

Recent Neutrino Oscillation Results from JUNO

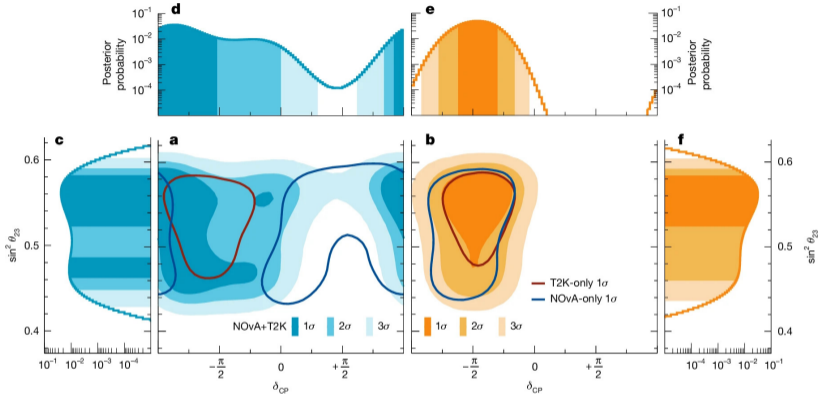
Benda Xu

on behalf of the JUNO collaboration

Tsinghua University, Center for High Energy Physics

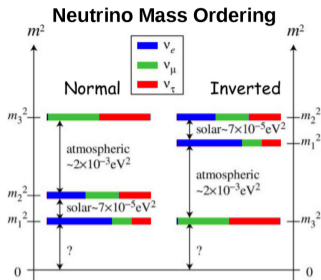


2026-05-26 NPN 2026 @ Mito



Nature 646, 818–824 (2025).

- T2K and NO ν A combined fit shows favor of inverted ordering.
- JUNO is committed to provide the independent & orthogonal NMO input.
 - \rightarrow Synergy with accelerator ν facilities



ν Mass Ordering (NMO): Sign of Δm_{31}^2

- $\Delta m_{21}^2 = m_2^2 - m_1^2 = 7.5 \times 10^{-5} \text{eV}^2$
- $\Delta m_{31}^2 = m_3^2 - m_1^2 = 2.4 \times 10^{-3} \text{eV}^2$
- normal order: $m_3^2 > m_1^2$;
 - inverted order: $m_3^2 < m_1^2$
- imprints subtle difference on energy spectra.

Jiangmen Underground Neutrino Observatory

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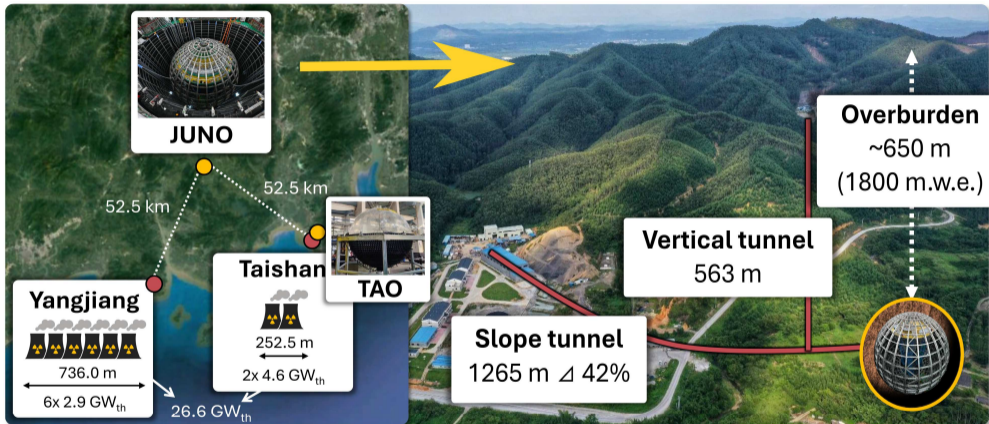
ν Oscillation
Machine

JUNO Design

Detector
Performance

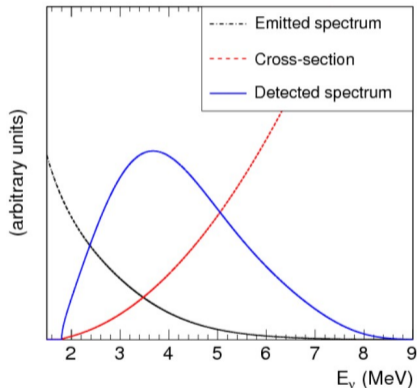
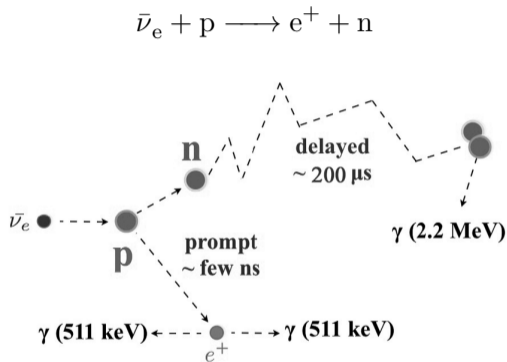
Oscillation
Measurement

Summary



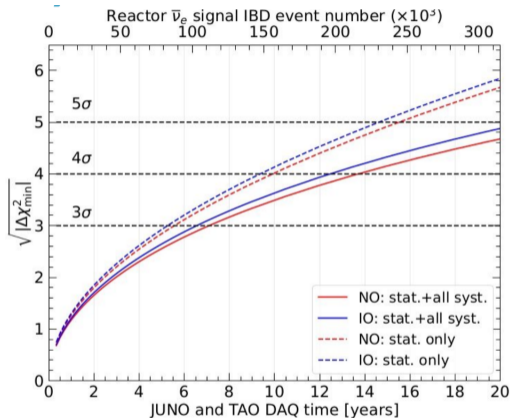
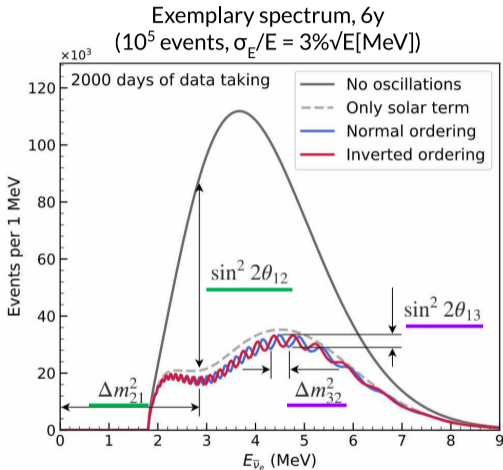
- 52.5 km captures the first minimal of ν_{12} oscillations.

$\bar{\nu}$ Detection with Inverse Beta Decay (IBD)



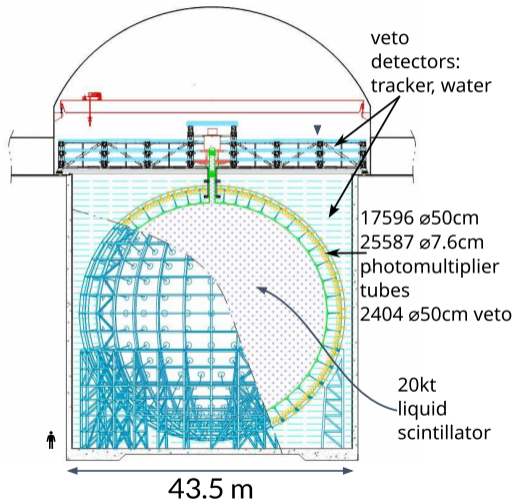
e^+/γ scintillation light are collected by photomultipliers.

Sensitivity to ν Mass Ordering



$\Delta\chi^2$ is the χ^2 difference between correct and wrong hypothesis.

3 σ determination in 6.5 years. Chin. Phys. C 49 033104



A Liquid Scintillator Detector

Large Exposure

- proven technology to scale up.
- world's largest by 20 \times .

Measurement of ν Energy

- bright liquid scintillator.
- maximize the coverage of photomultiplier tubes.
- > 1700 photo-electrons/MeV.

Control of Backgrounds

- effective impurity removal: filtration, distillation, absorption.
- active shield of water and μ tracker.

20 kton LS Target Providing High Statistics

Proximity to 8 nuclear reactors
and **unprecedented size**

evolution of LS detectors

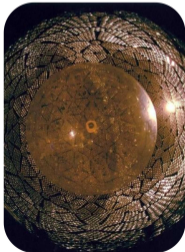
Reines &
Cowan
300 L



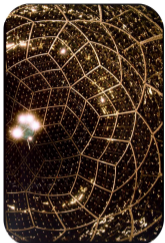
Daya Bay
8 x 20 t



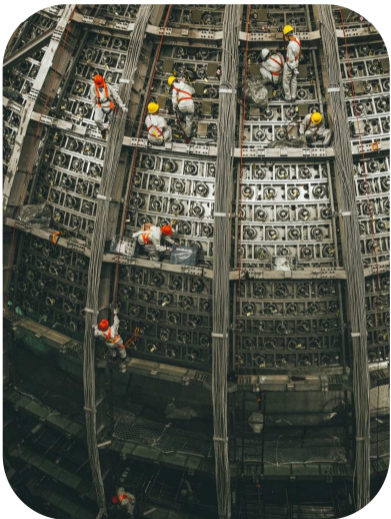
Borexino 300 t



SNO+ 780 t



KamLAND 1,000 t



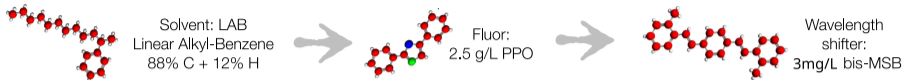
JUNO 20,000 t

High Energy Resolution with Dense PMT Array

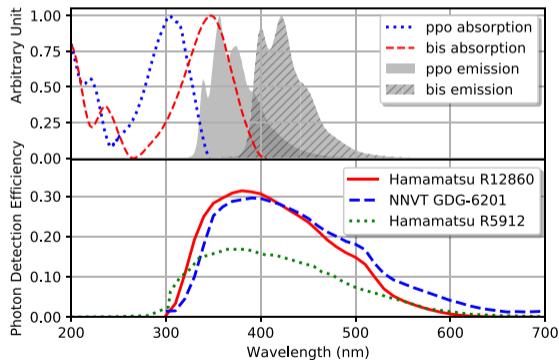
	KamLAND	JUNO
PE yield per MeV	250	>1700
Resolution @ 1MeV	6%–7%	~ 3%
Photocoverage/%	34	78
PMT QE × CE/%	12	30
Relative light yield	1	1.5



Bright Liquid Scintillator

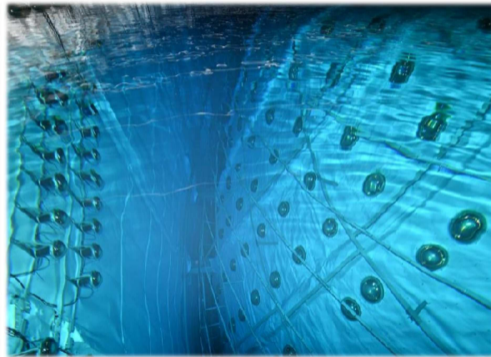


- Carefully designed liquid scintillation cocktail.
- Efficient transmission of ionization energy to scintillation photons.



NIM A 967, 163860 (2020), JINST 20, P05009 (2025).

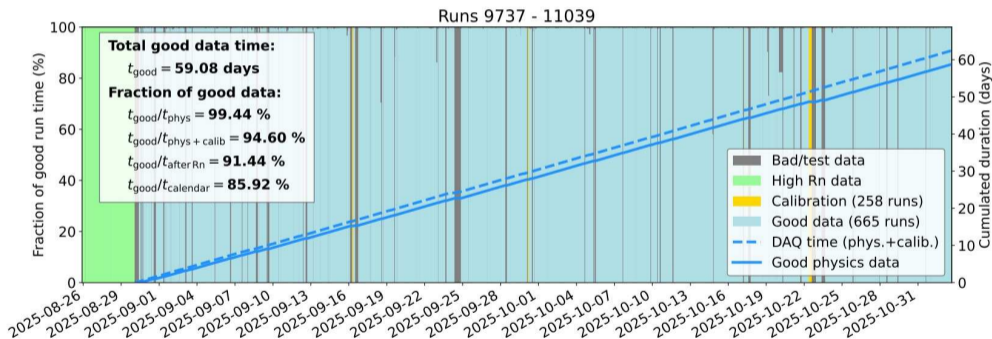
Decade-Long Endeavor



12-years from civil excavation to sub-percent precision physics.

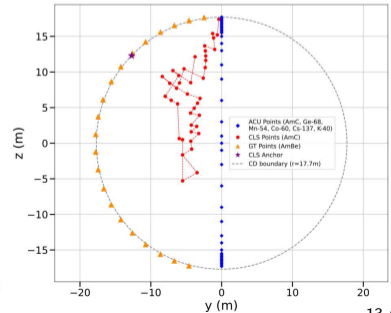
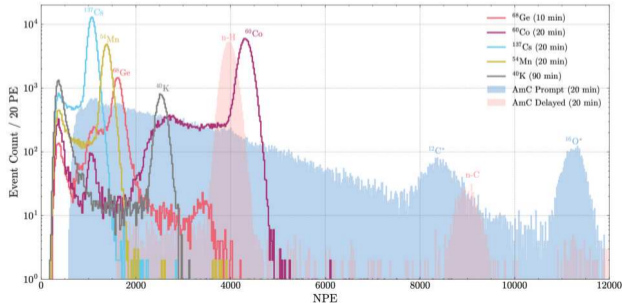
- total LS 23 231.6 m³; total water 41 225.1 m³.

- Data taking $> 98\%$.
 - Calibration twice a week with laser and AmC source¹.
- Physics data $> 92\%$ duty factor.



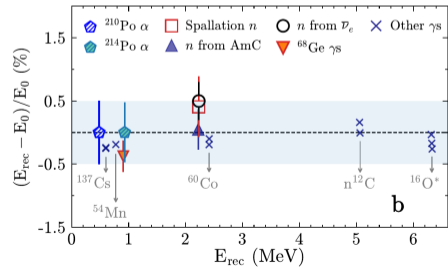
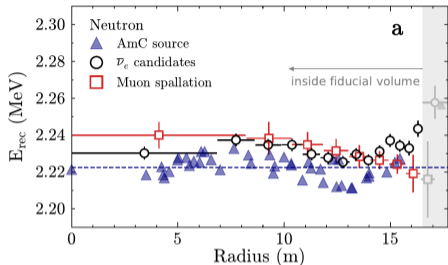
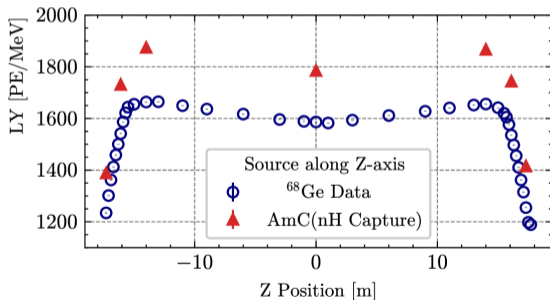
¹Detector maintenance performed during calibration source moving.

- Extensive calibration at the beginning of data taking of Aug. 2025
- Multiple calibration sources covering wide energies and positions.
- Natural radioactivity and cosmogenic sources complements
 - ^{210}Po , ^{214}Bi - ^{214}Po , ^{12}B , ^{11}C .



PE Yield and Residual Non-Uniformity

- Analytic/ML tools driving the energy resolution limit.
- Energy non-uniformity within $\pm 0.5\%$ for $r < 16.5$ m



Energy Resolution and Non-Linearity

- Energy non-linearity calibrated to $< 1\%$.
- Energy resolution is approaching the design level.

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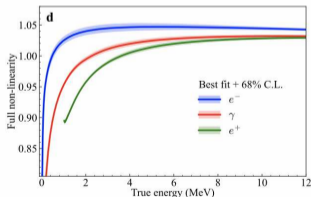
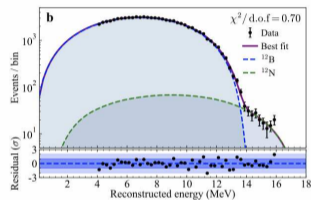
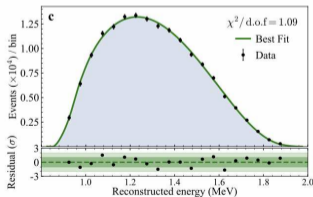
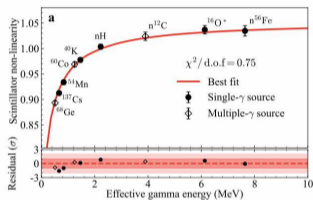
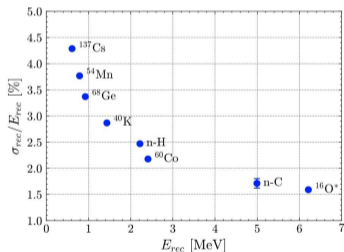
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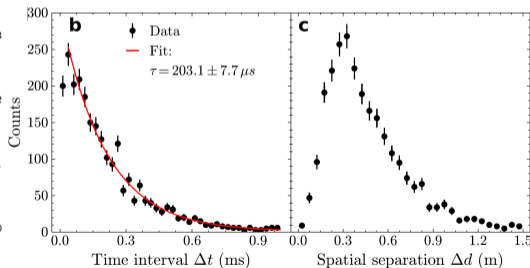
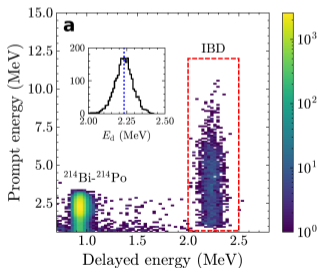
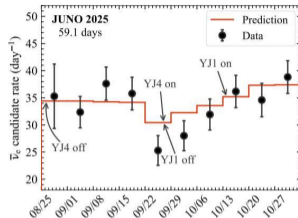
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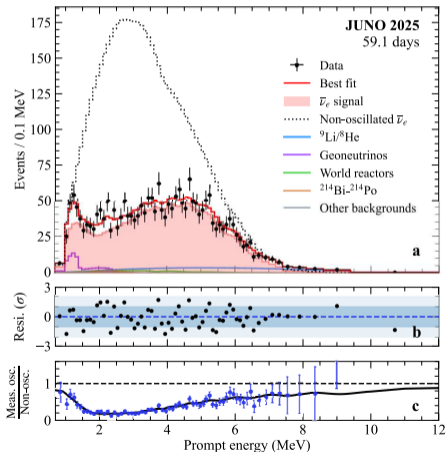
Chin. Phys. C 50, 043001 (2026); arXiv:2511.14593

- Final IBD coincidence selection is by tempo-spatial and energy criteria.
- Signal efficiency $\sim 70\%$, fiducial volume $R < 16.5$ m dominant.



The Reactor $\bar{\nu}_e$ Spectrum

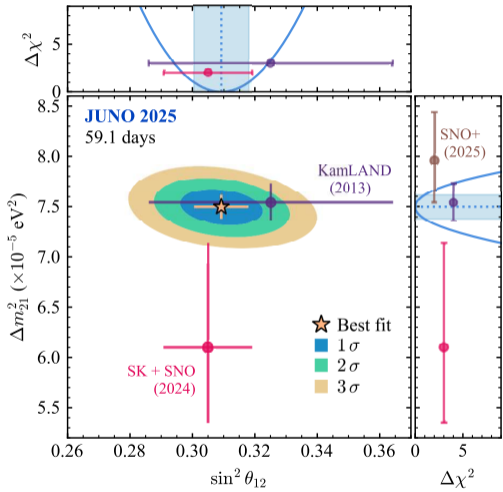
- Show the breathtaking data vs. un-oscillated prediction plot.
- The visible oscillation patterns.



arXiv:2511.14593

Antineutrinos ($\bar{\nu}_e$) Candidates Summary		
DAQ live time (days)	59.1	
$\bar{\nu}_e$ candidates	2379	
Selection Efficiencies (%)	ε	σ_{rel}
Fiducial volume	80.6	1.6
PMT flasher rejection	>99.9	negligible
μ veto	93.6	negligible
Multiplicity	97.4	negligible
Prompt-delayed coinc.	95.1	0.13
Total efficiency (ε_{tot})	69.9	1.6
$\bar{\nu}_e$ signal (cpd¹)		
w/o ε_{tot} corrected	33.5 ± 1.7	
w/ ε_{tot} corrected	47.9 ± 2.6	
Non-oscillated $\bar{\nu}_e$	150.9 ± 2.7	
Backgrounds (cpd)	Pre-fit	Best-fit
$^9\text{Li}/^8\text{He}$	4.3 ± 1.4	3.9 ± 0.6
Geoneutrinos	1.2 ± 0.5	1.4 ± 0.4
World reactors	0.88 ± 0.09	0.88 ± 0.09
$^{214}\text{Bi}-^{214}\text{Po}$	0.18 ± 0.10	0.20 ± 0.10
$^{13}\text{C}(\alpha, n)^{16}\text{O}$	0.04 ± 0.02	0.04 ± 0.02
Fast neutrons	0.02 ± 0.02	0.02 ± 0.02
Double neutrons	0.05 ± 0.05	0.07 ± 0.05
Atmospheric neutrinos	0.08 ± 0.04	0.07 ± 0.04
Accidentals ($\times 10^{-2}$)	4.9 ± 0.3	4.9 ± 0.3

Results of $\sin^2 \theta_{12}$ and Δm_{21}^2



Experiment	Value
SNO (2011)	$0.299^{+0.016}_{-0.015}$ 5.08%
SK (2024)	$0.324^{+0.027}_{-0.023}$ 7.72%
all solar ν (2024)	$0.306^{+0.013}_{-0.013}$ 4.25%
KamLAND (2013)	$0.325^{+0.039}_{-0.039}$ 12.02%
JUNO	$0.309^{+0.0087}_{-0.0087}$ 2.81%

Experiment	Value
SNO (2011)	$5.6^{+1.9}_{-1.4}$ 29.46%
SK (2024)	$6.10^{+1.26}_{-0.86}$ 17.38%
all solar ν (2024)	$6.10^{+0.95}_{-0.81}$ 14.43%
KamLAND (2013)	$7.54^{+0.19}_{-0.18}$ 2.45%
SNO+ (2025)	$7.96^{+0.48}_{-0.42}$ 5.65%
JUNO	$7.50^{+0.12}_{-0.12}$ 1.55%

Unprecedented precision achieved by JUNO. arXiv:2511.14593

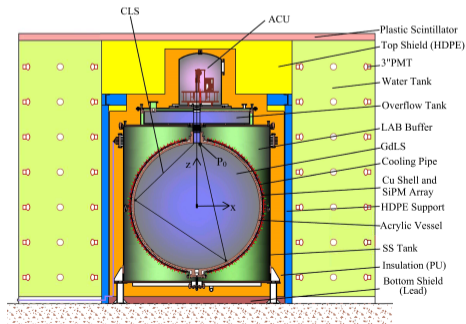
- JUNO is fully operational and delivering sub-percent precision on Δm^2_{21} .

Impact on the 3-flavor ν oscillation paradigm

- JUNO's precision tightly constrains the parameter space.
 - Powerful input to test leptonic flavor models.
- Setting the stage for the mass ordering measurement.

JUNO-TAO

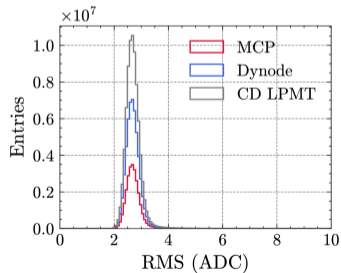
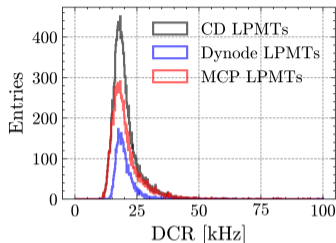
- JUNO-TAO as the ultimate reference for reactor spectrum.
- TAO started data-taking in Feb. 2026.



PMT electronics stability

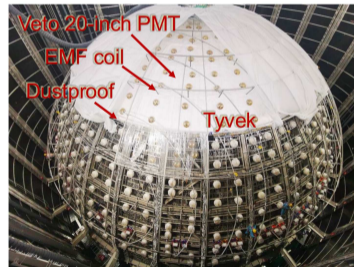
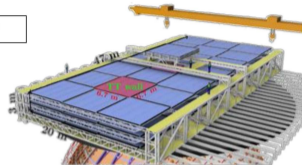
- Large PMTs (LPMTs):
 - Lost 0.1% LPMTs during installation in the Central Detector (CD)
 - Approximately 2-3% of large PMTs flashing at any one time
 - 20.6 kHz and 22.7 kHz of Dark Count Rate (DCR) for dynode and MPC-PMTs, respectively
- Electronics:
 - Low noise levels (mean pedestal RMS of 2.6 ADC counts roughly 0.05 PE)
 - PMT Threshold at 0.2-0.3 PE
- Trigger:
 - Threshold at 200 keV

[arXiv:2511.14590](https://arxiv.org/abs/2511.14590)



- **Water Cherenkov + Top tracker**
- Water Cherenkov detector
 - **35 kton** water to shield backgrounds from the rock
 - Instrumented w/ **2400 20-inch PMTs** on SS structure
 - Water pool lining: 5mm HDPE (black) to keep the clean water and to stop Rn from the rock, covered w/ tyvek
 - **100 ton/h pure water system**
Requirement: U/Th/K < 10^{-14} g/g and Rn < 10 mBq/m³, attenuation length > 40 m, temperature controlled to (21 ± 1) °C
- Top tracker
 - Refurbished OPERA scintillators
 - 3 layers, ~60% coverage on the top
 - $\Delta\theta \sim 0.2^\circ$, $\Delta D \sim 20$ cm
- Earth Magnetic Field compensation coil

NIMA 1057 (2023) 168680

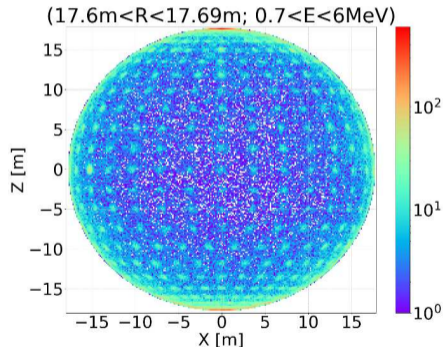


Measured radioactivity in LS and water better than specification

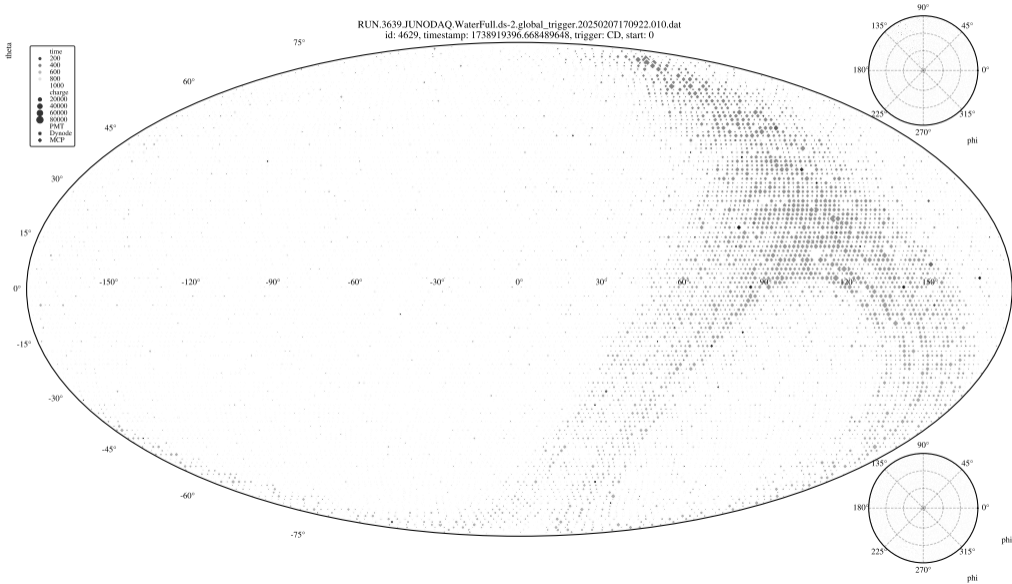
- Radiopurity control of raw material, cleaning of inner surface, ^{222}Rn control

Central Detector	Specification	Measure
Singles Rate (Hz)	< 7.2 Hz	< 7 Hz
$^{238}\text{U}/^{232}\text{Th}$ ($\times 10^{-15}$ g/g)	< 1	< 0.1
^{210}Po ($\times 10^4$ cpd/kt)	< 8	< 5

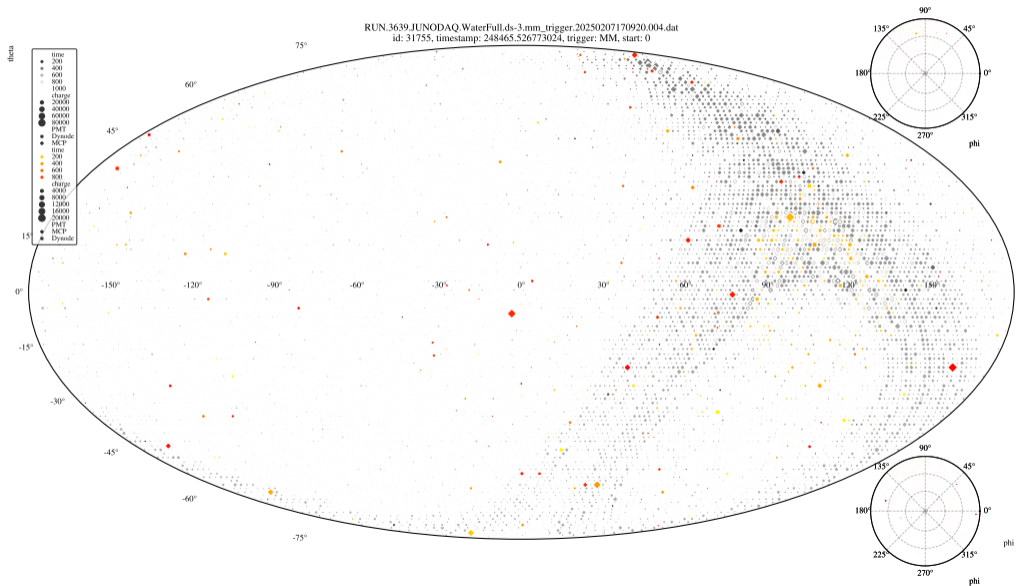
Muon Veto Detector	Specification	Measure
^{222}Rn (mBq/m ³)	< 10	< 10
^{226}Ra (mBq/m ³)	< 1	< 0.01
U/Th ($\times 10^{-15}$ g/g)	< 10	< 0.4



Water phase: stopping muon



Water phase: michel electron



Oscillation Fit Statistical and Systematic Breakdown

