
Radiative kaon decays: where do we stand?

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Rare (radiative) Kaon decays

- Rare kaon decays: extremely useful probes of
 - (a) new physics (when SM-suppressed and short-distance dominated).
 - (b) the SM itself in the nonperturbative regime (when long-distance dominated).
- Focus of this talk: hadronic decay modes with photons or $\ell^+\ell^-$ pairs (long-distance dominated).

$$K \rightarrow \pi\gamma^{(*)}, \pi\pi\gamma^{(*)}$$

- Main goal: determination of ChPT couplings at NLO, i.e., precise knowledge of the SM at low energies.

Theoretical description

- Kaon decays can be described with ChPT, the low-energy EFT of the strong interactions. $\Delta S = 1$ sector:

$$\mathcal{L}_{\Delta S=1} = G_8 f_\pi^4 \text{tr}[\lambda_6 D_\mu U^\dagger D^\mu U] + G_8 f_\pi^2 \sum_j N_j W_j(U, D_\mu U, \lambda_6) + \mathcal{O}(p^6)$$

with

$$U = \exp\left[i\frac{\phi^a \tau^a}{f_\pi}\right]; \quad D_\mu U = \partial_\mu U + ieA_\mu[Q, U]$$

- The LO is universal, NLO order contains nonperturbative (hadronic) information inside N_i .
- Radiative kaon decays: out of the 37 NLO operators, sensitive to combinations of W_{14}, \dots, W_{18} (CP-even) and W_{28}, \dots, W_{31} (CP-odd).
- General structure of the amplitudes:

$$\mathcal{M}(K \rightarrow X \gamma^{(*)}) = \underbrace{\mathcal{M}_B(\mathcal{O}(p^2))}_{\text{Brems.}} + \underbrace{\mathcal{M}_E(\mathcal{O}(p^4))}_{\text{electric, CP-even}} + \underbrace{\mathcal{M}_M(\mathcal{O}(p^4))}_{\text{magnetic, CP-odd}}$$

Radiative kaon decays

MEASURED MODES:

$$K^\pm \rightarrow \pi^\pm \pi^0 \gamma \quad \sim 10^{-6}$$

$$K^\pm \rightarrow \pi^\pm \gamma \gamma \quad \sim 10^{-6}$$

$$K^\pm \rightarrow \pi^\pm \gamma^* \quad \sim 10^{-7}$$

$$K_S \rightarrow \pi^0 \gamma^* \quad \sim 10^{-9}$$

$$K_S \rightarrow \pi^0 \gamma \gamma \quad \sim 10^{-8}$$

NEAR FUTURE:

$$K^\pm \rightarrow \pi^\pm \pi^0 \gamma^* \quad \sim 10^{-6} \quad \text{NA48, NA62}$$

$$K_S \rightarrow \pi^+ \pi^- \gamma^* \quad \sim 10^{-5} \quad \text{LHCb}$$

$$K_S \rightarrow \mu^+ \mu^- \quad < 10^{-9} \quad \text{LHCb}$$

$$K_S \rightarrow 4\ell \quad \text{LHCb}$$

Measured radiative kaon decays

- Weak chiral couplings are related to the slope of the differential decay rate, most easily accessed through the interference term between Bremsstrahlung and electric emission.
- Experimental effort for the last 15 years, and ongoing:

$$K^\pm \rightarrow \pi^\pm \gamma^* : \quad a_+ = -0.578 \pm 0.016 \quad [\text{NA48/2, 2009} - 11]$$

$$K_S \rightarrow \pi^0 \gamma^* : \quad a_S = (1.06_{-0.21}^{+0.26} \pm 0.07) \quad [\text{NA48/1, 2003} - 04]$$

$$K^\pm \rightarrow \pi^\pm \pi^0 \gamma : \quad X_E = (-24 \pm 4 \pm 4) \text{ GeV}^{-4} \quad [\text{NA48/2, 2010}]$$

$$K^+ \rightarrow \pi^+ \gamma \gamma : \quad \hat{c} = 1.56 \pm 0.23 \pm 0.11 \quad [\text{NA62, 2014}]$$

- The slopes are linked to ChPT through

[Ecker et al; D'Ambrosio et al]

$$\mathcal{N}_E^{(1)} \equiv N_{14}^r - N_{15}^r = \frac{3}{64\pi^2} \left(\frac{1}{3} - \frac{G_F}{G_8} a_+ - \frac{1}{3} \log \frac{\mu^2}{m_K m_\pi} \right) - 3L_9^r$$

$$\mathcal{N}_S \equiv 2N_{14}^r + N_{15}^r = \frac{3}{32\pi^2} \left(\frac{1}{3} + \frac{G_F}{G_8} a_S - \frac{1}{3} \log \frac{\mu^2}{m_K^2} \right)$$

$$\mathcal{N}_E^{(0)} \equiv N_{14}^r - N_{15}^r - N_{16}^r - N_{17}^r = -\frac{|\mathcal{M}_K| f_\pi}{2G_8} X_E$$

$$\mathcal{N}_0 \equiv N_{14}^r - N_{15}^r - 2N_{18}^r = \frac{3}{128\pi^2} \hat{c} - 3(L_9^r + L_{10}^r)$$

Status of NLO chiral counterterms

Decay mode	counterterm combination	expt. value
$K^\pm \rightarrow \pi^\pm \gamma^*$	$N_{14} - N_{15}$	$-0.0167(13)$
$K_S \rightarrow \pi^0 \gamma^*$	$2N_{14} + N_{15}$	$+0.016(4)$
$K^\pm \rightarrow \pi^\pm \pi^0 \gamma$	$N_{14} - N_{15} - N_{16} - N_{17}$	$+0.0022(7)$
$K^\pm \rightarrow \pi^\pm \gamma \gamma$	$N_{14} - N_{15} - 2N_{18}$	$-0.0017(32)$

- From $K \rightarrow \pi \gamma^*$ decays,

$$N_{14} = (-2 \pm 18) \times 10^{-4}; \quad N_{15} = (1.65 \pm 0.22) \times 10^{-2}$$

- Adding $K \rightarrow \pi \gamma \gamma$,

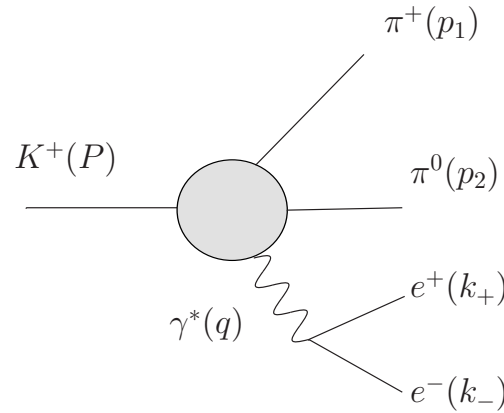
$$N_{18} = (-7.5 \pm 2.3) \times 10^{-3}$$

- So far, only the combination $N_{16} + N_{17}$ constrained. One extra combination needed:

$$K \rightarrow \pi \pi \ell^+ \ell^-$$

- Most promising channel: $K^\pm \rightarrow \pi^\pm \pi^0 e^+ e^-$, under analysis at NA48. Important alternative: $K_L \rightarrow \pi^+ \pi^- e^+ e^-$.

$$K^+ \rightarrow \pi^+ \pi^0 e^+ e^-$$



- The hadronic piece contains 3 form factors:

$$H^\mu(p_1, p_2, q) = F_1 p_1^\mu + F_2 p_2^\mu + F_3 \varepsilon^{\mu\nu\alpha\beta} p_{1\nu} p_{2\alpha} q_\beta$$

- Relevant weak coupling combinations:

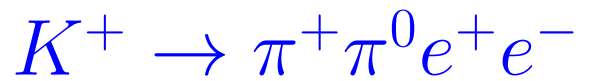
$$\mathcal{N}_E^{(0)} \equiv N_{14}^r - N_{15}^r - N_{16}^r - N_{17} = +0.0022(7)$$

$$\mathcal{N}_E^{(1)} \equiv N_{14}^r - N_{15}^r = -0.0167(13)$$

$$\mathcal{N}_E^{(2)} = N_{14}^r + 2N_{15}^r - 3(N_{16}^r - N_{17})$$

e.g.

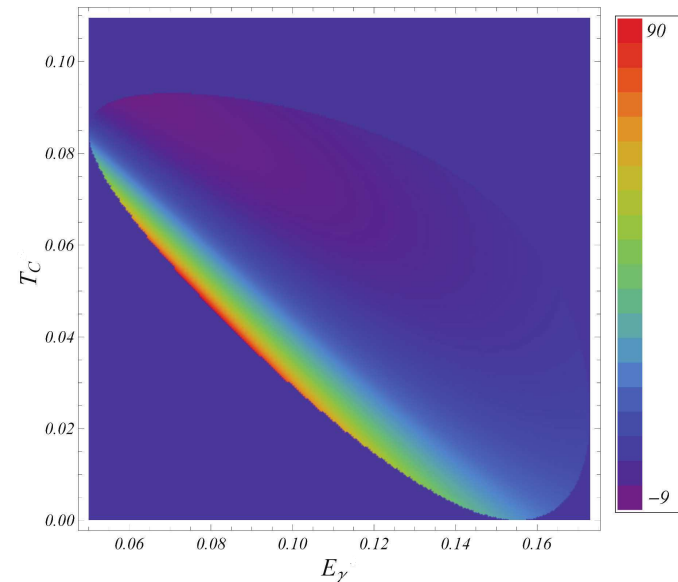
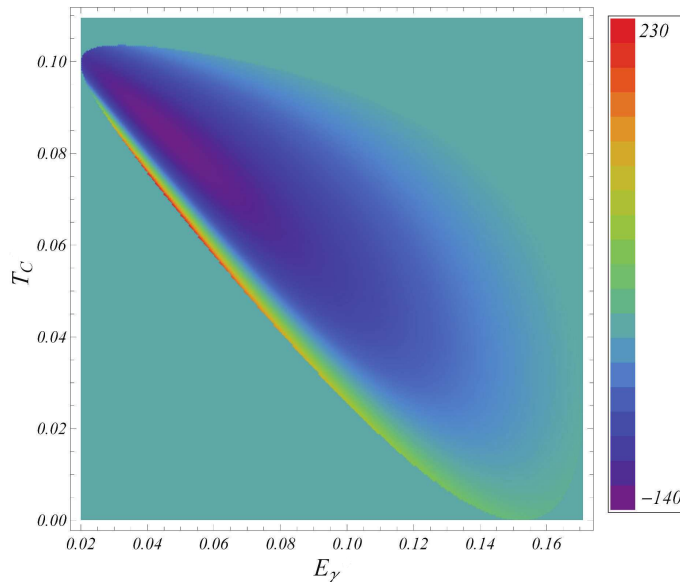
$$F_2 = -\frac{2ie}{2q \cdot p_K - q^2} \mathcal{M}_K e^{i\delta_0^2} + \frac{2ieG_8 e^{i\delta_1^1}}{f_\pi} \left\{ q \cdot p_+ \mathcal{N}_E^{(0)} - \frac{1}{3} q^2 \mathcal{N}_E^{(2)} \right\}$$



- Challenge: how to overcome the $\mathcal{O}(p^2)$ contribution. In terms of integrated branching ratios:

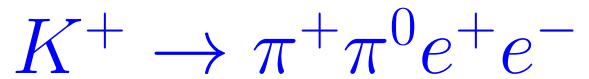
$$\Gamma_M \sim \frac{1}{70} \Gamma_B; \quad \Gamma_{\text{INT}} \sim 10^{-2} \Gamma_B$$

- Best strategy: Bremsstrahlung is peaked at low q^2 . Use cuts in the photon energy. [Pichl'01; Capiello, OC, D'Ambrosio'11,'17]



- The gain depends on the size of $\mathcal{N}_E^{(2)}$, but it can be at least a factor 10.
- Using the already measured radiative decays:

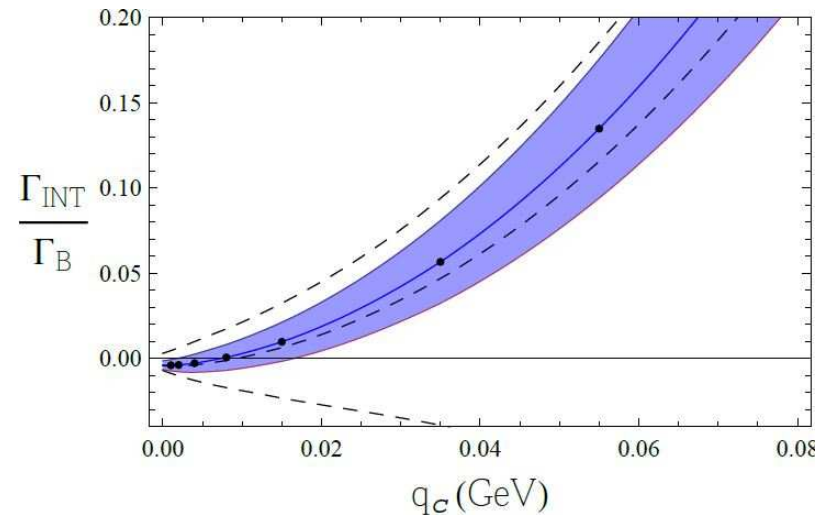
$$\mathcal{N}_E^{(2)} = +0.089(11) + 6N_{17}$$



- Prediction: a very large $\mathcal{N}_E^{(2)}$ counterterm.

$$\mathcal{N}_E^{(2)} \sim +\mathcal{O}(10^{-1})$$

- Rather robust: unless N_{17} is large (theoretically disfavored) and negative.
- An analysis shows that the interference has a characteristic pattern depending mostly on $\mathcal{N}_E^{(0)}$ and $\mathcal{N}_E^{(2)}$.



- NA48 has collected 5000 events. Unclear whether this will allow for an informative fit (analysis in progress) or one would need the larger statistics of NA62.

$$K_S \rightarrow \pi^+ \pi^- e^+ e^-$$

- A similar analysis can be performed for $K_S \rightarrow \pi^+ \pi^- e^+ e^-$. Total branching ratio measured:

$$BR(K_S \rightarrow \pi^+ \pi^- e^+ e^-)_{\text{exp}} = (4.79 \pm 0.15) \times 10^{-5}$$

Access to interference term requires typically a percent precision. Challenging but in the LHCb agenda.

- On top of $\mathcal{N}_E^{(0)}$ (known), new counterterm, but not independent:

$$\mathcal{N}_E^{(3)} = N_{14} - N_{15} - 3(N_{16} + N_{17}) = -2\mathcal{N}_E^{(1)} + 3\mathcal{N}_E^{(0)} = +0.040(5)$$

- Prediction of the LO and NLO chiral contributions:

$$BR(K_S \rightarrow \pi^+ \pi^- e^+ e^-) = \underbrace{4.74 \cdot 10^{-5}}_{\text{Brems.}} + \underbrace{4.39 \cdot 10^{-8}}_{\text{Int.}} + \underbrace{1.33 \cdot 10^{-10}}_{\text{DE}}$$

- Extraction of N_{17} from neutral kaon decays possible with information on $K_L \rightarrow \pi^+ \pi^- e^+ e^-$:

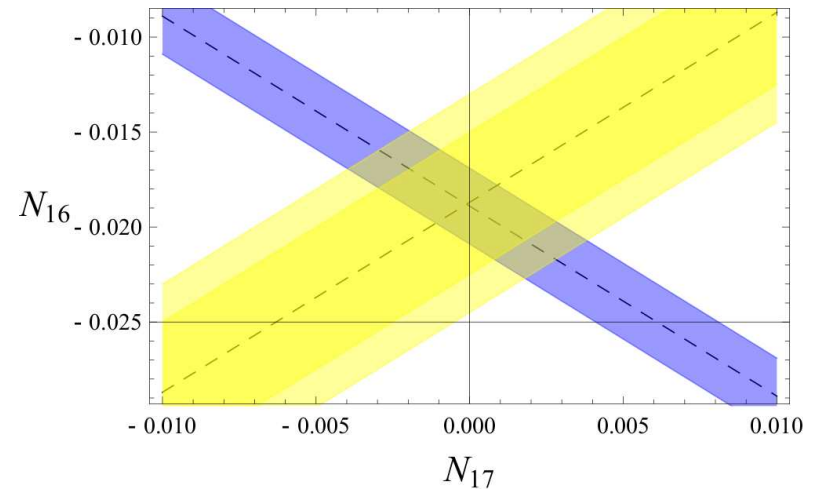
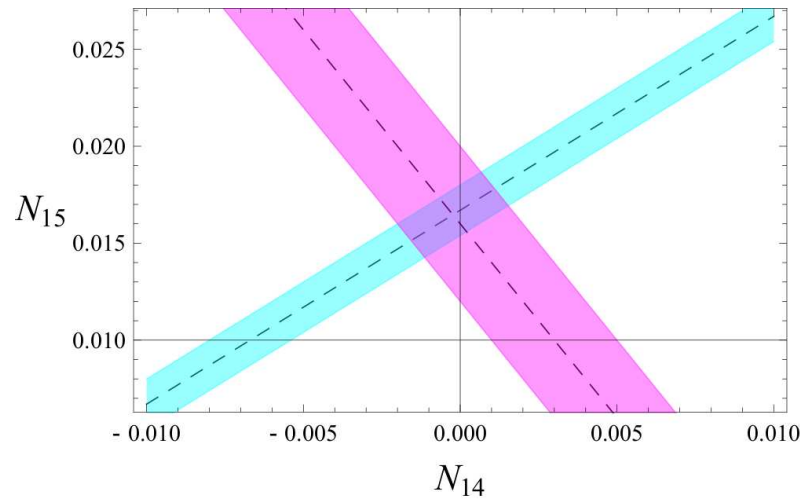
$$\mathcal{N}_E^{(4)} - \mathcal{N}_E^{(3)} = 6N_{17}$$

- Muon decay extremely suppressed by phase space:

$$BR(K_S \rightarrow \pi^+ \pi^- \mu^+ \mu^-) = \underbrace{4.17 \cdot 10^{-14}}_{\text{Brems.}} + \underbrace{4.98 \cdot 10^{-15}}_{\text{Int.}} + \underbrace{2.17 \cdot 10^{-16}}_{\text{DE}}$$

Radiative weak couplings

Current situation:



- Strong hierarchies:

$$N_{14} = (-2 \pm 18) \times 10^{-4}; \quad N_{15} = (1.65 \pm 0.22) \times 10^{-2}$$

- N_{16} vs N_{17} based on an educated guess ($N_{17} \lesssim 10^{-3}$). If experiment contradicts it, rather exceptional failure of all the theoretical frameworks.

Weak counterterms (theory)

- ChPT built on chiral symmetry: operator catalog at each loop order. Coefficients depend on physics above the GeV scale: predictions are model-dependent.
- Main ideas: (explicit) resonance models and weak deformation models.

counterterm combinations	decay mode	WDM/FM/HEW	R^μ
$N_{14} - N_{15}$	$K^\pm \rightarrow \pi^\pm \gamma^*$	$-3L_9 - L_{10} - 2H_1$	$-0.02\eta_V$
$2N_{14} + N_{15}$	$K_S \rightarrow \pi^0 \gamma^*$	$-2L_{10} - 4H_1$	$0.08\eta_V$
$N_{14} - N_{15} - N_{16} - N_{17}$	$K^\pm \rightarrow \pi^\pm \pi^0 \gamma$	$-2(L_9 + L_{10})$	$-0.01\eta_A$
$N_{14} - N_{15} - 2N_{18}$	$K^\pm \rightarrow \pi^\pm \gamma \gamma$	$-3(L_9 + L_{10})$	$-0.01\eta_A$
$N_{14} + 2N_{15} - 3(N_{16} - N_{17})$	$K^\pm \rightarrow \pi^\pm \pi^0 \gamma^*$	$6L_9 - 4L_{10} + 4H_1$	$0.16\eta_V + 0.01\eta_A$
$N_{14} - N_{15} - 3(N_{16} - N_{17})$	$K_L \rightarrow \pi^+ \pi^- \gamma^*$	$-4L_{10} + 4H_1$	$0.04\eta_V + 0.01\eta_A$
$N_{14} - N_{15} - 3(N_{16} + N_{17})$	$K_S \rightarrow \pi^+ \pi^- \gamma^*$	$-4L_{10} + 4H_1$	$0.04\eta_V - 0.04\eta_A$
$7(N_{14} - N_{16}) + 5(N_{15} + N_{17})$	$K_S \rightarrow \pi^+ \pi^- \pi^0 \gamma$	$10L_9 - 14L_{10}$	$0.48\eta_V + 0.01\eta_A$

$$N_{14} - N_{15} \simeq +\mathcal{O}(10^{-2}) ;$$

$$2N_{14} + N_{15} \simeq -\mathcal{O}(10^{-2}) ;$$

$$N_{14} - N_{15} - N_{16} - N_{17} \simeq \mathcal{O}(10^{-3}) ;$$

$$N_{14} - N_{15} - 2N_{18} \simeq \mathcal{O}(10^{-3}) ;$$

$$N_{14} + 2N_{15} - 3(N_{16} - N_{17}) \simeq +\mathcal{O}(10^{-2}) ;$$

$$N_{14} - N_{15} - 3(N_{16} - N_{17}) \simeq +\mathcal{O}(10^{-2}) ;$$

$$N_{14} - N_{15} - 3(N_{16} + N_{17}) \simeq +\mathcal{O}(10^{-2}) ;$$

$$7(N_{14} - N_{16}) + 5(N_{15} + N_{17}) \simeq +\mathcal{O}(10^{-1})$$

Conclusions

- First fully experimentally-based determination of all the NLO weak chiral couplings relevant for radiative kaon decays soon to be there.
- An extraction of N_{16} and N_{17} with NA48 statistics requires (a) some expectation for the size of the counterterm combinations to be tested; (b) a strategy to compensate the overwhelming dominance of the Bremsstrahlung contributions.
- So far, results are in qualitative agreement with models (once their limitations are taken into account). Highly nontrivial.
- LHCb and NA62 will continue data taking on radiative kaon decays. More precise knowledge of the SM at low energies achievable through the measurements of $K^\pm \rightarrow \pi^\pm \pi^0 e^+ e^-$, $K_S \rightarrow \pi^+ \pi^- e^+ e^-$, and other modes in the near future.