

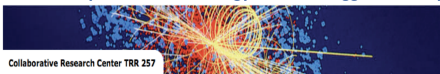
New Results for inclusive $b \rightarrow c$ semileptonic transitions

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Introduction

- V_{cb} is one of the best known CKM matrix elements
- Inclusive determination:
 - Based on the Heavy Quark Expansion (HQE)
 - Close to the 1% theoretical uncertainty
- Exclusive determination
 - Based on Lattice determinations of Form Factors
 - Simulations with finite quark masses
 - Not yet at the 1% level of uncertainties

Recent data indicate that there is no V_{cb} problem

Part I

Alternative V_{cb} Determination

ThM, K. K. Vos: arXiv:1802.09409

M. Fael, ThM, K. K. Vos: arXiv:1812.07472

Inclusive V_{cb} Determination (See Paolo Gambino's Talk)

- Based on the HQE for the inclusive rates and for moments of spectra
- (Cut) moments of the charged lepton energy, hadronic energy and hadronic invariant mass spectra
- Extract the HQE parameters from this data
- Obtain V_{cb} from the total semileptonic rate

Problem: Number of HQE parameters in higher orders!

- 4 up to $1/m^3$
- 13 up to $1/m^4$ (tree level)
- 31 up to order $1/m^5$ (tree level)
- **Factorial Proliferation**

Reparametrization in HQE (Dugan, Golden, Grinstein, Chen, Luke, Manohar...)

Start from the operator:

$$R(q) = \int d^4x e^{iqx} T[\bar{Q}(x)\Gamma q(x) \bar{q}(0)\Gamma^\dagger Q(0)]$$

and replace $Q(x) = \exp(-im(v \cdot x))Q_v(x)$

$$R(S) = \int d^4x e^{-iSx} T[\bar{Q}_v(x)\Gamma q(x) \bar{q}(0)\Gamma^\dagger Q_v(0)]$$

with $S = mv - q$.

These expressions are independent of v !

Perform the OPE \longrightarrow HQE

$$\begin{aligned}
 R(S) &= \sum_{n=0}^{\infty} [C_{\mu_1 \dots \mu_n}^{(n)}(S)]_{\alpha\beta} \bar{Q}_{V,\alpha}(iD_{\mu_1} \dots iD_{\mu_n}) Q_{V,\beta} \\
 &= \sum_{n=0}^{\infty} C_{\mu_1 \dots \mu_n}^{(n)}(S) \otimes \bar{Q}_V(iD_{\mu_1} \dots iD_{\mu_n}) Q_V
 \end{aligned}$$

These expressions are still invariant under reparametrization of v : (as long as the sum is not truncated)

$$\delta_{\text{RP}} v_\mu = \delta v_\mu \quad \text{with} \quad v \cdot \delta v = 0$$

$$\delta_{\text{RP}} iD_\mu = -m \delta v_\mu$$

$$\delta_{\text{RP}} Q_V(x) = im(x \cdot \delta v) Q_V(x) \quad \text{in particular} \quad \delta_{\text{RP}} Q_V(0) = 0.$$

The RP connects different orders in $1/m$, which yields the master relation between the coefficients $n = 0, 1, 2, \dots$

$$\delta_{\text{RP}} C_{\mu_1 \dots \mu_n}^{(n)} = m \delta v^\alpha \left(C_{\alpha \mu_1 \dots \mu_n}^{(n+1)} + C_{\mu_1 \alpha \mu_2 \dots \mu_n}^{(n+1)} + \dots + C_{\mu_1 \dots \mu_n \alpha}^{(n+1)} \right)$$

Use these coefficients, integrate over phase space, get a total rate $\Gamma = \text{Im} \langle B | R | B \rangle = \text{Im} \langle R \rangle$

The coefficients of the OPE will depend only on v

$$R = \sum_{n=0}^{\infty} c_{\mu_1 \dots \mu_n}^{(n)}(v) \otimes \bar{Q}_v(iD_{\mu_1} \dots iD_{\mu_n}) Q_v$$

and satisfy the master relation between different orders in the HQE

Making use of RPI ...

- RPI is a consequence of Lorentz invariance of QCD
- RPI is an exact symmetry:
the relations must hold to all order in α_s
- Resummation of towers of terms from different orders
- For Lorentz invariant observables:
 - The master relations are identical for all observables
 - “Rigid” relations between coefficients
 - Reduction of HQE parameters due to RPI

How does this happen? A Toy example without gluons

Look at the partonic result for the rate

$$\begin{aligned} R(p) &= R(p^2) = R((mv + k)^2) = R(m^2 + 2m(vk) + k^2) \\ &= R(m^2) + R'(m^2)(2m(vk) + k^2) + \frac{1}{2}R''(m^2)(2m(vk) + k^2)^2 \\ &= R(m^2) \end{aligned}$$

if there are no gluons: **Equation of motion:**

$$(2m(vk) + k^2) \rightarrow 2m(iv\partial) + (i\partial)^2 \rightarrow 0$$

All the fully symmetrized (zero gluon) contributions are contained in the partonic result

HQE parameters (for the total rate) to $O(1/m^4)$

$$2m_H \mu_3 = \langle H(p) | \bar{Q}_v Q_v | H(p) \rangle = \langle \bar{Q}_v Q_v \rangle$$

$$2m_H \mu_G = \langle \bar{Q}_v (iD^\mu) (iD^\nu) (-i\sigma_{\mu\nu}) Q_v \rangle$$

$$2m_H \rho_D = \langle \bar{Q}_v \left[(iD^\mu), \left[\left((ivD) + \frac{(iD)^2}{2m} \right), (iD_\mu) \right] \right] Q_v \rangle$$

$$2m_H r_G^4 = \langle \bar{Q}_v [(iD_\mu), (iD_\nu)] [(iD^\mu), (iD^\nu)] Q_v \rangle$$

$$2m_H r_E^4 = \langle \bar{Q}_v [(ivD), (iD_\mu)] [(ivD), (iD^\mu)] Q_v \rangle$$

$$2m_H s_B^4 = \langle \bar{Q}_v [(iD_\mu), (iD_\alpha)] [(iD^\mu), (iD_\beta)] (-i\sigma^{\alpha\beta}) Q_v \rangle$$

$$2m_H s_E^4 = \langle \bar{Q}_v [(ivD), (iD_\alpha)] [(ivD), (iD_\beta)] (-i\sigma^{\alpha\beta}) Q_v \rangle$$

$$2m_H s_{qB}^4 = \langle \bar{Q}_v [iD_\mu, [iD^\mu, [iD_\alpha, iD_\beta]]] (-i\sigma^{\alpha\beta}) Q_v \rangle$$

- Less that had been identified before
- Depend on the quark mass
- Are expressed NOT in terms of iD_{\perp} , rather “full” derivatives
- Can be expressed in terms of full QCD operators via

$$i\not{D} Q_v \rightarrow \left(i\not{D} + \frac{(iD)^2}{m} \right) Q_v = \frac{1}{2m} ((iD)^2 - m^2) Q$$

Thus

$$2m_H \mu_3 = \langle \bar{Q} Q \rangle$$

$$2m_H \mu_G = \langle \bar{Q} (iD^\mu) (iD^\nu) (-i\sigma_{\mu\nu}) Q \rangle$$

$$2m_H \rho_D = \frac{1}{2m} \langle \bar{Q} [(iD^\mu), [(iD)^2, (iD_\mu)]] Q \rangle$$

....

Example 1: $b \rightarrow s\gamma$ (O_7 contribution only)

$$\Gamma_{b \rightarrow s\gamma} = \frac{\lambda^2 m_b^3}{4\pi} \left[\mu_3 - \frac{2}{m_b^2} \mu_G^2 - \frac{10\rho_D^3}{3m_b^3} - \frac{1}{3m_b^4} \left(4r_G^4 + 4r_E^4 + \frac{1}{4}s_{qB}^4 - 4s_E^4 \right) + \mathcal{O}(1/m_b^5) \right]$$

Example 2: $b \rightarrow c\ell\bar{\nu}$ ($\rho = m_c^2/m_b^2$)

$$\begin{aligned} \frac{\Gamma}{\Gamma_0} &= \mu_3 z(\rho) - 2 \frac{\mu_G^2}{m_b^2} (\rho - 1)^4 + d(\rho) \frac{\tilde{\rho}_D^3}{m_b^3} + \frac{2}{3} (-1 + \rho)^3 (1 + 5\rho) \frac{s_B^4}{m_b^4} \\ &\quad - \frac{8}{9} \frac{r_E^4}{m_b^4} \left(2 + 9\rho^2 - 20\rho^3 + 9\rho^4 + 6 \log \rho \right) \\ &\quad + \frac{4}{9} \frac{r_G^4}{m_b^4} \left(16 - 21\rho + 9\rho^2 - 7\rho^3 + 3\rho^4 + 12 \log \rho \right) \\ &\quad + \frac{1}{9} \frac{s_E^4}{m_b^4} \left(50 - 72\rho + 40\rho^3 - 18\rho^4 + 24 \log \rho \right) \\ &\quad + \frac{1}{36} \frac{s_{qB}^4}{m_b^4} \left(-25 + 48\rho - 36\rho^2 + 16\rho^3 - 3\rho^4 - 12 \log \rho \right) + \mathcal{O}(1/m_b^5) \end{aligned}$$

Differential Rates and Moments

Define generalized moments with weight function w :

$$\langle M[w] \rangle = \int \frac{d^4 q}{(2\pi)^4} \widetilde{dk} \widetilde{dk}' w(v, k, k') \langle R(s) \rangle L(k, k') (2\pi)^4 \delta^4(q - k - k')$$

which have an OPE in term of the “operator kernel”

$$M[w] = \sum_{n=0}^{\infty} a_{\mu_1 \dots \mu_n}^{(n)} \otimes \bar{b}_v (iD_{\mu_1} \dots iD_{\mu_n}) b_v$$

If the weight function is RPI (i.e. is independent of v)

$$\delta_{\text{RP}} w(v, k, k') = 0$$

the coefficients $a^{(n)}$ of the operator kernel satisfy the same “rigid” relations as the total rates. Thus they depend on the same reduced set of HQE parameters.

For the case of $b \rightarrow c$ semileptonics, the weight function

$$w(v, k, k') = \delta(q^2 - (k + k')^2)$$

satisfies this: **The leptonic invariant mass spectrum will depend on the reduced set of HQE parameters**

This thus also holds for moments:

$$\frac{1}{\Gamma_0} \int d\hat{q}^2 (\hat{q}^2)^n \frac{d\Gamma}{d\hat{q}^2}$$

and for cut moments, **as long as the cut is Lorentz invariant**, e.g.

$$\frac{1}{\Gamma_0} \int_{q_{\text{cut}}^2} d\hat{q}^2 (\hat{q}^2)^n \frac{d\Gamma}{d\hat{q}^2}$$

In the massless limit:

$$\begin{aligned}
 Q_1 = & \frac{3}{10} \mu_3 - \frac{7}{5} \frac{\mu_G^2}{m_b^2} + \frac{\tilde{\rho}_D^3}{m_b^3} (19 + 8 \log \rho) - \frac{r_E^4}{m_b^4} \left(\frac{1292}{45} + \frac{40}{3} \log \rho \right) - \frac{s_B^4}{m_b^4} (8 + 2 \log \rho) \\
 & + \frac{13}{120} \frac{s_{qB}^4}{m_b^4} + \frac{s_E^4}{m_b^4} \left(\frac{63}{5} + 4 \log \rho \right) + \frac{r_G^4}{m_b^4} \left(\frac{827}{45} + \frac{22}{3} \log \rho \right), \quad (4.10)
 \end{aligned}$$

$$\begin{aligned}
 Q_2 = & \frac{2}{15} \mu_3 - \frac{16}{15} \frac{\mu_G^2}{m_b^2} + \frac{\tilde{\rho}_D^3}{m_b^3} \left(\frac{358}{15} + 8 \log \rho \right) - \frac{r_E^4}{m_b^4} \left(\frac{2888}{45} + \frac{64}{3} \log \rho \right) - \frac{s_B^4}{m_b^4} \left(\frac{259}{15} + 4 \log \rho \right) \\
 & + \frac{s_{qB}^4}{m_b^4} \left(\frac{251}{180} + \frac{1}{3} \log \rho \right) + \frac{s_E^4}{m_b^4} \left(\frac{908}{45} + \frac{16}{3} \log \rho \right) + \frac{r_G^4}{m_b^4} \left(\frac{1373}{45} + \frac{28}{3} \log \rho \right), \quad (4.11)
 \end{aligned}$$

etc.

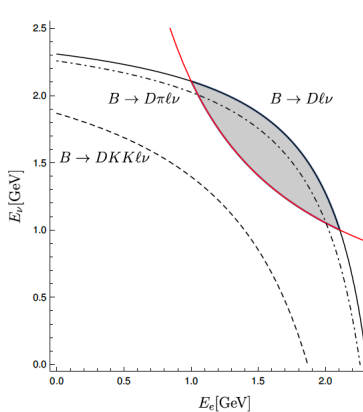
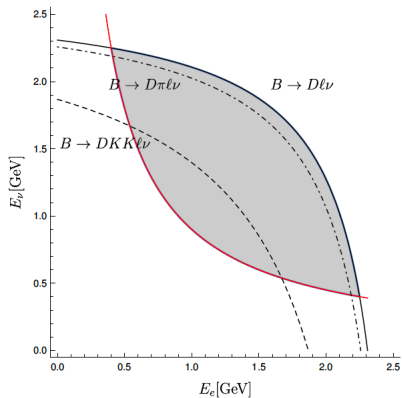
NB.: In the massless limit one need to include also four quark operators, and

$\log \rho = \log(\mu^2 / m_b^2)$, the μ dependence is compensated by the four quark operators

We have computed the full expression (for finite ρ) for all Q_n , as well as for the spectrum $d\Gamma / (dq^2)$.

Cut Moments

Implementing a q^2 cut:



The

expression for the cut moments are available!

Conclusion on Part I

- V_{cb} may be determined with a smaller number of independent HQE parameters
- Strategy is the same as before, but based on a different set of observables
- **Perspective for a data-driven analysis up to $1/m_b^4$**

Part II

Preliminary Results on $\alpha_s \Lambda_{\text{QCD}}^3 / m_b^3$

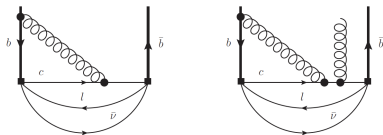
ThM, A. A. Pivovarov: SI-HEP-2018-36, arXiv:1903.xxxxx

Status of the $b \rightarrow c$ semileptonic HQE Calculation

$$\Gamma = \Gamma_0 \left(C_0(\rho) \left[1 + \frac{\mu_\pi^2}{2m_b^2} \right] + C_G(\rho) \frac{\mu_G^2}{2m_b^2} + C_{rD}(\rho) \frac{\rho_D^3}{6m_b^3} + \dots \right)$$

- Tree level terms up to and **including $1/m_b^5$** known
- $\mathcal{O}(\alpha_s)$ and full $\mathcal{O}(\alpha_s^2)$ for the partonic rate known
- $\mathcal{O}(\alpha_s)$ for the μ_π^2/m_b^2 and μ_G^2/m_b^2 is known
- QCD inspired modelling for the HQE matrix elements
- In the pipeline:
 - $\mathcal{O}(\alpha_s)$ for the ρ_D/m_b^3 and ρ_{LS}/m_b^3
 - Relations among the coefficients from RPI

Some technical remarks ...



$$\tilde{T} = C_0 \mathcal{O}_0 + C_V \frac{\mathcal{O}_V}{m_b} + C_\pi \frac{\mathcal{O}_\pi}{2m_b^2} + C_G \frac{\mathcal{O}_G}{2m_b^2} + C_D \frac{\mathcal{O}_D}{2m_b^3}$$

$$\mathcal{O}_0 = \bar{h}_V h_V \quad \mathcal{O}_V = \bar{h}_V \not{v} \pi h_V \quad \mathcal{O}_\pi = \bar{h}_V \pi_\perp^2 h_V$$

$$\mathcal{O}_G = \bar{h}_V \frac{1}{2} [\gamma_\mu, \gamma_\nu] \pi_\perp^\mu \pi_\perp^\nu h_V \quad \mathcal{O}_D = \bar{h}_V [\pi_\perp^\mu, [\pi_\perp^\mu, \pi \not{v}]] h_V$$

- Compute the three-loop Feynman Diagrams
- Perform the renormalization

... a few subtleties ...

(a) Rewrite

$$\bar{b}\psi b = \mathcal{O}_0 + \tilde{C}_\pi \frac{\mathcal{O}_\pi}{2m_b^2} + \tilde{C}_G \frac{\mathcal{O}_G}{2m_b^2} + \tilde{C}_D \frac{\mathcal{O}_D}{2m_b^3} + O(\Lambda_{\text{QCD}}^4 / m_b^4)$$

(b) Renormalization:

- Heavy Quark Fields (on-shell renormalization)
- Charm mass in the $\overline{\text{MS}}$ scheme
- Renormalization of ρ_D in $\overline{\text{MS}}$
- Mixing with time-ordered products like

$$(-i) \int d^4x \text{T}[\mathcal{L}_{\text{HQET}}(x) \mathcal{O}_\pi]$$

All poles cancel ... (I still have to check this!)

Result (preliminary!)

We have the analytic result for C_{rD}

Numerically we have ($m_c^2/m_b^2 = 0.07$, $\alpha_s(m_b) = 0.2$)

$$\begin{aligned} C_{rD} &= -57.1588 + \frac{\alpha_s}{4\pi} (-56.5941 C_A + 446.793 C_F) \\ &= -57.1588 + \frac{\alpha_s}{4\pi} (425.942) \\ &= -57.1588 \left(1 - \frac{\alpha_s}{4\pi} 7.4519 \dots\right) \\ &= -57.1588 (1 - 0.12) \end{aligned}$$

- $\rho_D = \rho_D(m_b)$
- Leading order has a sizeable coefficient
- QCD corrections have the expected size.
- Impact on V_{cb} will be small but visible

Conclusion on Part II

- $\alpha_s \Lambda_{\text{QCD}}^3 / m_b^3$ is almost completed for the total rate
- We can also compute moments of kinematic distributions
- **We will not have the full phase space distributions**, so we cannot implement an electron energy cut
- Can we use RPI to obtain the coefficient of ρ_{LS} ?

Overall Conclusions

Inclusive V_{cb} is getting more and more precise:

- The problem exclusive vs inclusive seems to disappear
- Subtle problems will arise once we want to increase the precision further:
 - Convergence of the HQE
 - Duality Violations
 - α_s^3 for the partonic rate
 - ...