$,
a new universal fundamental constant in Nature characterizing Big Collapse.

How to extract the (low energy) particle physics from SGM.

4. Linearization of NLSUSY and the NL/L SUSY relation (Equivalence)

@ BC is the ignition of BB and the evolution of the universe:

At BC the universal attractive force graviton would produce all possible gravitational-composites of superons, which span the (massless) irreducible representation of the LSUSY $SO(N)$ super-Poincaré(sP) algebra of nature.

- This can be regarded as the ignition of the Big Bang of the universe.
- The vacuum structure of SGM is studied by linearizing NLSUSY in SGM and creating the equivalent global LSUSY theory in terms of the superon-composite global SUSY supermultiplet (called NL/L SUSY relation).
- The global (NL and L) SUSY algebra plays an essential role.
- How to find NL/L SUSY relation;: We can adopt ,
Commutator based linearization and/or Superfield based one.

As an example of NL/L SUSY relation:

- NL/L SUSY relation(**equivalence**) for N=2 SUSY in $d = 2$ flat space;

$$L_{\text{N=2LSUSYQED}}(v_a, \lambda^i, A, \phi^i D, \chi, F^i, \dots) \\ = f(\xi, \xi^i) L_{\text{N=2NLSUSY}}(\psi^j) + [\text{surface terms}],$$

$$v^a = -\frac{i}{2}\xi\kappa\epsilon^{ij}\bar{\psi}^i\gamma^a\psi^j|w|, \lambda^i = \xi\psi^i|w|, A = \frac{1}{2}\xi\kappa\bar{\psi}^i\psi^i|w|, D = \frac{\xi}{\kappa}|w| \quad \text{etc.} \\ \chi = \xi^i [\psi^i|w| + \frac{i}{2}\kappa^2\partial_a\{\gamma^a\psi^i\bar{\psi}^j\psi^j|w|\}], \dots$$

SUSY compositeness

- NL/L SUSY relation bridges the cosmology and the low energy particle physics in NLSGR scenario. e.g., N=2SUSYQED : $e = \tilde{f}(\xi, \xi^i, G, \Lambda)$

**Broken LSUSY(QED) gauge theory is encoded
in the vacuum of NLSUSY theory as composites of NG fermion.**

- For linearizing /analyzing the vacuum of NLSGR/SGM
we need **broken global LSUSY SUGRA-analogue model** which realizes:

NL/L SUSY relation (equivalence):

$$L_{\text{NLSGR}}(w) = L_{\text{SGM}}(e, \psi) = L_{\text{LSUSY}}(e^a{}_\mu, \psi_\nu(e, \psi), M(e, \psi), N(e, \psi), \dots).$$

5. Cosmological implications of SGM scenario

The variation of SGM action $L_{\text{SGM}}(e, \psi)$ with respect to $e^a{}_\mu$ yields Einstein equation equipping with matter and cosmological term:

$$R_{\mu\nu}(e) - \frac{1}{2}g_{\mu\nu}R(e) = \frac{8\pi G}{c^4}\{\tilde{T}_{\mu\nu}(e, \psi) - g_{\mu\nu}\frac{c^4\Lambda}{8\pi G}\}. \quad (5)$$

where $\tilde{T}_{\mu\nu}(e, \psi)$ abbreviates the stress-energy-momentum of superon(NG fermion) including the gravitational interaction.

- Note that the cosmological term $-\frac{c^4\Lambda}{8\pi G}$ can be interpreted as the negative energy density of space-time, i.e. the dark energy density ρ_D .

- Big collapse(BC) induces 3 dimensional (space-like region) rapid expansion of space-time by Pauli principle:

$$ds^2 = s_{\mu\nu}(x)dx^\mu dx^\nu = \{g_{\mu\nu} + \underline{\Phi_{\mu\nu}(e, \psi)}\}dx^\mu dx^\nu.$$

$$\begin{aligned} \{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = 0 \quad \Rightarrow \quad & \{\psi(x)_a, \bar{\psi}(y)_b\}_{x_0=y_0} = \delta^{(3)}(\mathbf{x} - \mathbf{y})\delta_{ab} \\ (\{\psi(x)_a, \bar{\psi}(y)_b\} = 0 \quad \Rightarrow \quad & \{\psi(x)_a, \bar{\psi}(y)_b\} = iS(x - y)\delta_{ab}) \end{aligned}$$

- BC produces gravitational composite (massless) eigenstates of $SO(N)$ sP angebra, which is the ignition of the Big Bang(BB) SM scenario.
- As shown in the toy model, the vacuum of SGM may explain naturally observed mysterious (numerical) relations:

$$dark\ energy\ density\ \rho_D \sim O(\kappa^{-2}) \sim m_\nu^{-4} \sim (10^{-12} GeV)^4 \sim g_{sv}^{-2},$$

provided λ_D^0 is identified with neutrino and $f\xi \sim O(1)$.

6. Summary

NLSGR(SGM) for unity of nature:

- Ultimate entity; **Unstable** $d = 4$ **space-time**: $[x^a, \psi_\alpha^N; x^\mu]$ described by $[L_{\text{NLSGR}}(w^a_\mu)]$: Global SUSY NLSGR on New space-time with $\Lambda > 0$

⇒ **Big Collapse (BC)**

- The creation of ordinary Riemann space-time [graviton e^a_μ] and massless NG fermionic matter [superon ψ_α^N], rapid expansion by Pauli principle $[L_{\text{SGM}} = L_{\text{EH}}(e) - \Lambda + T(\psi \cdot e)]$: Einstein GR with $\Lambda > 0$ and N superon

- The universal attractive force graviton dictates the evolution of $[L_{\text{SGM}}$ by forming gravitational composite **global** LSUSY supermultiplet corresponding to (massless) eigenstates of space-time symmetry $\text{SO}(10)$ sP.

⇒ **Ignition of Big Bang of the universe toward Λ CDM scenario**

- In flat space-time, broken N -LSUSY theory emerges from the N -NLSUSY cosmological term of $L_{\text{SGM}}(e, \psi)$ [NL/L SUSY relation]. \longleftrightarrow BCS vs GL

The cosmological term is the origin of everything!

- NLSGR((SGM) paradigm can bridge naturally the cosmology and the low energy particle physics, which provides new insights into unsolved problems of cosmology and particle physics,, e.g. the space-time dimension *four*, the origin of mass and SUSY breaking, the dark energy and the dark matter, the dark energy density \approx (neutrino mass)⁴ the tiny neutrino mass, inflation of space-time, the three-generations structure of quarks and leptons, the rapid expansion of space-time, the magnitude of the bare gauge coupling constant, matter dominance of nature , the shape and the role of black hole etc. $NLSGR \rightarrow (BC) \rightarrow SGM \rightarrow (BB) \rightarrow SMs \rightarrow \dots \rightarrow BH \rightarrow NLSGR \rightarrow \dots$

- Many important problems are yet to be studied, e.g., global SUSY for higher spins (5/2 3) and SUGRA
- Direct linearization of NLSGR/SGM
- The revisit of global SUSY superfield formulation from SGM/SQM. viewpoint. \leftrightarrow C-asymmetry, proton decay, etc.

For some details:

1. K. Shima, Invited talk at International Conference, *100 Years Werner Heisenberg - Works and Impact-*, September 26-30, 2001., Bamberg, Germany. *Fortschr. Phys* **50**(2002) 5-7. 517-523, eds. D. Papenfuss, D. Luest and W. P. Schleich.
2. K. Shima, Invited talk at Conference on Cosmology, Gravitational Waves and Particles ,2017, NTU, Singapore (Uploaded at YouTube by IAS). Proceedings of CCGWP, ed. Harald Fritzsch, (World Scientific, Singapore, 2017), 301.
3. K. Shima, Preliminary temporal report: arXiv:2012.01646[hep-th](2020)